

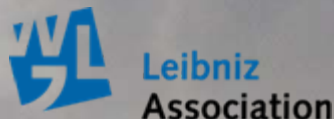
# Simulating aerosol particles' influence on clouds and precipitation with the model system COSMO-SPECS

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**COSMO User Workshop**  
Langen, 11.03.09



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# Motivation

- Aerosol particles have a great influence on weather and climate. They influence
  - cloud microstructure,
  - precipitation formation,
  - cloud albedo.

- Aerosol particles have a great influence on weather and climate. They influence
  - cloud microstructure,
  - precipitation formation,
  - cloud albedo.
- Detailed cloud microphysics schemes in mesoscale models can serve to gain a deeper understanding of the complex aerosol-cloud-precipitation interaction under different meteorological conditions.
- But: Such schemes are pretty slow! Thus efficient numerical schemes have to be found to make them suitable for realistic case studies and larger sensitivity studies. → [Part 2, Matthias Lieber](#)

# COSMO-SPECS

**Spectral microphysics & coupling scheme**

## SPECTral bin cloud microphysicS

### ➤ Spectra:

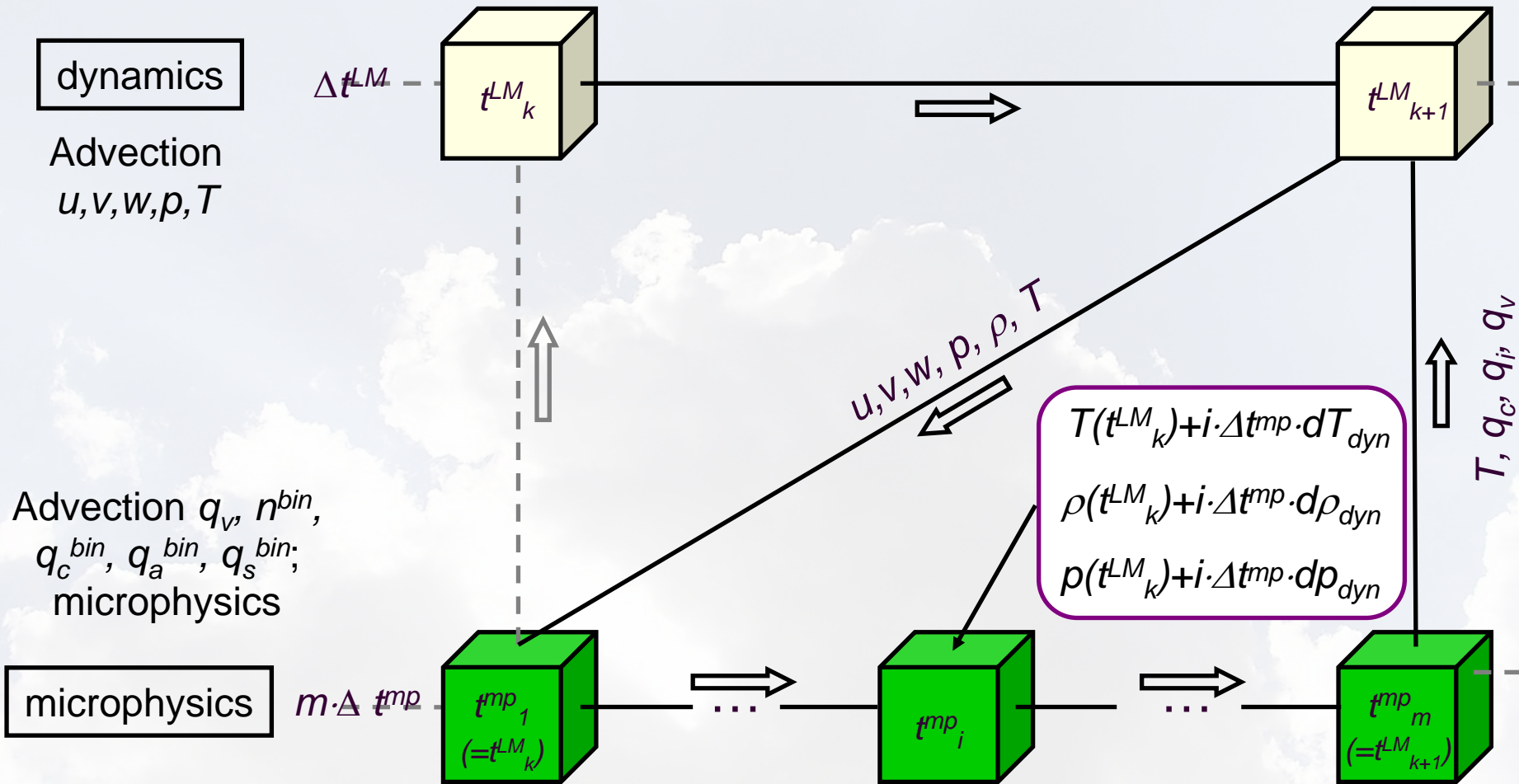
- Partly insoluble aerosol particles, cloud & rain drops (joint spectrum, number density, mixing ratios (MRs) of water, soluble and insoluble aerosol)
- Mixed phase hydrometeors (as above+frozen water MR)
- Insoluble aerosol particles (number and mass MR)
- 66 mass-doubling size bins each, starting at 10-15  $\mu\text{g}$  ( $r=1$  nm)

### ➤ Microphysical processes:

- Droplet nucleation, immersion/contact freezing
- Condensation/evaporation, deposition/sublimation
- Collisions between the various hydrometeors
- Freezing/melting

# Coupling scheme

Additionally: mass conserving, flux based advection for hydrometeors  
(Knoth and Wolke, 1998)



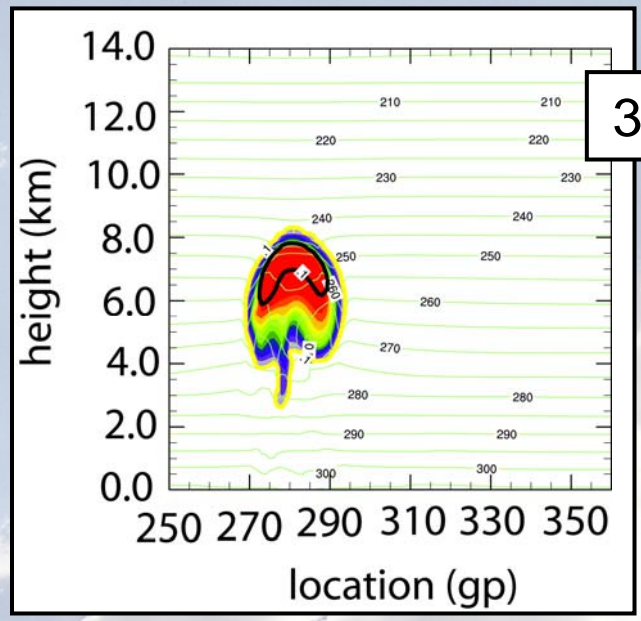
# Results

# Case setup

- Deep convective cloud basing on observations made in Midland, Texas, 13th August '99 (Rosenfeld and Woodley, 2000)
- 2D Model domain: horizontal 125 km, vertical 22 km  
Resolution: horizontal 250 m, vertical 125 m  
Time step: dynamical 2 s, microphysical 1 s
- Initial aerosol distribution unimodal  
number densities: 100/566/1000/5000 cm<sup>-3</sup>
- Simulated time: 3 hours, 100 CPUs, Altix (TU Dresden)  
(Time to simulate: 33-39 hours run time... ☹)

# Results

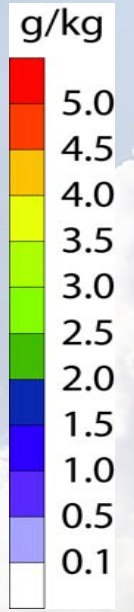
➤ Deep convective cloud observed over Midland, Texas, 1999  
(Rosenfeld and Woodley, 2000)



