

Status report of

WG 2 – Numerical aspects

COSMO General meeting, Rome, Italy
09-12 Sept. 2019

Michael Baldauf, Daniel Reinert, Günther Zängl (DWD)

- PP EX-CELO → [Zbigniew Piotrowski](#)
- PT CCE final report → [Damian Wójcik](#)

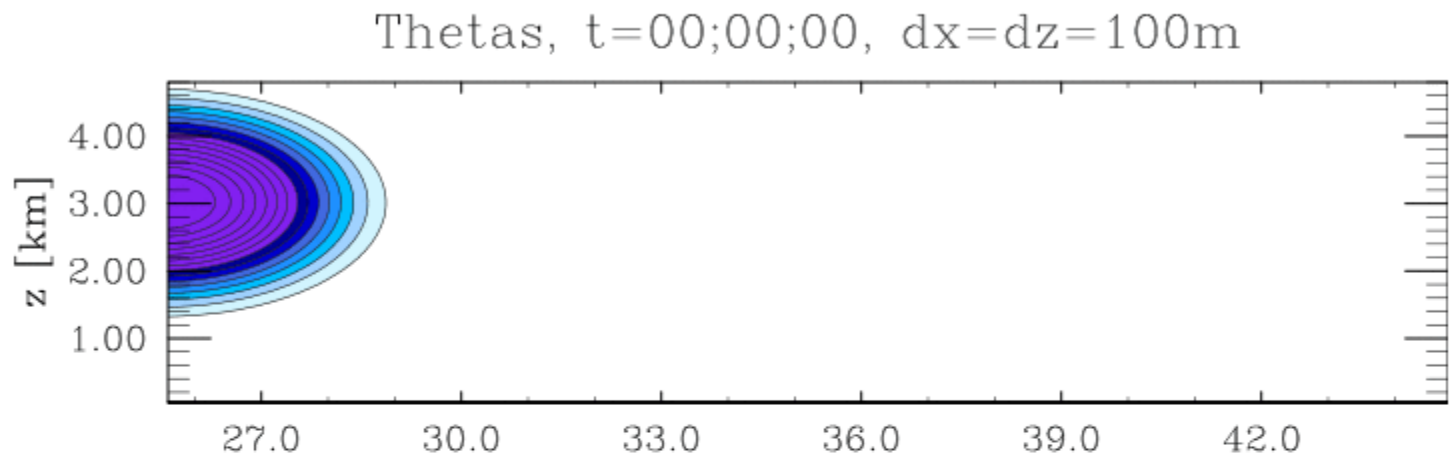
Addendum to the Priority Project ,Comparison between the dynamical cores of COSMO and ICON‘ (CDIC)

- PP CDIC officially finished in Aug. 2018
- however, the Straka et al. case did not work correctly
this is now solved (thanks to D. Reinert (DWD))
→ all relevant idealized test cases are working correctly with ICON
- The code for all the new test cases is available in the
`icon-nwp-dev-branch`
- Final report is still overdue
(until now, only contribution by D. Wójcik, (thank you!) received)

Test case 3: cold bubble

R. Dumitrache, A. Iriza (NMA), M. Baldauf (DWD)

Testsetup by *Straka et al (1993)*



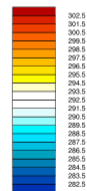
Test properties:

- test of dry Euler equations (without Coriolis force)
- unstationary
- strongly nonlinear
- comparison with reference solution from paper

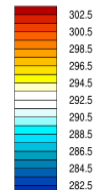
t=00:15:00, dx=dz=200m,

Theta t=00:15:00 dx=200m

Theta (in K)

 θ at t=15 min.for $\Delta x = 200, 100, 50, 25\text{m}$

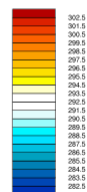
Theta



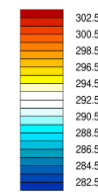
t=00:15:00, dx=dz=100m,

Theta t=00:15:00 dx=100m

Theta (in K)



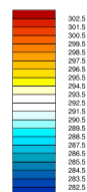
Theta



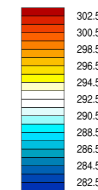
t=00:15:00, dx=dz=50m,

Theta t=00:15:00 dx=50m

Theta (in K)



