

# The implementation of the Klemp–Durrán–Bougeault radiative upper boundary condition in the LM with consistent time-split numerical scheme

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LM User Seminar 2006, Langen

# Outline

Radiative  
Upper  
Boundary  
Condition  
(RUBC)

Gassmann and  
Herzog

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# Motivation – Upper Boundary Condition

## Radiative Upper Boundary Condition (RUBC)

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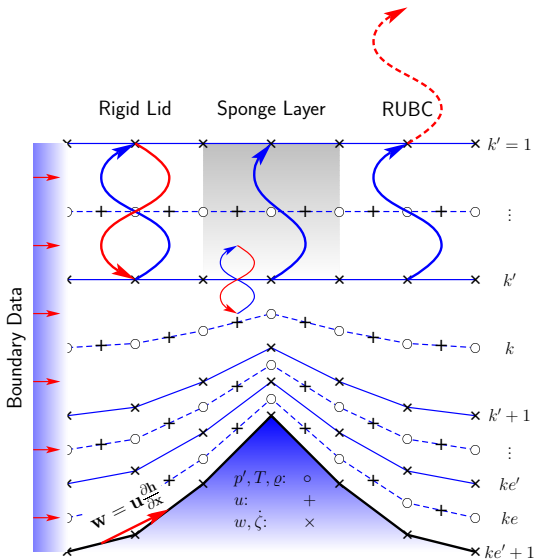
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- Rigid lid ( $w = 0$ ) with a sponge layer
- Spurious reflections
- "Useless" grid points
- Upper boundary values sometimes inaccessible

Alternative



RUBC

# Physical constraint at the upper boundary

Waves propagating upwards may be approximated as hydrostatic gravity waves with  $\omega = \pm \frac{Nk}{m}$  which gives us

$$\hat{w} = \frac{k}{N\varrho} \hat{p}'$$

in Fourier-space at the upper boundary for outward radiating energy (group velocity  $\frac{\partial\omega}{\partial m} > 0$ ).

This condition is local in time and global in horizontal space.

But our model is:

nonhydrostatic

compressible

nonlinear, inhomogeneous

limited area

rotating sphere

$$\omega = \pm \frac{Nk}{\sqrt{k^2 + m^2 + \frac{1}{4H^2}}}$$

sound waves ?

wave solutions ?

periodicity ?

Coriolis effects ?

# Historical Development

- Klemp and Durran (1983)
  - Application to a (non)hydrostatic (non)linear shallow(deep) atmosphere model without(with) Coriolis effects
  - Not suitable for Rossby waves
  - Implementation in an explicit numerical scheme
- Bougeault (1983)
  - Boussinesq hydrostatic model
  - Implementation in an implicit numerical scheme
  - Accounting for nonperiodicity by adding buffer zones
- Herzog and Binder (1995)
  - Implementation into the EM/DM/SM
  - Filtering of high frequent noise
  - Detrending of FFT-input fields
  - Operational use at SMA (Switzerland), and in versions of Hirlam and Eta models

# Historical development for the LM

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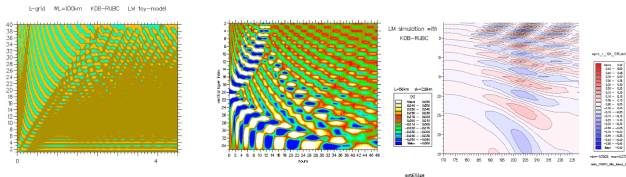
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Toy model (isothermal fast-waves model) studies  
successful with 2 methods:

- 1 Incorporation of the RUBC into the vertical implicit solver of the fast-waves system (internal method).
- 2 A direct method independent of a given time scheme (external method).

LM studies only partly successful with method (2).

# Reasons for failure in the LM ... and resort

- 1 The internal method (1) failed due to incomplete gravity wave representation in the fast-modes system. Fast-waves subsystem solved for quasi-base-state ( $p_0, T = const$ ) conditions.

→ ALM fast-waves equations solve for the current slow background state

- 2 The external method (2) relies on a back-lagging crudely interpolated pressure field which is not consistent with the implicit solution method.

→ Prefer method (1)

- 3 The usually applied upper layers are too thick for resolving wave structures. Waves were reflected before reaching the top.