Initial aerosol  
conc.=  $566 \text{ cm}^{-3}$

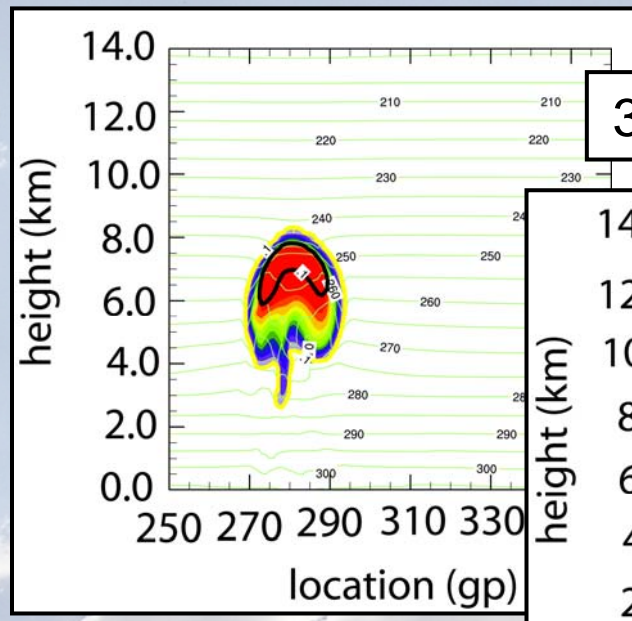
2D simulation  
 $\Delta x=250 \text{ m}$   
 $\Delta z=125 \text{ m}$   
 $\Delta t=2 \text{ bzw. } 1 \text{ s}$

— 0.1 g/kg ice  
— 0.1 g/kg cloud water  
— 0.1 g/kg rain water  
**colors:** condensed water MR



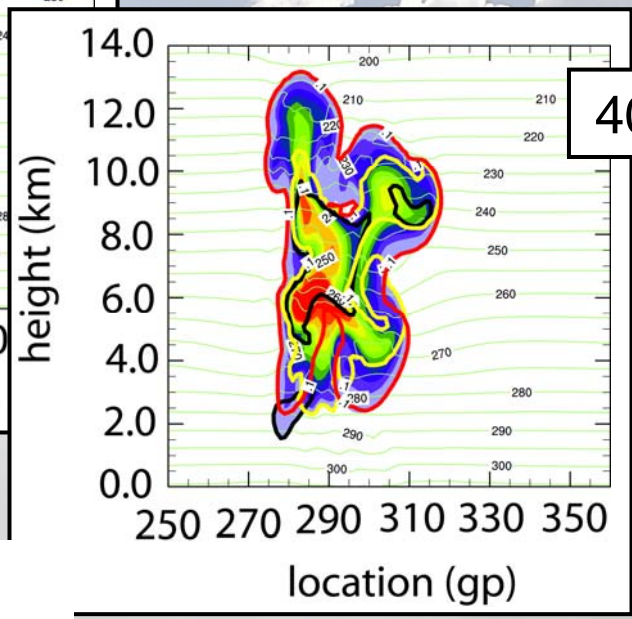
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➤ Deep convective cloud observed over Midland, Texas, 1999  
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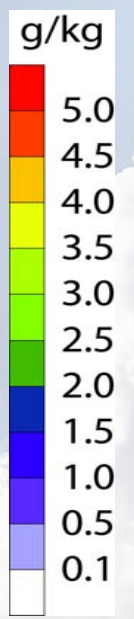


Initial aerosol  
conc. =  $566 \text{ cm}^{-3}$

2D simulation  
 $\Delta x = 250 \text{ m}$   
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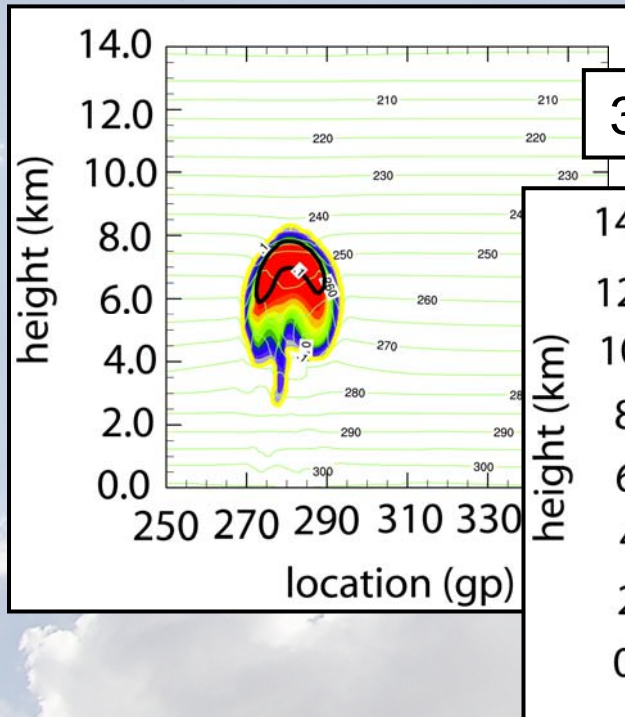


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— 0.1 g/kg rain water  
**colors:** condensed water MR



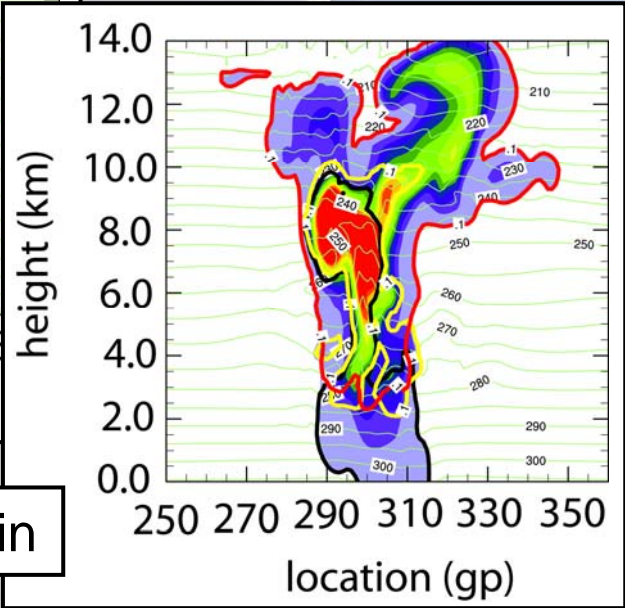
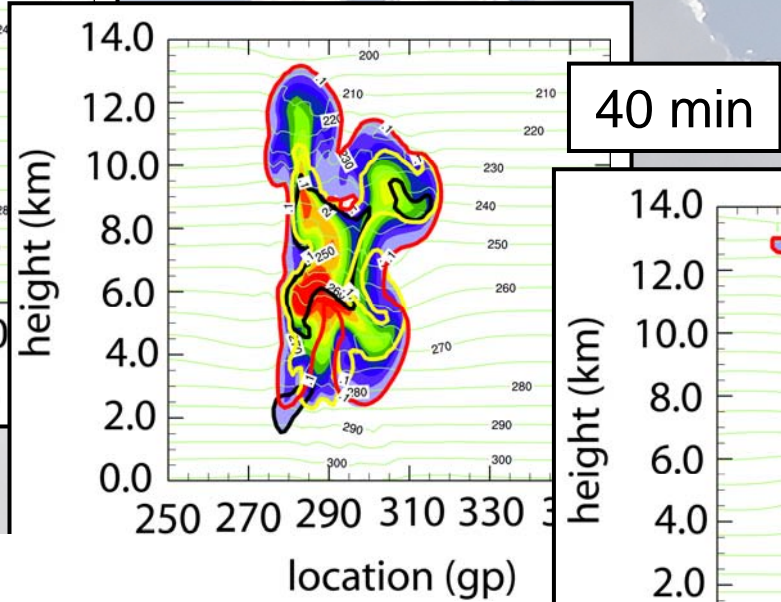
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- Deep convective cloud observed over Midland, Texas, 1999 (Rosenfeld and Woodley, 2000)



