

# Around the 3D diffusion: stability and testing

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#### 3D diffusion in spherical and terrain-following coordinates

scalar flux divergence

$$\rho \, \frac{\partial s}{\partial t} = -\nabla_j H^j = -\frac{\partial}{\partial x^j} H^j - \Gamma^j_{jk} H^k$$

vectorial flux divergence

$$\frac{\partial v^i}{\partial t} = -\frac{1}{\rho} \nabla_j T^{ij} = -\frac{1}{\rho} \left( \frac{\partial}{\partial x^j} T^{ij} + \Gamma^i_{jk} T^{kj} + \Gamma^j_{jk} T^{ik} \right)$$

with gradient expressions for scalar diffusion flux H and momentum flux tensor T

$$H^{i} = -\rho K_{s} g^{ij} \nabla_{j} s$$
$$T^{ij} = -\rho K \left( g^{il} \nabla_{l} v^{j} + g^{jl} \nabla_{l} v^{i} \right)$$

These formulae cannot be used directly since many meteorological models use *spherical* (i.e. *non-terrain-following*) and *normalized* base vectors  $\rightarrow$ 





#### **Metric terms of 3D-diffusion**

#### scalar flux divergence:



earth curvature

scalar fluxes:

$$H^{*1} = -\rho K_s \frac{1}{r \cos \phi} \left( \frac{\partial s}{\partial \lambda} + \frac{J_\lambda}{\sqrt{G}} \frac{\partial s}{\partial \zeta} \right),$$
  

$$H^{*2} = -\rho K_s \frac{1}{r} \left( \frac{\partial s}{\partial \phi} + \frac{J_\phi}{\sqrt{G}} \frac{\partial s}{\partial \zeta} \right),$$
  

$$H^{*3} = +\rho K_s \frac{1}{\sqrt{G}} \frac{\partial s}{\partial \zeta},$$

analogous: ,vectorial' diffusion of *u*, *v*, *w* 

Baldauf (2005), COSMO-Newsl. Nr. 5





### **Current implementation, Numerics**

- all metric terms are treated explicitly -> implemented in Subr. explicit\_horizontal\_diffusion
- This is done in a two-step approach:
  1. discretize the fluxes at their ,own' positions
  2. discretize the divergence of these fluxes
- PHYCTL-namelist-parameter 13dturb, 13dturb\_metr

Positions of turbulent fluxes in staggered grid:







#### Disadvantages of the current implementation in COSMO

- due to the above mentioned ,two-step' approach, some terms use a wider stencil than necessary solution: analytically insert fluxes i.e. boil down the formulae to derivatives of the prognostic variables and discretize afterwards (results in a huge amount of terms!)
- the scalar diffusion has been tested (*Baldauf (2005) COSMO-Newsl. No 6*) but not the vector diffusion!
- though spherical metric terms are contained, only shallow atmosph. approx is used (i.e. 1/r · … → 1/a · …). However, this is less important.
- the current implementation is not stable in steep terrain (*Langhans ...*)

> new implementation of subroutines explicit\_horiz\_diffusion
and implicit\_vert\_diff\_uvwt





## Stability





# Linear von-Neumann stability analysis of 2D (i.e. x-z-diffusion) in tilted terrain



 $\Delta x$ 

$$\frac{\Delta h}{\Delta z} = m \cdot M_z$$

dimensionless variables: diffusion-Courant number  $C_{diff} := K \frac{\Delta t}{\Delta r^2}$ steepness  $m := \left. \frac{\partial z}{\partial x} \right|_{\mathcal{L}} = \frac{\Delta h}{\Delta x}$ grid anisotropy  $M_z :=$ 