→ Redefine layer geometry

# Method (1) for the ALM

ALM vs. LM:            fast waves            slow modes

$$\frac{\partial p'}{\partial t} - w \frac{1}{\sqrt{G}} \frac{\partial p_0 + p'}{\partial \zeta} + \frac{c_{pd}}{c_{vd}} p D = -\vec{v}_h \cdot \nabla_h p' + Q_p$$

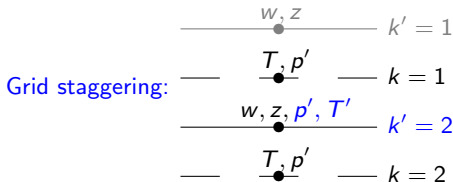
$$\frac{\partial T'}{\partial t} - w \frac{1}{\sqrt{G}} \frac{\partial T}{\partial \zeta} + \frac{1}{c_{vd}} \frac{p}{\varrho} D = -\vec{v}_h \cdot \nabla_h T + Q_T$$

3-diagonal matrix equation:

$$B_2 w_2^{\nu+1} + C_2 w_3^{\nu+1} = D_2 + \mu p'_{k'=2}{}^{\nu+1}$$

$$A_{k'} w_{k'-1}^{\nu+1} + B_{k'} w_{k'}^{\nu+1} + C_{k'} w_{k'+1}^{\nu+1} = D_{k'} \quad k' \in [3, ke' - 1]$$

$$A_{ke'} w_{ke'-1}^{\nu+1} + B_{ke'} w_{ke'}^{\nu+1} = D_{ke'} - C_{ke'} w_{ke'+1}^{\nu+1}$$



# Method (1) for the ALM continued

Gaussian elimination algorithm gives

$$w_2 = F_2 + G_2 p'_{k'=2}$$

This is unfortunately nonlinear! Linearize  $G_2$  and apply Fourier-transform and RUBC:

$$\hat{w}_2 = \hat{F}_2 + \bar{G}_2 \hat{p}'_2 = \hat{F}_2 + \bar{G}_2 \frac{\overline{N_{\rho_2}}}{K} \hat{w}_2, \quad K = \sqrt{k^2 + l^2}$$

At this point we get

$$\hat{w}_2 = \frac{K \hat{F}_2}{K - \overline{N_{\rho_2}} \bar{G}_2}$$

Back in grid point space that is the appropriate boundary condition for  $w$ .

# Fourier or Cosine transform?

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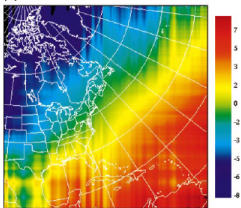
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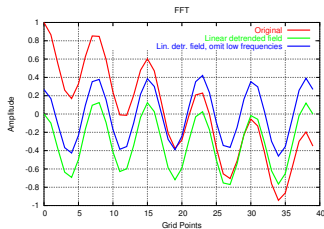
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Problem for real data: no periodicity, low frequency variability

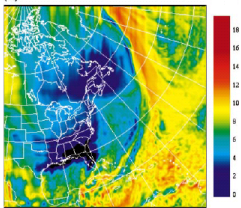
(a)



FFT on de-  
trended fields



(b)



Discrete  
Cosine  
Transform,  
DCT →

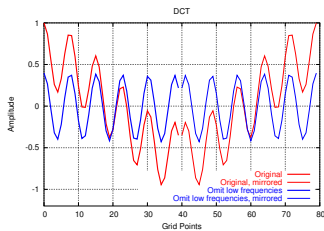


FIG. 3. (a) Trend component that is removed from the humidity field (Fig. 1) to get the (b) detrended field used as input for the detrending-DFT method. The line pattern is not generated by a printing problem but by the detrending itself.

Denis et al., 2002, MWR 130, 1812-1829

# Filtering

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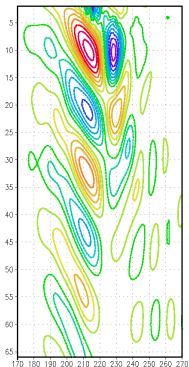
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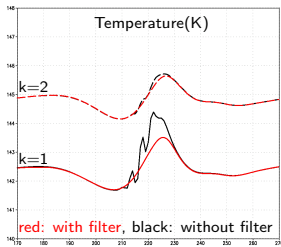
To avoid nonlinear instabilities, a **filter** is applied to the **vertical velocity** field and the **temperature** field.