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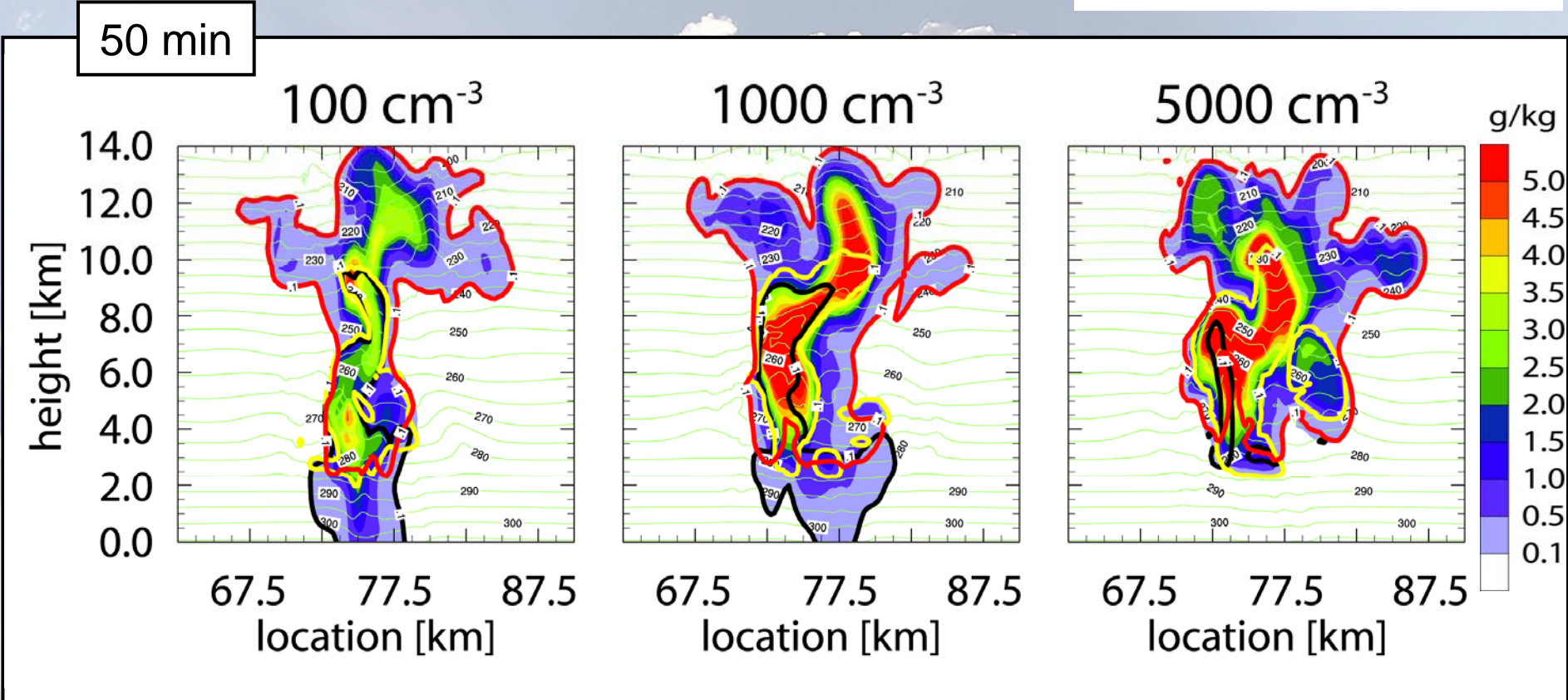
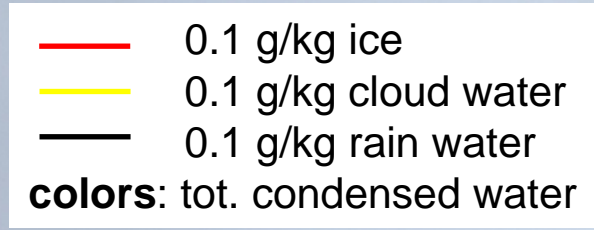
— 0.1 g/kg ice  
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# Results

➤ Deep convective cloud observed over Midland, Texas, 1999  
(Rosenfeld and Woodley, 2000)

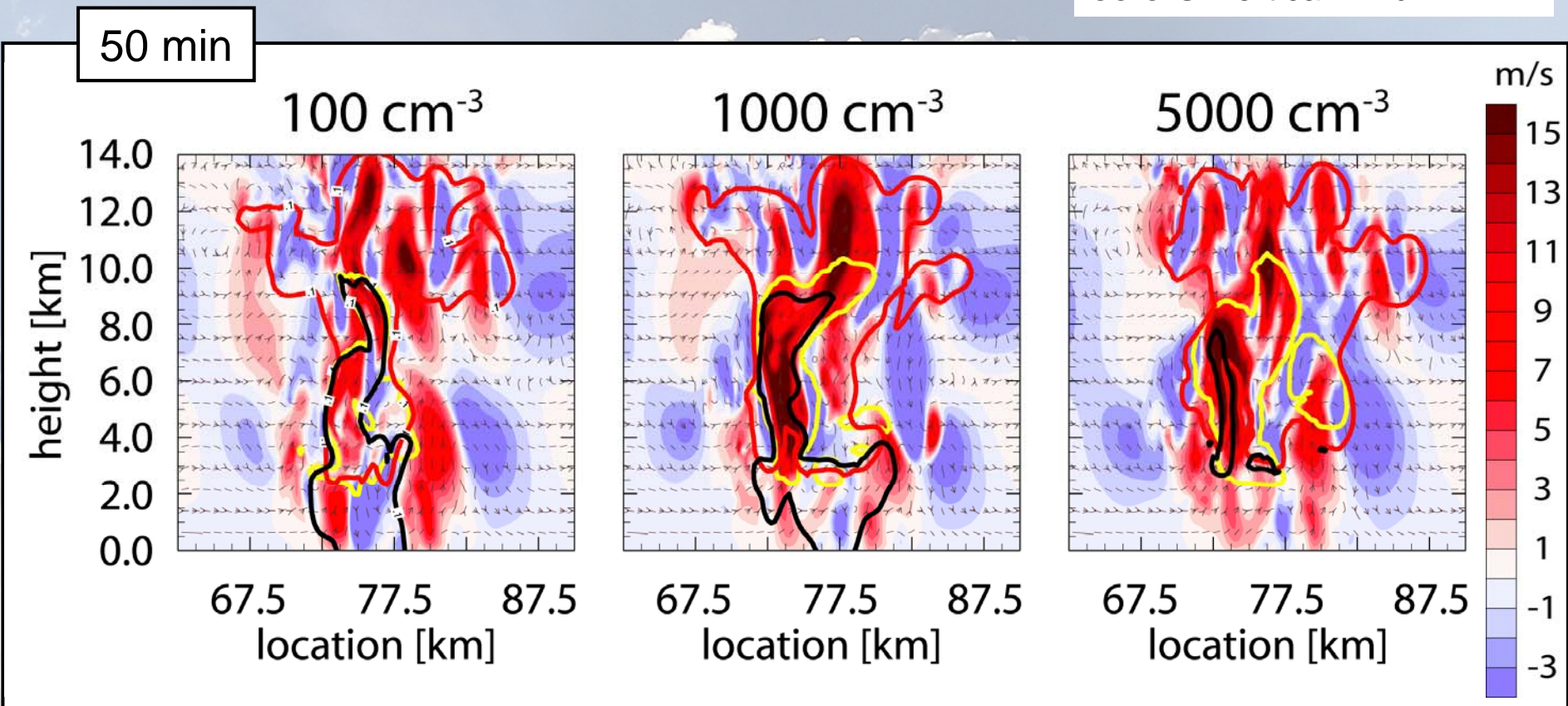
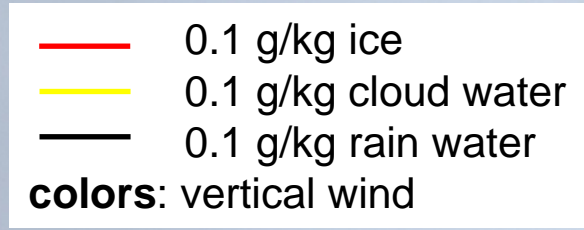
➤ **Condensed water**