Theta



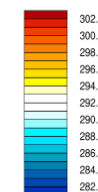
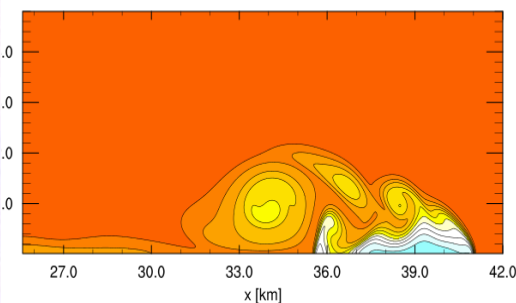
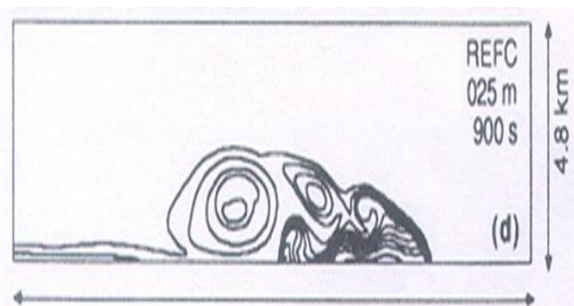
t=00:15:00, dx=dz=25m,

Theta t=00:15:00 dx=25m

Theta (



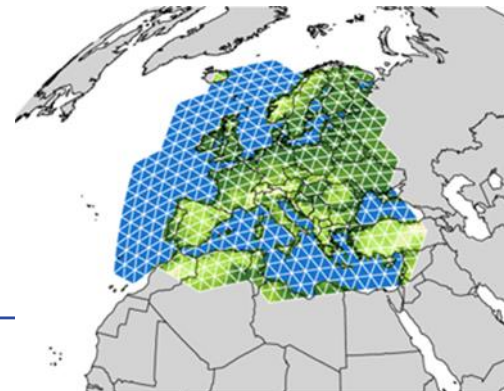
Theta

Reference solution
from Straka et al.:

Explore a more advanced vertical discretization for ρ , θ_v , and π

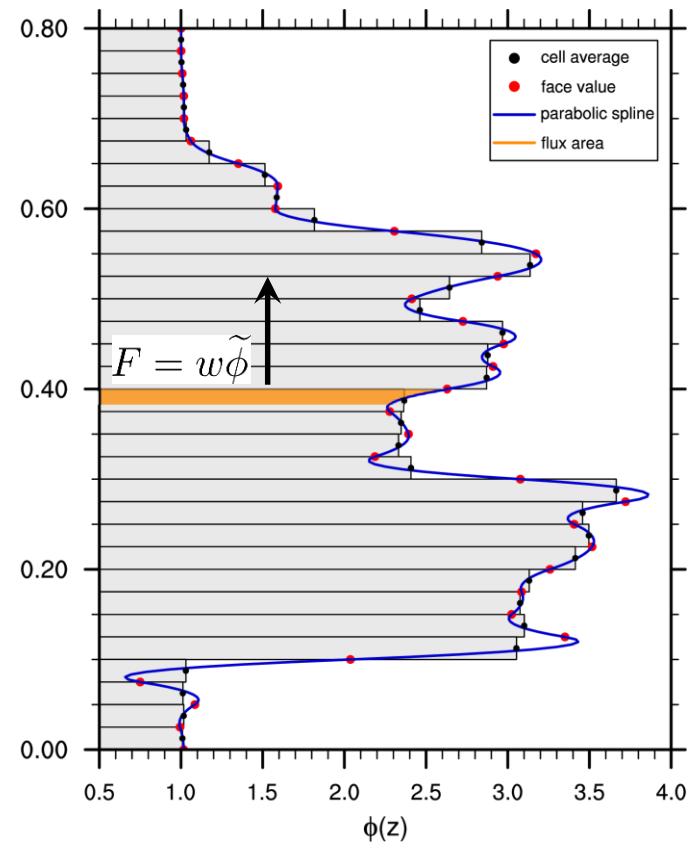
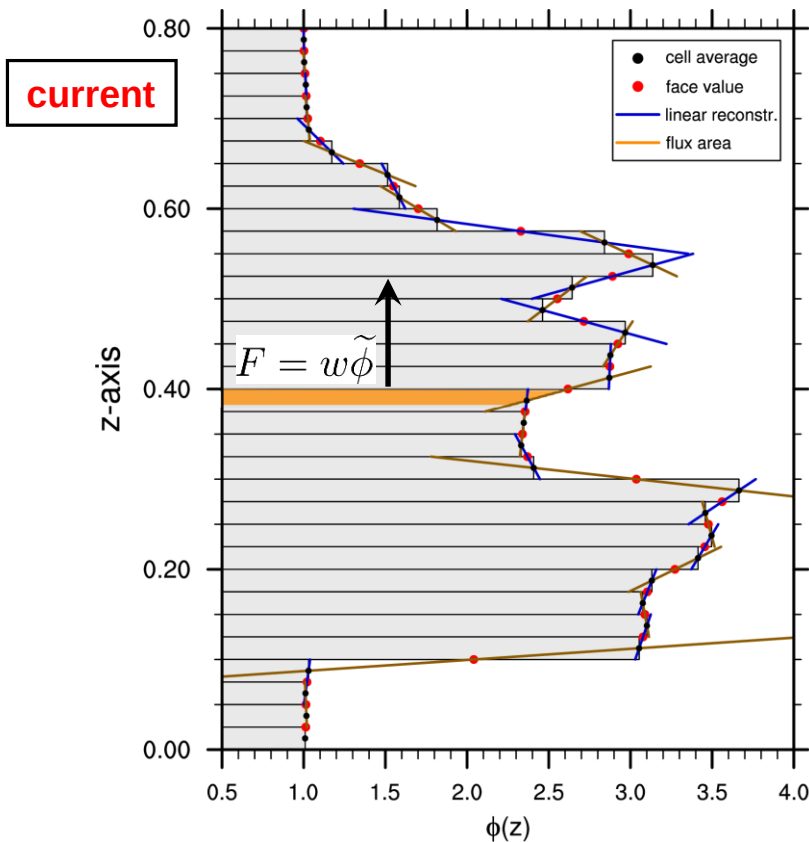
D. Reinert (DWD)