vertical

# Linear von-Neumann stability analysis of 2D (i.e. x-z-diffusion) in tilted terrain

#### scalar flux divergence:

$$\rho \frac{\partial s}{\partial t} = -\frac{1}{r \cos \phi} \frac{\partial H^{*1}}{\partial \lambda} - \frac{J_{\lambda}}{\sqrt{G}} \frac{1}{r \cos \phi} \frac{\partial H^{*1}}{\partial \zeta} - \frac{1}{r} \frac{\partial H^{*2}}{\partial \phi} - \frac{J_{\phi}}{\sqrt{G}} \frac{1}{r} \frac{\partial H^{*2}}{\partial \zeta} + \frac{1}{\sqrt{G}} \frac{\partial H^{*3}}{\partial \zeta} \\ -\frac{2}{r} H^{*3} + \frac{\tan \phi}{r} H^{*2}.$$

scalar fluxes:

$$H^{*1} = -\rho K_s \frac{1}{r \cos \phi} \left( \frac{\partial s}{\partial \lambda} + \frac{J_\lambda}{\sqrt{G}} \frac{\partial s}{\partial \zeta} \right),$$
  

$$H^{*2} = -\rho K_s \frac{1}{r} \left( \frac{\partial s}{\partial \phi} + \frac{J_\phi}{\sqrt{G}} \frac{\partial s}{\partial \zeta} \right),$$
  

$$H^{*3} = +\rho K_s \frac{1}{\sqrt{G}} \frac{\partial s}{\partial \zeta},$$

analogous: ,vectorial' diffusion of *u*, *v*, *w* 

Baldauf (2005), COSMO-Newsl. Nr. 5





max C 1000 800 1000 800 600 600 .52 400 400 .48 .44 200 200 .4 100 .36 .32 Mz .28 .24 .2 10 .16 .12 .08 .04 .04 12 1 \_ 0.0 0.0 1.0 2.0 3.0 4.0 5.0 m max C: min=0 max=0.250001 (60, 50, 50, 50, 0, 0.5, 1e-05) (60, 50, 50, 50, 0, 0, 1e-05)

#### purely explicit

vertically implicit, only ,pure' z-deriv no off-centering (0.5)



MZ







# Linear von-Neumann stability analysis of 2D (i.e. x-z-diffusion) in tilted terrain

scalar flux divergence:  

$$\rho \frac{\partial s}{\partial t} = -\frac{1}{r \cos \phi} \frac{\partial H^{*1}}{\partial \lambda} \left[ -\frac{J_{\lambda}}{\sqrt{G}} \frac{1}{r \cos \phi} \frac{\partial H^{*1}}{\partial \zeta} - \frac{1}{r} \frac{\partial H^{*2}}{\partial \phi} \right] \left[ -\frac{J_{\phi}}{\sqrt{G}} \frac{1}{r} \frac{\partial H^{*2}}{\partial \zeta} + \frac{1}{\sqrt{G}} \frac{\partial H^{*3}}{\partial \zeta} \right] - \frac{2}{r} H^{*3} + \frac{\tan \phi}{r} H^{*2}.$$
vertical

scalar fluxes:

$$H^{*1} = -\rho K_s \frac{1}{r \cos \phi} \left( \frac{\partial s}{\partial \lambda} + \frac{J_\lambda}{\sqrt{G}} \frac{\partial s}{\partial \zeta} \right),$$
  

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$$H^{*3} = +\rho K_s \frac{1}{\sqrt{G}} \frac{\partial s}{\partial \zeta},$$

analogous: ,vectorial' diffusion of *u*, *v*, *w* 

Baldauf (2005), COSMO-Newsl. Nr. 5





now: all vertical terms implicit !







now: all vertical terms implicit !