# Results

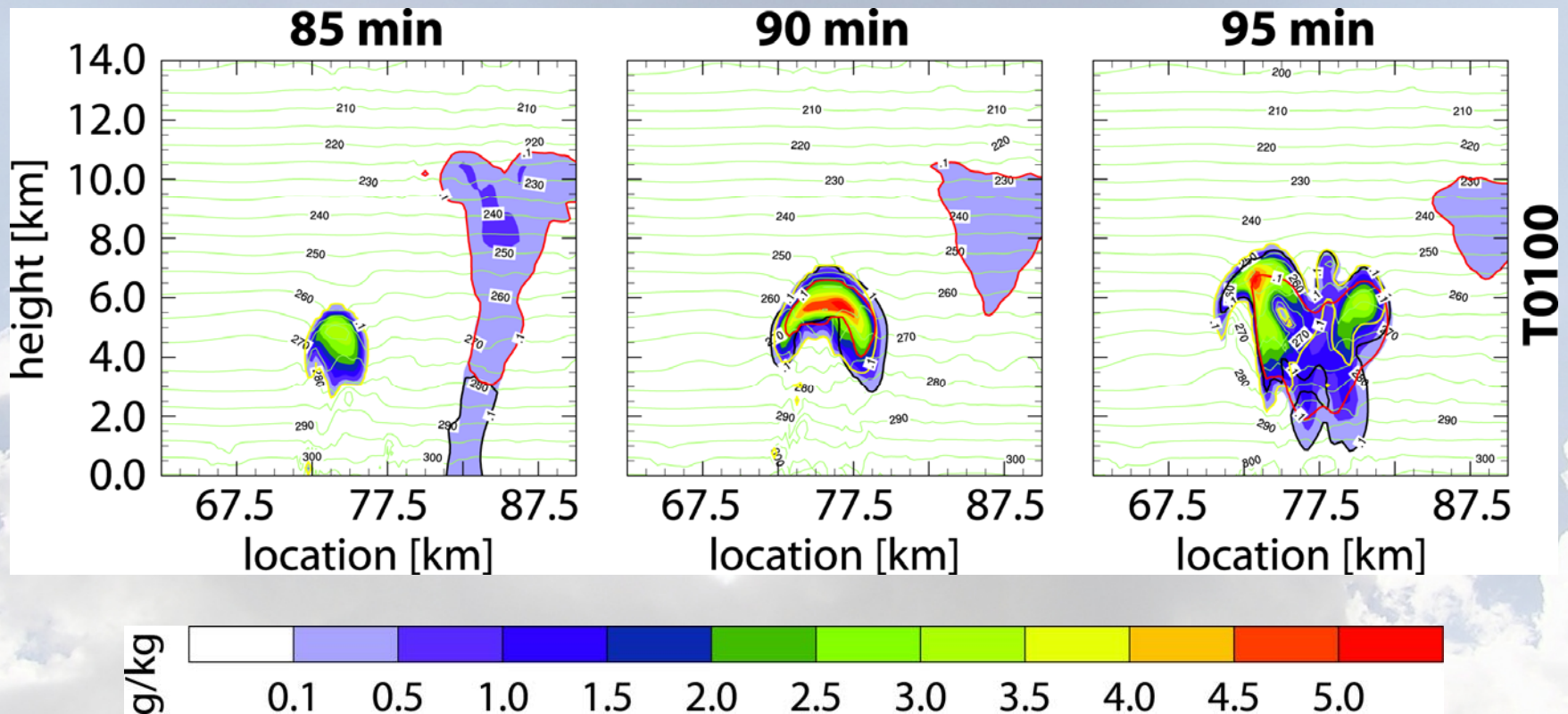
➤ Deep convective cloud observed over Midland, Texas, 1999 (Rosenfeld and Woodley, 2000)

➤ **Vertical wind**



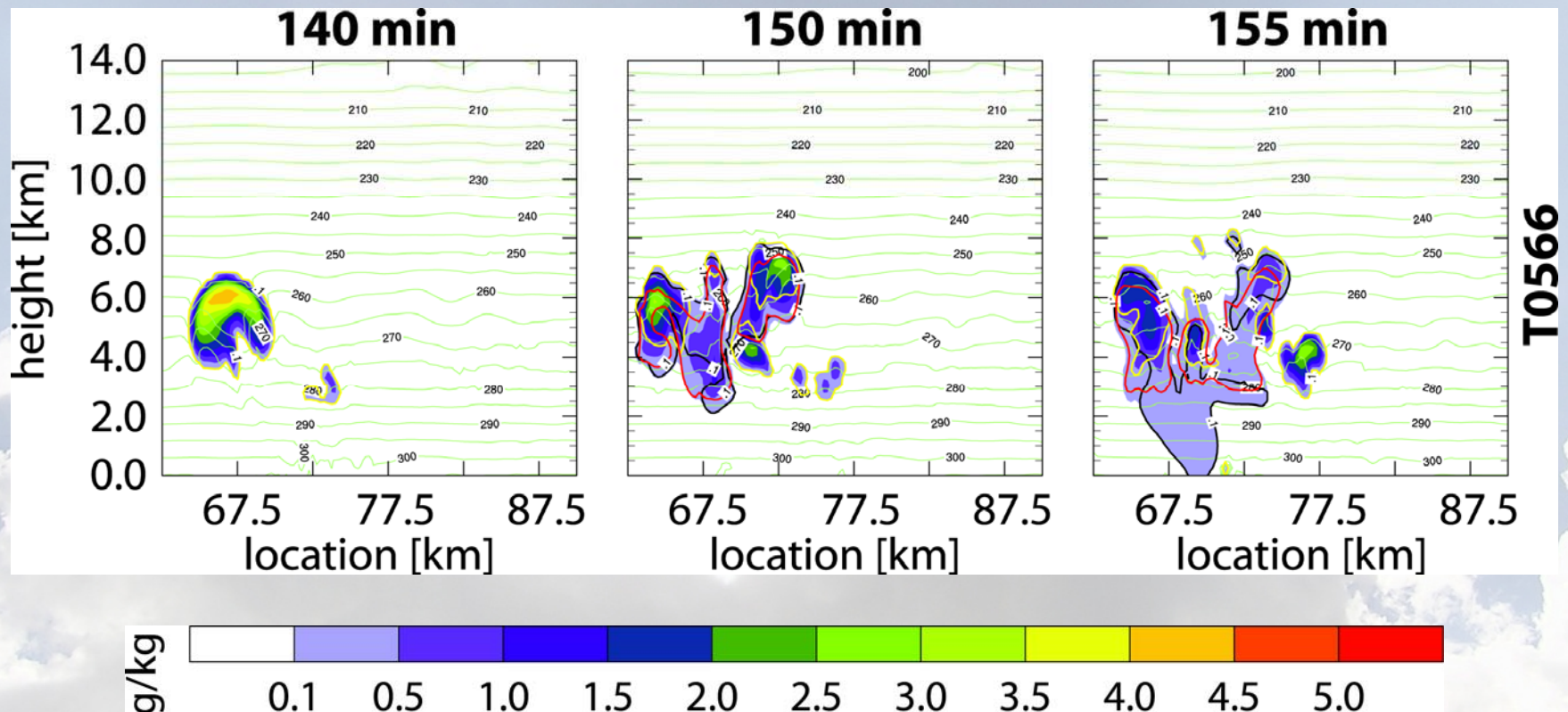
# Results

- Deep convective cloud observed over Midland, Texas, 1999 (Rosenfeld and Woodley, 2000)
- **Secondary cells –  $100 \text{ cm}^{-3}$**



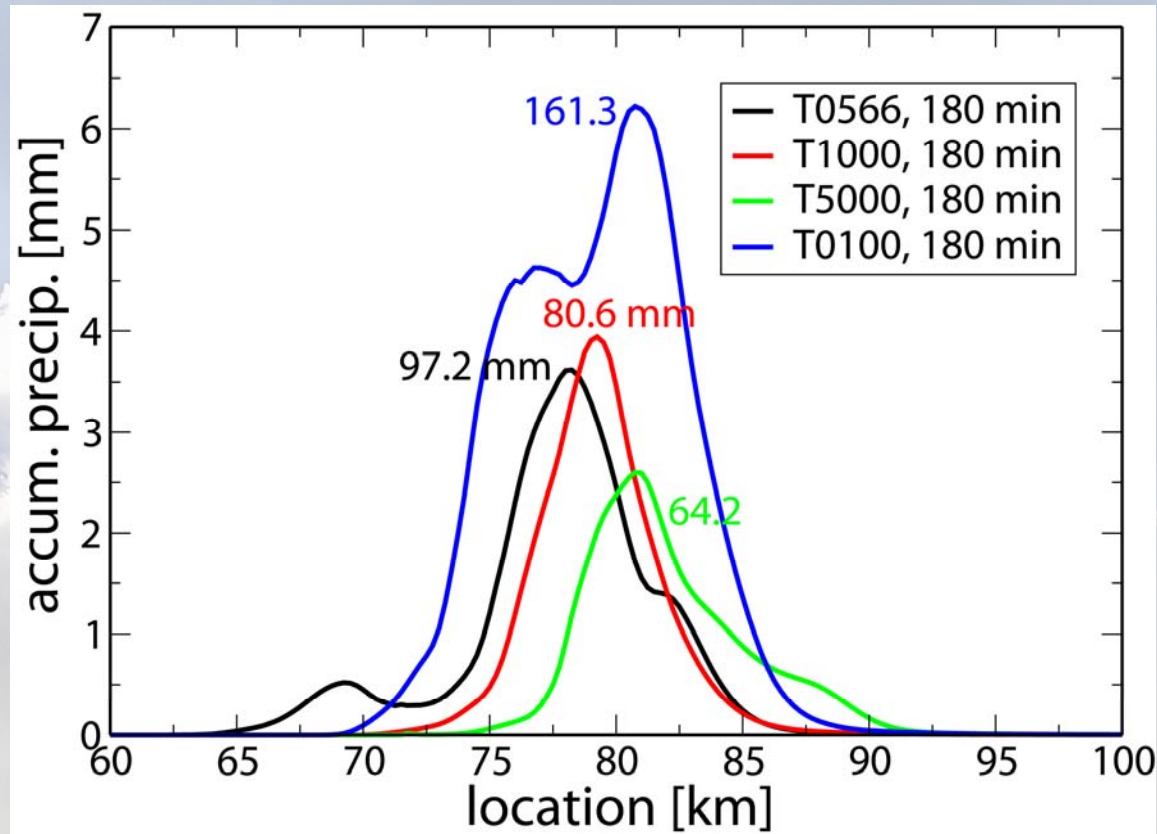
# Results

- Deep convective cloud observed over Midland, Texas, 1999 (Rosenfeld and Woodley, 2000)
- **Secondary cells – 566 cm<sup>-3</sup>**