- Replace the 2nd order (linear)
 - vertical interpolation operator (cell to face)
 - vertical advective flux operator
- by 3rd order operators based on reconstructed parabolic splines
(Zerroukat et al., 2006)
- Why parabolic splines?
 - successfully tested and used for tracer transport in ICON (vertical PSM-scheme)
 - Operators are already available in ICON and ,only‘ have to be applied within the dynamical core



Old vs. new vertical discretization

Example: dycore-like reconstruction of an irregular 1D test signal *D. Reinert (DWD)*



2nd order reconstruction

Continuously differentiable at cell interfaces

Reconstruction not unique in cell interior

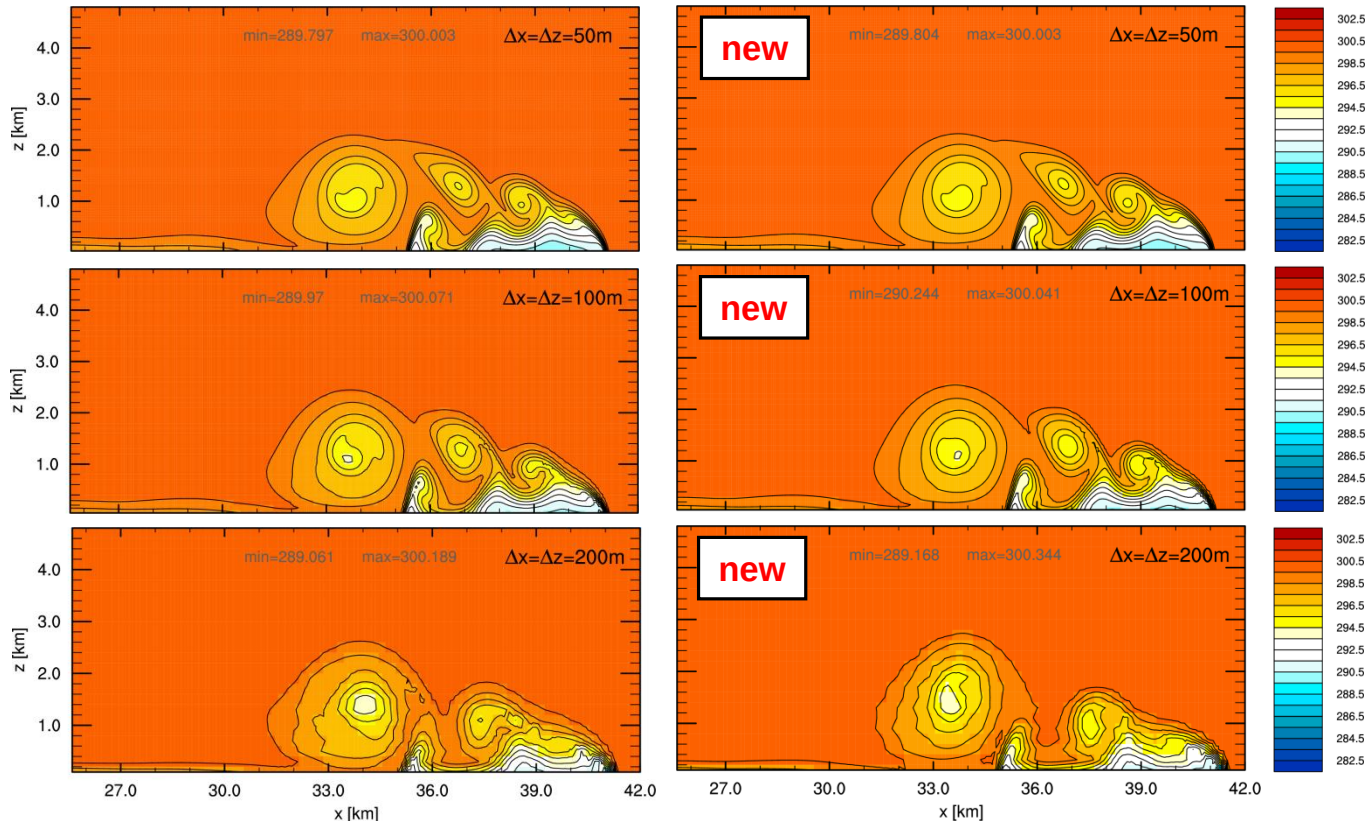


3rd order reconstruction

Continuously differentiable everywhere

Unique reconstruction

D. Reinert (DWD)