**Vector diffusion** 





max C 1888 600 .26 400 .24 .22 200 .2 100 .18 .16 Mz .14 .12 .1 10 .08 .06 .04 .02 .02 1 \_ 0.0 2.0 8.0 4.0 6.0 m max C: min=4.76837e-06 max=0.125003 (0, 0, 0, 0, 0, 0)

purely explicit

#### vertically implicit, only ,pure' z-deriv off-centering=0.7 (=current COSMO)



<sup>(60, 50, 50, 50, 0, 0.7, 0, 0, 0, 0, 0, 7, 1</sup>e-05)





now: vertically implicit treatment of all possible terms

off-centering (0.6)



no off-centering (0.5)





vertically implicit treatment of all possible terms







vertically implicit treatment of all possible terms, off-centering 0.6







### Testing



DWC

### Test of scalar diffusion: 3-dim. isotropic gaussian tracer distribution

3D diffusion equation:

$$\frac{\partial \phi}{\partial t} = K \Delta \phi$$

analytic Gaussian solution for K=const.:

$$\phi(r,t) = \frac{\Phi_0}{\sqrt{4\pi K(t+t_0)^3}} \exp\left(-\frac{r^2}{4K(t+t_0)}\right),$$
  
$$r := \sqrt{x^2 + y^2 + z^2}$$

Baldauf (2005) COSMO-Newsl. no. 6











#### nearly isotropic grid

goal: show false diffusion in the presence of orography



2400 2200 2000 1800 1600 E 1400 N 1200 1000 800 600 400 0.003 0.006 0.009 0.012 0.015 0.018 0.021 0.024 Lon. [Deg.]

t=3 min., z=825m, 3D, ohne Metrik, mit Orographie











### Case 4: 3D-diffusion, with metric terms, with orography

nearly isotropic grid

goal: show correct implementation of the new metric terms



t=3 min., z=825m, 3D, mit Metrik, mit Orographie

















T,analy: min=273 max=276.52 T,simul: min=273 max=275.866

5.1r14b\_2.8\_s\_h1000m\_3dneu\_3dturbT\_3dmetrTi0.75

analytic solution: solid lines COSMO solution: colors, dashed lines

















































#### A proposal for an analogous test case for 3D vector diffusion:

isotropic, purely radial vector field:

 $\mathbf{v}(\mathbf{r},t) = v_r(r,t)\hat{\mathbf{e}}_r$ 

with

$$v_r(r,t) = const. \frac{r}{\sqrt{K}(t+t_0)^{5/2}} \cdot e^{-\frac{r^2}{8K(t+t_0)}}$$







v,analy: min=1.16256e-10 max=12.3223 v,simul: min=4.30403e-10 max=12.3223

5.1r14b\_2.8\_v\_h1000m\_3dneu\_3dturbT\_3dmetrTi0.75





|v|, t=0500 timesteps 12 18000 10.8 9.6 8.4 16000 7.2 <sup>N</sup> 14000 6 4.8 3.6 12000 2.4 1.2 10000 0 -8000 -4000 8000 4000 0 Х

v,analy: min=3.9194e-15 max=7.8975 v,simul: min=2.71722e-11 max=6.67046

5.1r14b\_2.8\_v\_h1000m\_3dneu\_3dturbT\_3dmetrTi0.75























#### Summary

- The current version of 3D-diffusion is not stable in steep terrain (finding by MeteoCH) ← instability of metric terms
- A new stability analysis indicates how to cure the problem:
  - treat **all** of the possible vertical terms in the tridiagonal solver
  - some off-centering (~0.7) is necessary and leads to

$$K\Delta t\left(rac{1}{\Delta x^2}+rac{1}{\Delta y^2}
ight)\leq$$
 0.45 (scalar case) / 0.146 (vector case)

- in the v-equations additional forward-backward even slightly increases stability (not planned yet)
- new subroutines explicit\_horiz\_diffusion (replaces the old one) and implicit\_vert\_diff\_uvwT\_3D
- A new test case for vector diffusion has been proposed





#### Outlook

- Resharp the test case definition a bit (less time steps, perhaps numerical diffusivity of the diffusion is too prominent?)
- or still bugs in implementation? ٠
- proper lower and upper boundary treatment ٠
- implement also spherical metric terms into explicit horiz diffusion ٠
- deliver a version of explicit horiz diffusion without shallow approx. ٠ i.e. using the *deep* form  $(1/a \rightarrow 1/r)$
- possibly available in COSMO 5.2 ٠ (volunteers for 4-eyes principle?)