# Results

- Deep convective cloud observed over Midland, Texas, 1999 (Rosenfeld and Woodley, 2000)
- **Accumulated precipitation**



# Conclusions and Outlook (Part 1)

# Conclusions (Part 1)

- For the “Texas case” presented it was shown that
  - Aerosol particles have a great influence on the cloud microstructure, i.e., a higher aerosol particle number yields more numerous and smaller droplets and ice particles.
  - Precipitation onset is delayed in more polluted environments and precipitation amount decreased.
  - Through these changes, the dynamics is altered as well.
  - Cleaner cases yield stronger secondary cells.

# Outlook (Part 1)

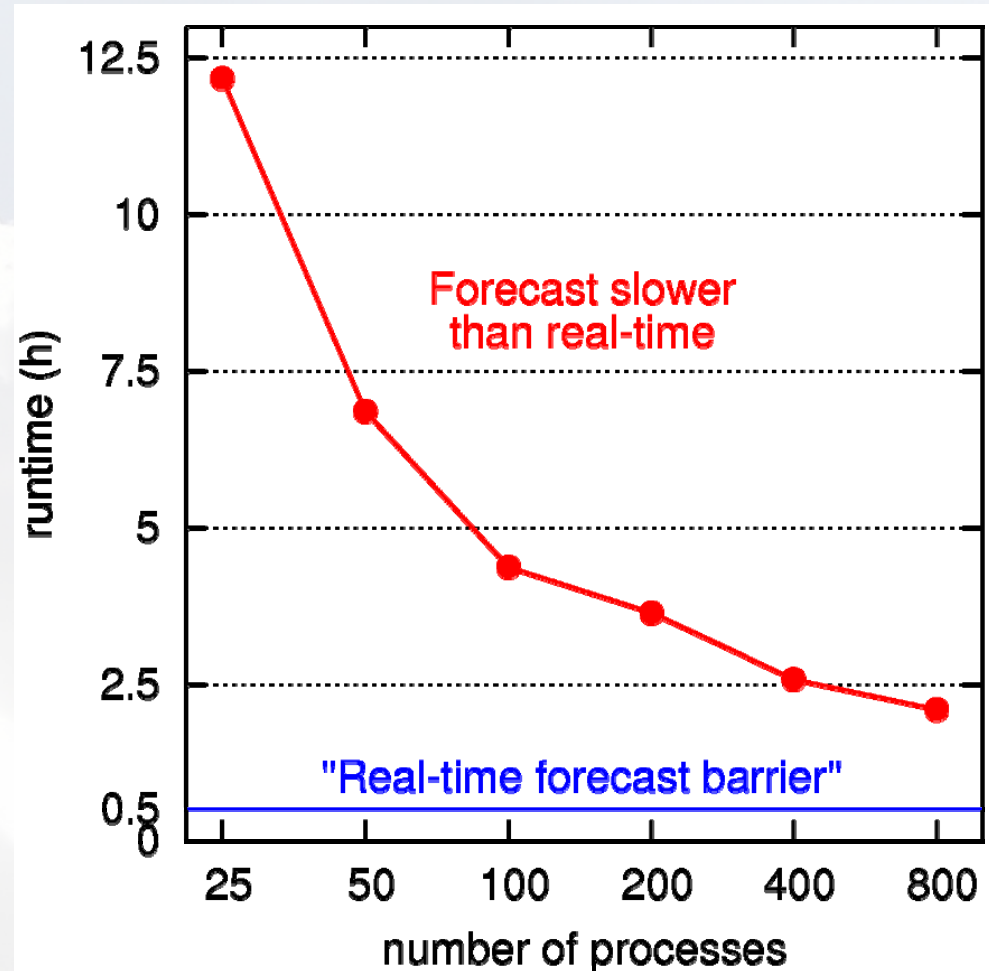
- COSMO-SPECS will be used to investigate aerosols' influence on clouds under different meteorological conditions. E.g.,
  - deep convective clouds in tropics/midlatitudes (sensitivity on humidity and temperature),
  - shallow (warm) cumuli under the influence of differing aerosol particle populations.
  - Playground is very large....
- Further freezing mechanisms will be implemented into SPECS to account for a broader range of cloud types.
- The module SPECS is currently being implemented into the LES model ASAM (All Scale Atmospheric Model, <http://asam.tropos.de>)

# Computational time

- Simple test case, heat bubble with  $80 \times 80 \times 48 = 307200$  grid points.
- Far from “real time forecast barrier”.
- Doesn't even scale anymore...!
- Let's see what an expert in computational science can do about this!

⇒ **Part 2, Matthias Lieber**

COSMO-SPECS Runtime for 30min Forecast  
80x80 bubble scenario, SGI Altix 4700



# Performance Problems in COSMO-SPECS

- Microphysical computations
  - Very expensive
  - Insufficient load balance
- Large amount of microphysical data per grid cell
  - Communication overhead
  - Writing GRIB output files

**Scalability?**

**We need special methods  
and tuning for HPC systems**

**“Parallel coupling framework and advanced time integration methods for detailed cloud processes in atmospheric models”**

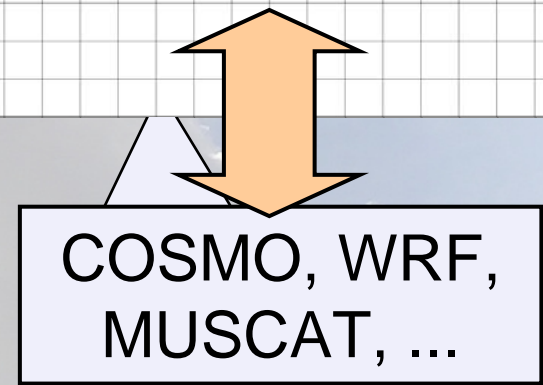
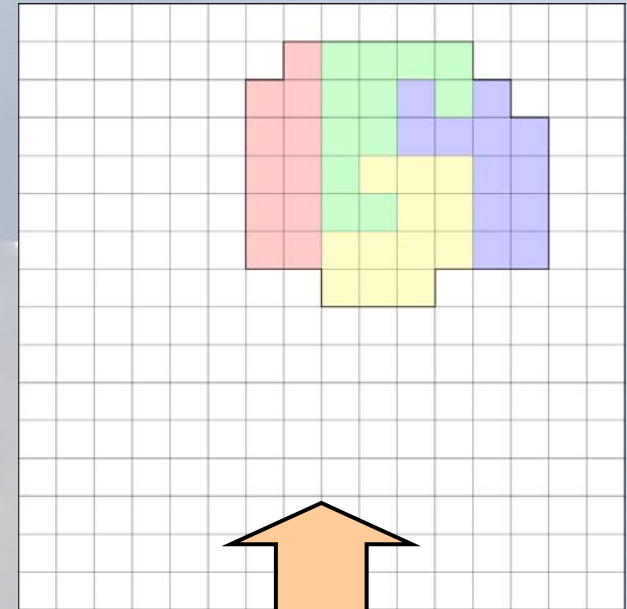
- Leibniz Institute for Tropospheric Research (IfT), Leipzig
- Center for Information Services and High Performance Computing (ZIH), TU Dresden
- Project started in September 2007
- Funded by