Filter function in Fourier space:

w(m/s), incr=0.03m/s



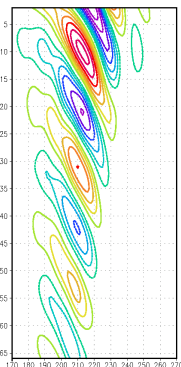
$$\frac{1}{1 + 0.1 \tan\left(\frac{k\Delta x}{2}\right)}$$



$\Delta x = 500m, \Delta z = 300m$

$\Delta t = 8s, h_{stop} = 5h$

w(m/s), incr=0.03m/s



# Redefine Layer Geometry

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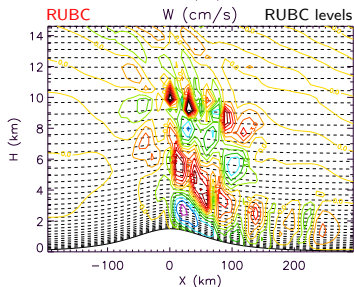
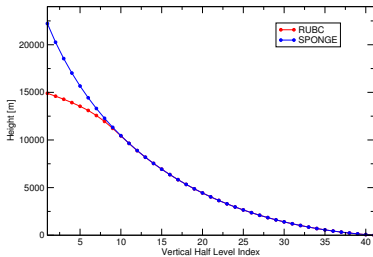
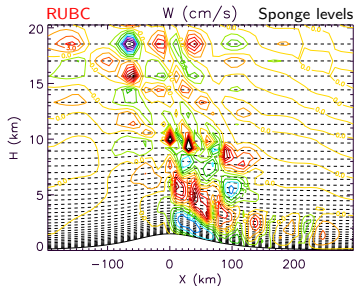
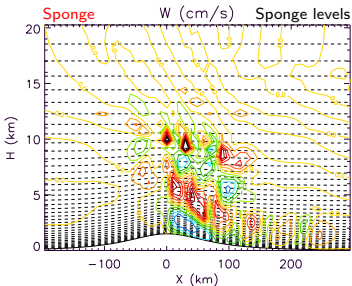
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# Case study – Initial model adjustment

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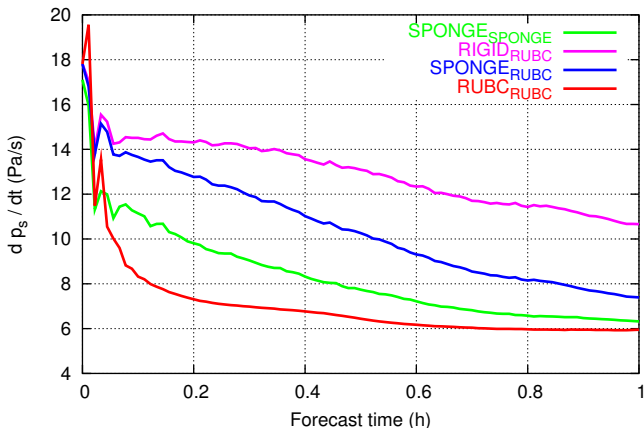
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21 January 2006, 325x325 grid points

$\Delta x=7\text{km}$ ,  $\Delta t=40\text{s}$ ,  $h_{\text{stop}}=24\text{h}$

Initial data interpolated from GME

Mean surface pressure tendency



# Case study – Surface pressure differences

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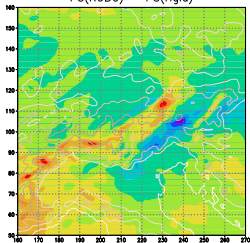
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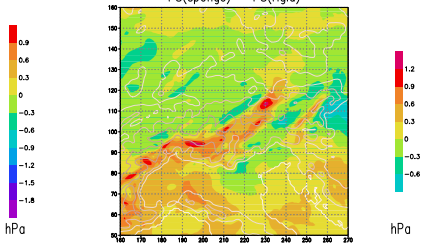
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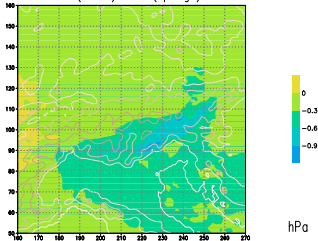
PS(RUBC) – PS(rigid)



PS(sponge) – PS(rigid)



PS(RUBC) – PS(sponge)



The rigid run uses RUBC levels. No wave pattern is visible in the difference between the RUBC run and the sponge run.

# Conclusions

- RUBC successfully implemented!
- Prerequisites:
  - ALM equations
  - Appropriate layer thickness
- Difficulties:
  - Nonlinearity  $\rightarrow$  filtering
  - Nonperiodicity  $\rightarrow$  Discrete Cosine Transform (DCT)
- Real case study is promising!
- To be done: code optimization for DCT