- similarity of the (almost) converged solution at 50m suggests that the 3rd order operators are implemented correctly.
- Unfortunately, only small improvements in simulation quality (if any) are noticeable (see e.g. middle rotor at $\Delta x = \Delta z = 200\text{m}$).

Implementation of the supercell detection index (SDI) into ICON

Wicker et al. (2005):

M. Baldauf, G. Zängl (DWD)

$$SDI_1(x, y) = \rho_{w\zeta}(x, y) \cdot \bar{\zeta}(x, y) \quad (SDI_2 \text{ similar})$$

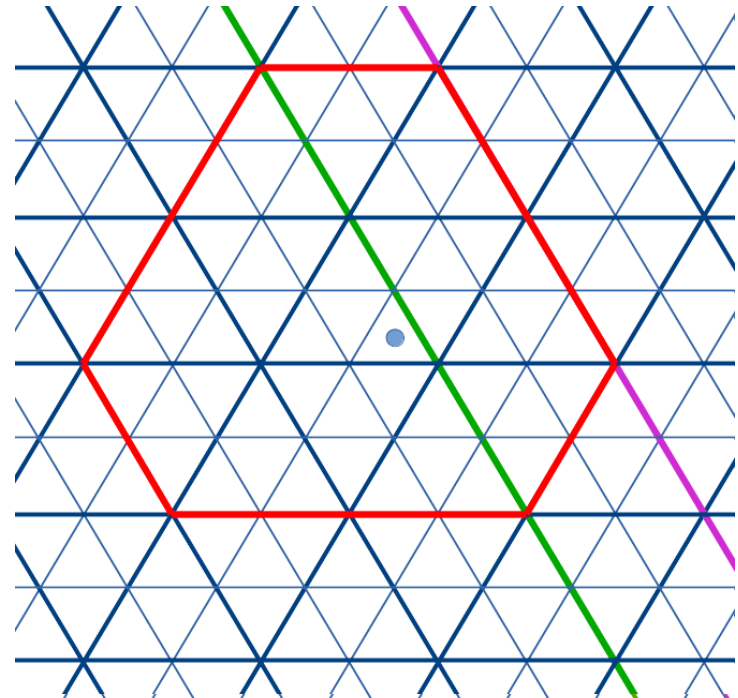
with the velocity-vorticity correlation:

$$\rho_{w\zeta}(x, y) = \frac{\langle w' \zeta' \rangle_{(x,y)}}{\sqrt{\langle w'^2 \rangle_{(x,y)}} \sqrt{\langle \zeta'^2 \rangle_{(x,y)}}}$$

<...> = volume average
= horiz. average of vertical averages

use the parent grid for a
larger horiz. averaging area!