Deutsche  
Forschungsgemeinschaft

**DFG**

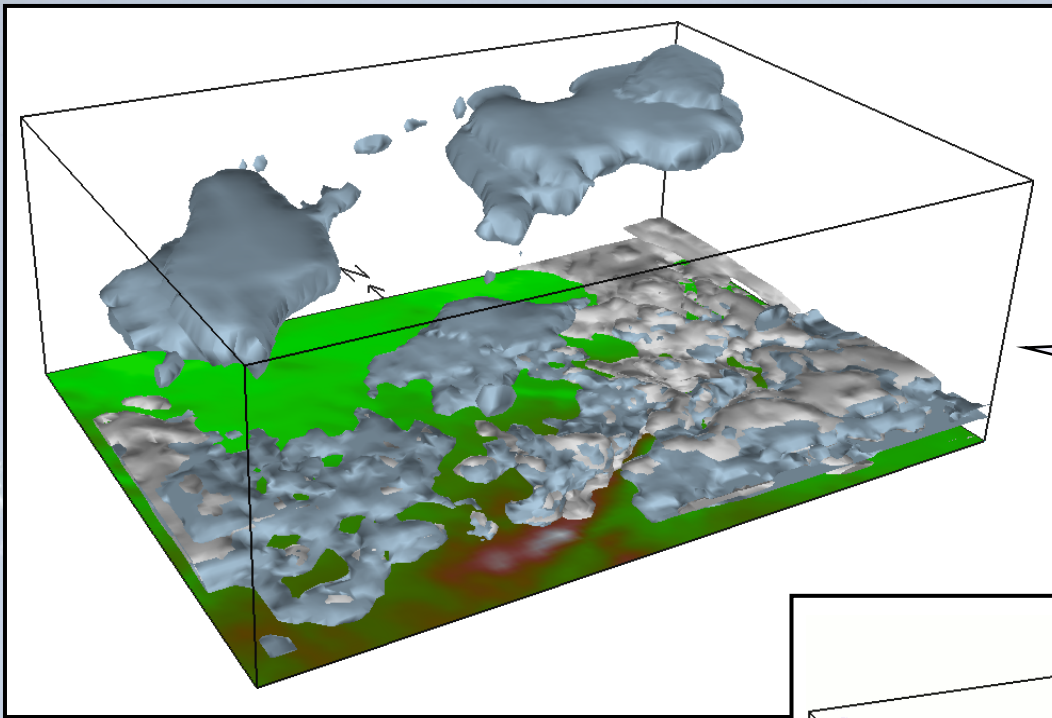
# Framework FD4 = “Four-Dimensional Distributed Dynamic Data Structures”

- Work in Progress
- Fortran 95, MPI parallel
- Block-based 3D decomposition of rectangular grid
- Adaption of grid to spatial cloud structure
- Parallelization
- Dynamic load balancing (Space-filling Curve, ParMetis)
- Coupling to atmospheric Model
- General applicable for multiphase modeling



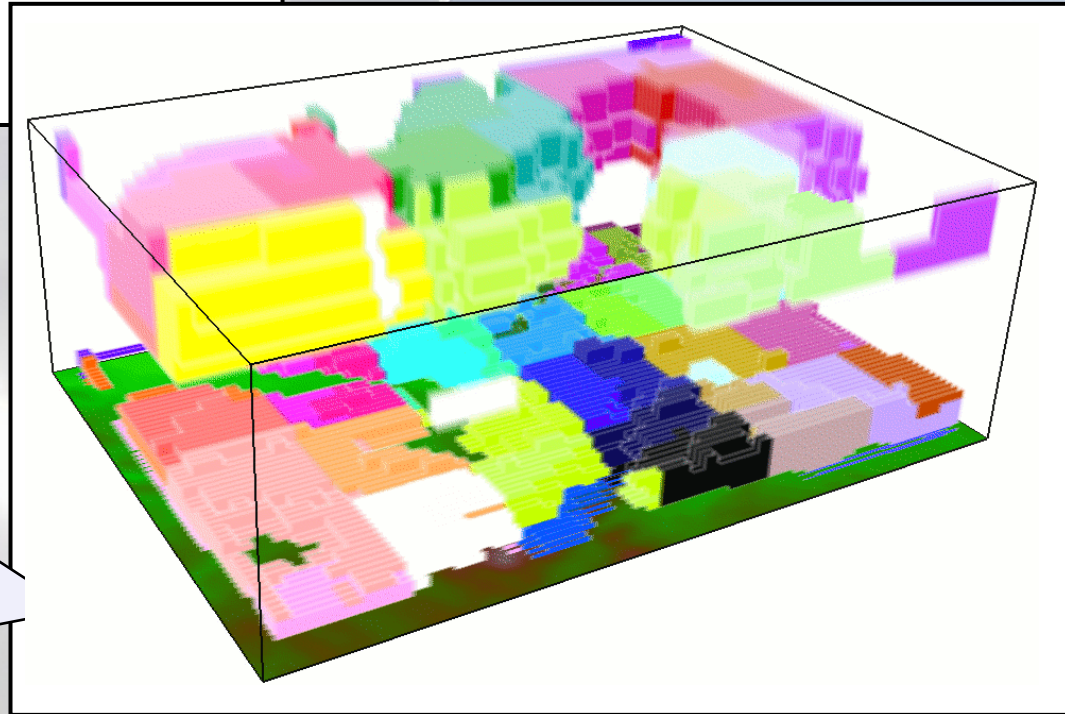
# Test Application: Adaption to Cloud Structure of a real Case

- COSMO real case
  - Saxony, 102x72x50 grid
  - 2,8 km horizontal Resolution
  - 20 December 2007
- Adaption to cloud structure
  - Transfer qc and qi to FD4 each step
  - FD4 creates blocks where  $qc > 0$  or  $qi > 0$
  - Blocks and partitioning adapt dynamically to cloud structure

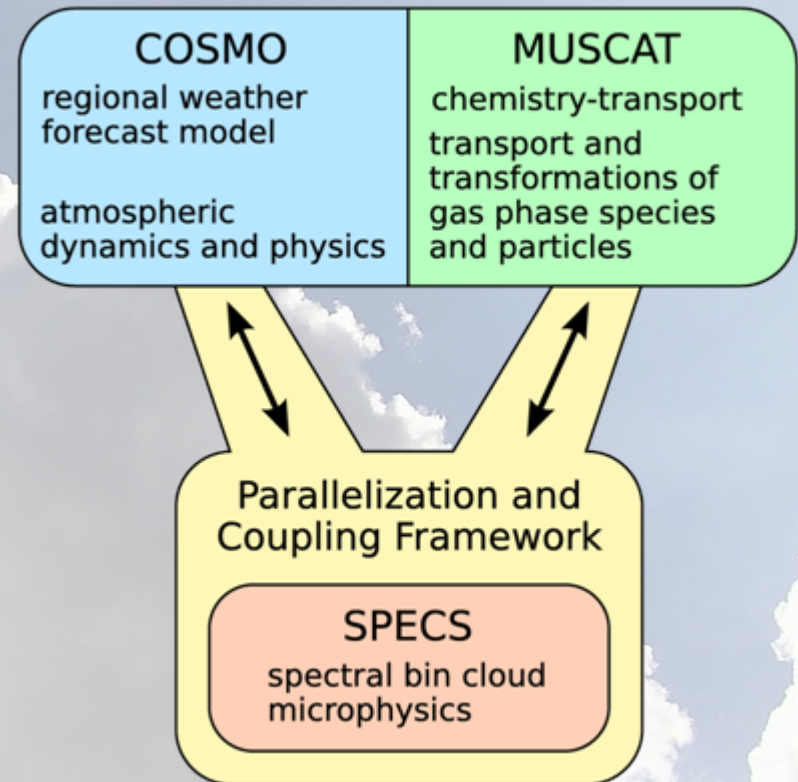


Cloud structure  
derived from  $q_c$   
and  $q_i$

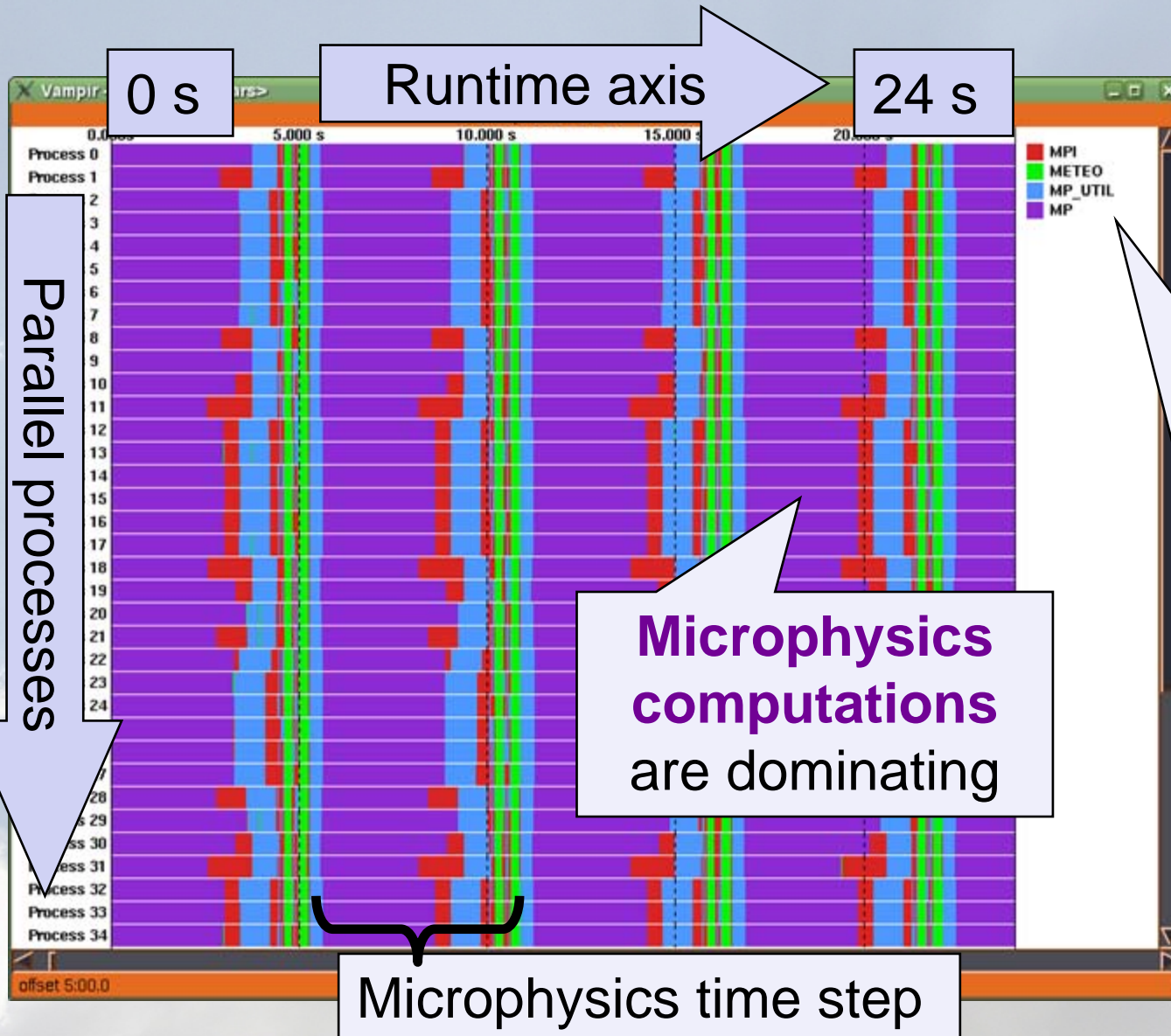
Adapted block  
structure with  $\sim 10,000$   
blocks (of max. 34,848)  
64 Partitions calculated  
by Hilbert SFC



- Parallel NetCDF I/O
- Use FD4 in COSMO-SPECS
- Multirate time stepping methods
- Couple MUSCAT to provide aerosol spectra and chemistry



# COSMO-SPECS seen in Vampir



No cloud existing

Activities:

**MPI**

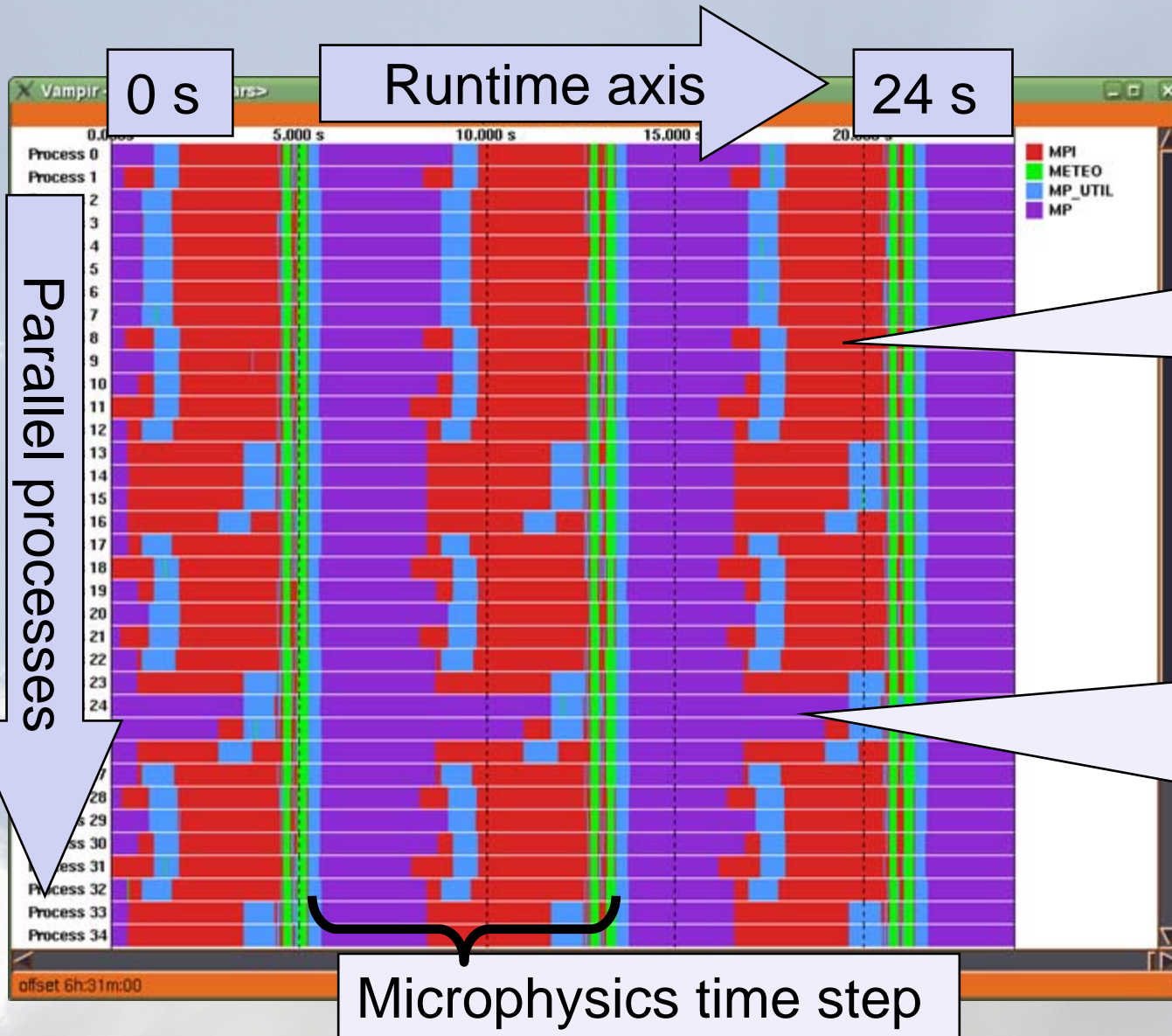
(comm & sync)

**Microphysics Computations**  
(SPECS)

**Microphysics Utilities**

**Meteorology**  
(COSMO)

# COSMO-SPECS seen in Vampir



**Cloud present**

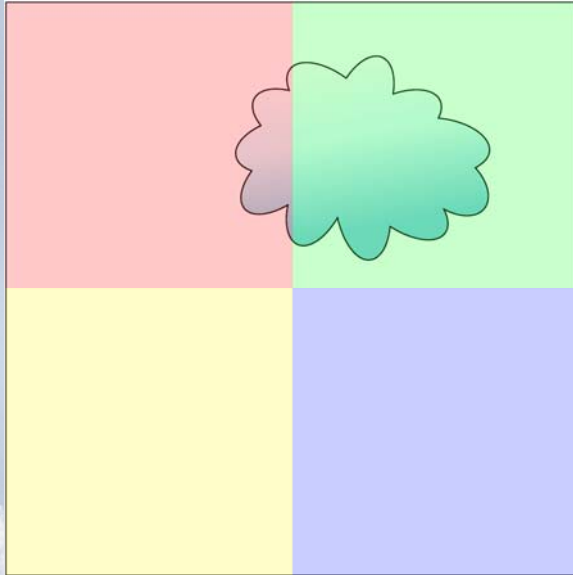
**MPI overhead**  
is large (40%)  
due to load  
imbalances