1. calc. values on fine grid
2. average to parent grid
3. exchange parent cells
4. average on parent grid
5. write back to fine grid

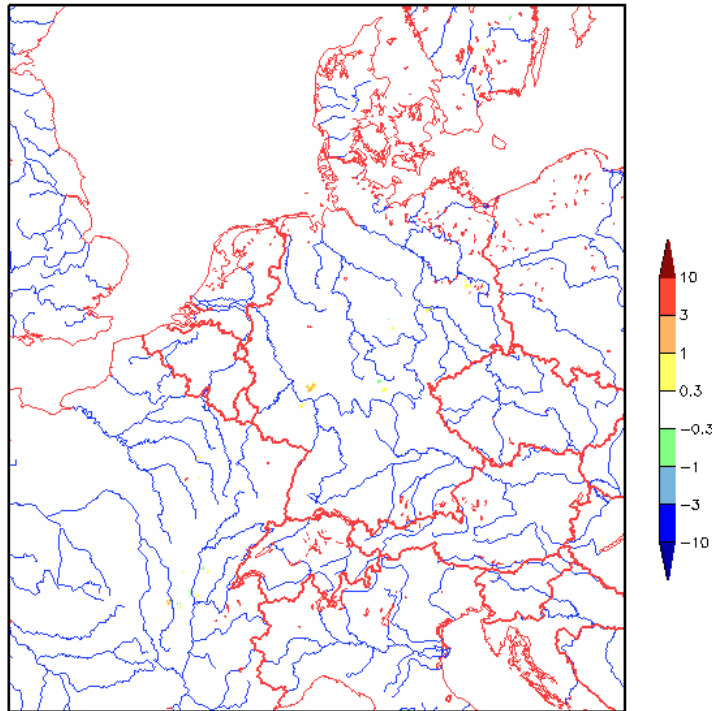


Similar averaging method is used for the lightning potential index (LPI)

Case study 18 Aug. 2019, SDI2 for a heavy storm event in the vicinity of Frankfurt

ICON-D2 (init. by ICON-EU)

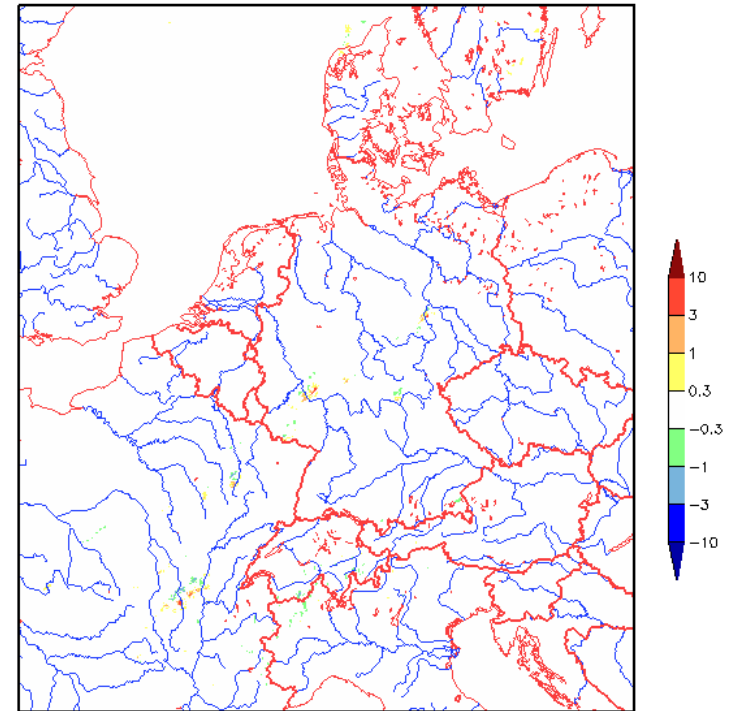
Start time: 18.08.2019 00:00 UTC ICON-D2 (urstart) 10647
Forecast time: 18.08.2019 16:00 UTC
SDI_2 [0.001 1/s]