**Microphysics  
computations**  
are unbalanced:  
2 processes do  
substantially  
more work

**Workload  
follows clouds!**

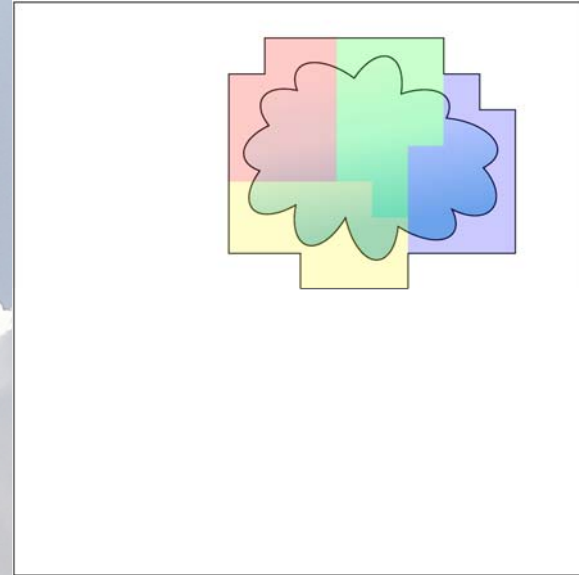
# Adaptive Block Structure

## COSMO-SPECS:



- Cloud causes high workload
- Static partitioning leads to load imbalance

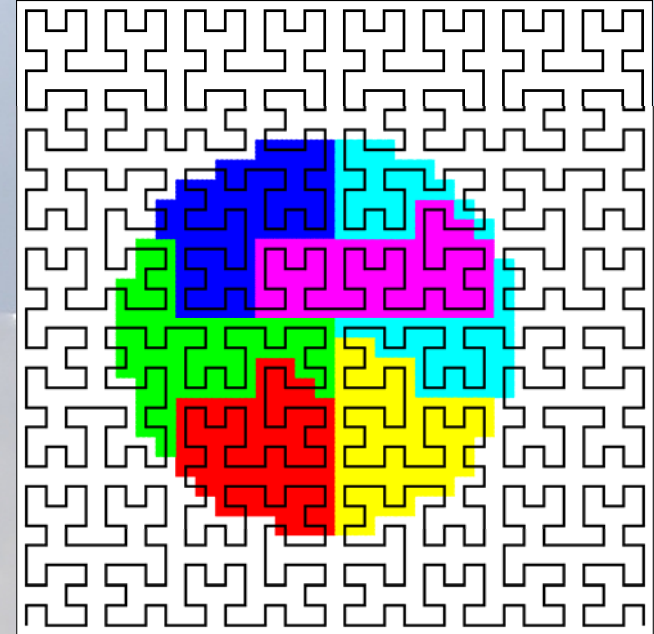
## Goal:



- Run cloud model only where clouds exist
- Adapt block structure and partitioning dynamically to actual cloud structure

# Load Balancing Methods

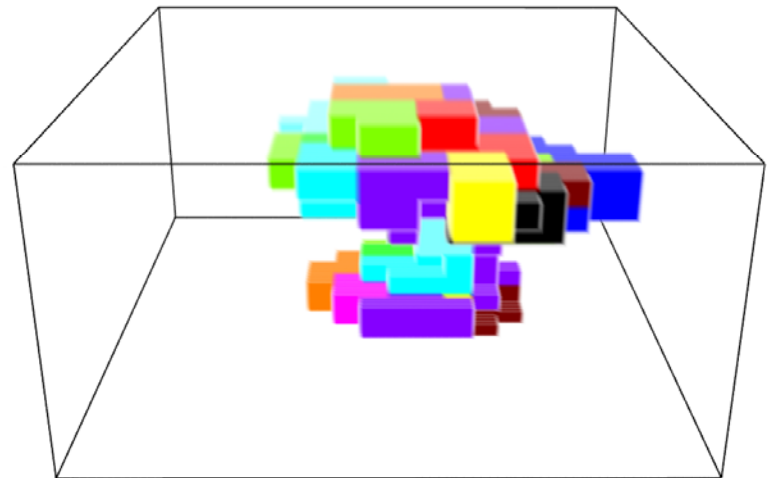
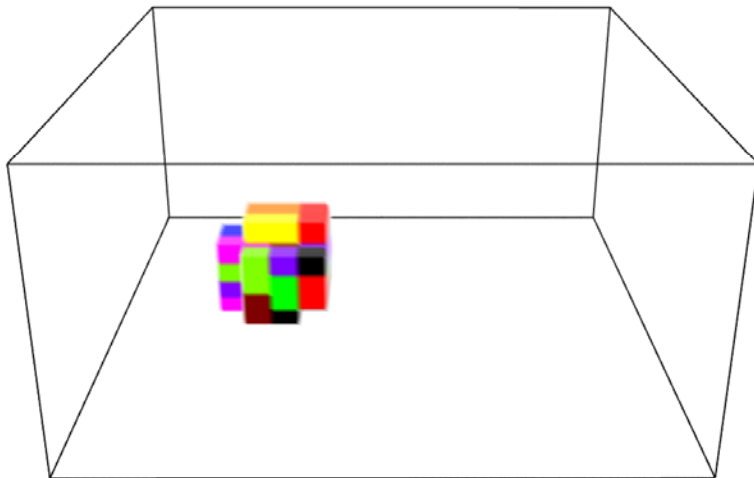
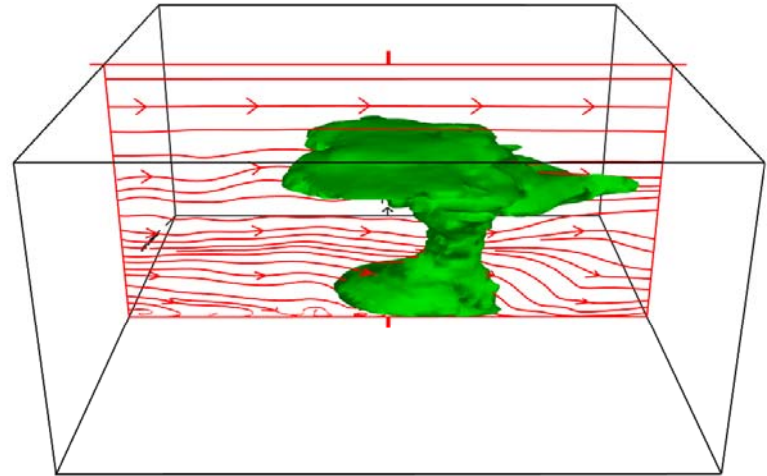
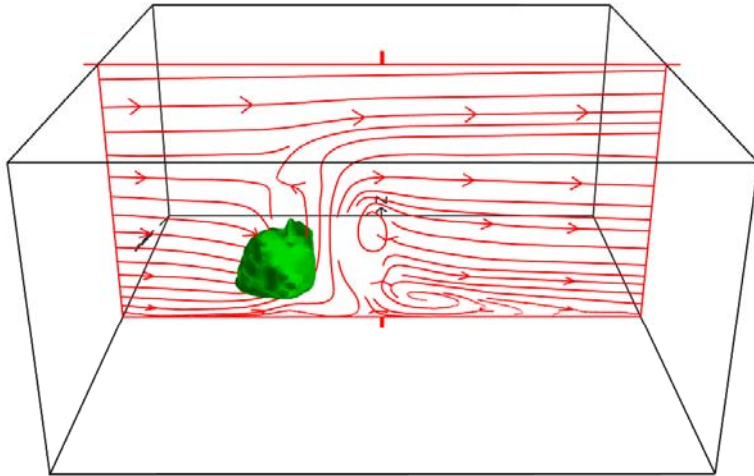
- Hilbert space filling curve
  - Continuous  $nD \rightarrow 1D$  mapping
  - Cut in equal-sized parts to obtain  $nD$  partitioning
- ParMetis
  - Library for parallel graph partitioning



# Test Application: Tracer Simulation

- Artificial case
  - Vertical wind
  - Heat bubble
- Tracer advection
  - Tracer stored in FD4 block structure
  - 2nd order advection scheme uses wind fields from COSMO
  - FD4 blocks and partitioning adapt dynamically to spreading of tracer

## Tracer and vertical wind



## FD4 blocks and 32 partitions