SDI_2: Mean: 0.000988631 Min: -0.916585 Max: 2.85591 Sigma: 0.026618

operat. COSMO-D2

Start time: 18.08.2019 00:00 UTC COSMO-D2_Routine
Forecast time: 18.08.2019 16:00 UTC
SDI_2 [0.001 1/s]

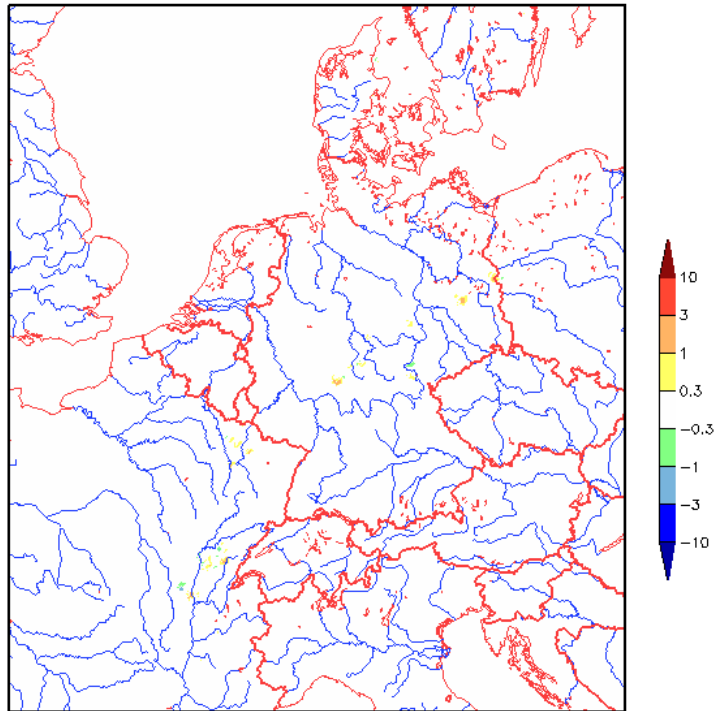


SDI_2(C-DE): Mean: -0.000966312 Min: -2.75059 Max: 6.41326 Sigma: 0.074752
SDI_2: Mean: 0.000285536 Min: -2.75059 Max: 6.41326 Sigma: 0.063663

Case study 18 Aug. 2019, SDI2 for a heavy storm event in the vicinity of Frankfurt

ICON-D2 (init. by ICON-EU)

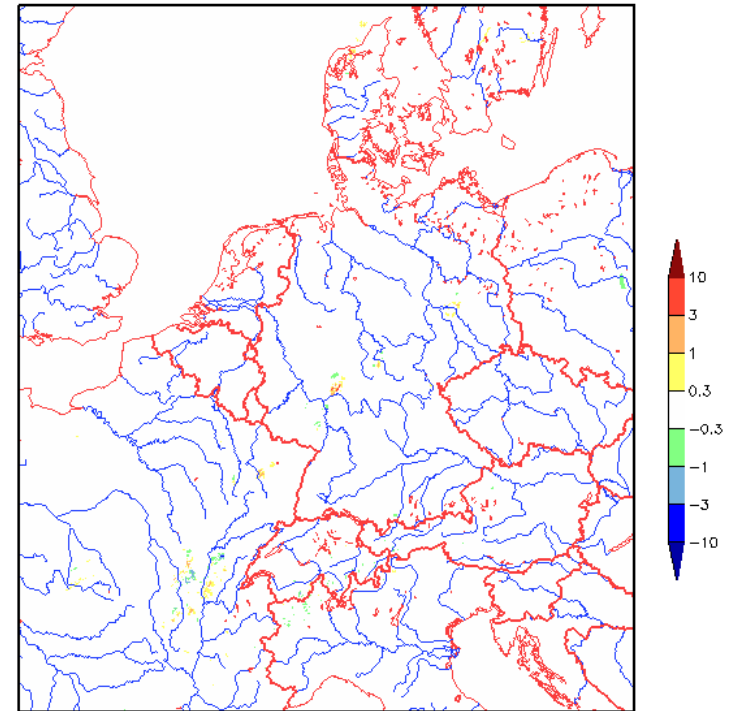
Start time: 18.08.2019 00:00 UTC ICON-D2 (urstart) 10647
Forecast time: 18.08.2019 17:00 UTC
SDI_2 [0.001 1/s]



SDI_2: Mean: 0.00125222 Min: -2.00685 Max: 2.62741 Sigma: 0.0360141

operat. COSMO-D2

Start time: 18.08.2019 00:00 UTC COSMO-D2_Routine
Forecast time: 18.08.2019 17:00 UTC
SDI_2 [0.001 1/s]



SDI_2(C-DE): Mean: -0.00146846 Min: -3.75508 Max: 4.60865 Sigma: 0.0646341
SDI_2: Mean: 0.000199474 Min: -3.75508 Max: 4.60865 Sigma: 0.0562111

Higher order discretization for COSMO

A. Will (Univ. Cottbus)

- currently: migration of the code from v5.0 to v5.6
unfortunately still bugs present in v5.6
(slow progress due to other tasks at BTU Cottbus)
- testing in hindcast mode and in the NUMEX-system
for a COSMO-D2 setup (summer case) must be done.
- However, expectations are:
max w ~ twice as large; more sound wave activity;
stronger diffusion properties (stronger PBL growth, ...)
→ probably no successful simulation without adaptation
at least of the turbulence scheme

A possible alternative dynamical core for ICON based on Discontinuous Galerkin Discretisation

Michael Baldauf (FE13)

MetStröm



Discontinuous Galerkin (DG) methods in a nutshell

$$\frac{\partial q^{(k)}}{\partial t} + \nabla \cdot \mathbf{f}^{(k)}(q) = S^{(k)}(q), \quad k = 1, \dots, K$$

weak formulation

$$\int_{\Omega_j} dx \, v(\mathbf{x}) \cdot \dots$$

Finite-element ingredient

$$q^{(k)}(x, t) = \sum_{l=0}^p q_{j,l}^{(k)}(t) p_l(x - x_j)$$

e.g. Legendre-Polynomials

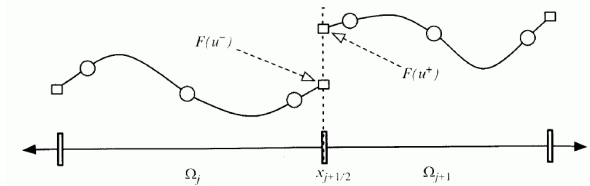
Finite-volume ingredient

$$\mathbf{f}(q) \rightarrow \mathbf{f}^{num}(q^+, q^-) = \frac{1}{2} (\mathbf{f}(q^+) + \mathbf{f}(q^-) - \alpha(q^+ - q^-))$$

Lax-Friedrichs flux

Gaussian quadrature for the integrals of the weak formulation

→ ODE-system for $q^{(k)}_{jl}$



From Nair et al. (2011) in
'Numerical techniques for global atm.
models'

e.g.

Cockburn, Shu (1989) *Math. Comput.*

Cockburn et al. (1989) *JCP*

Hesthaven, Warburton (2008):
Nodal DG Methods

- **local conservation**
- any **order of convergence** possible
- flexible application on **unstructured grids** (also dynamic adaptation is possible, h-/p-adaptivity)
- very good **scalability**
- **explicit** schemes are easy to build and are quite well understood
- higher accuracy helps to **avoid several awkward approaches** of standard 2nd order schemes: staggered grids (on triangles/hexagons, vertically heavily stretched), numerical hydrostatic balancing, grid imprints by pentagon points or along cubed sphere lines, ...

- high computational costs due to
 - (apparently) **small Courant numbers**
 - higher number of DOFs
- **well-balancing** (hydrostatic, perhaps also geostrophic?) in Euler equations is an issue (can be solved!)
- basically ,only‘ an **A-grid-method**, however, the ,spurious pressure mode‘ is very selectively damped!

Target system: ICON model

(Zängl et al. (2015) QJRMS)

- operational at DWD since Jan. 2015
(global (13km) and nest over Europe (6.5km))
- convection-permitting (2.2km): Q4/2020

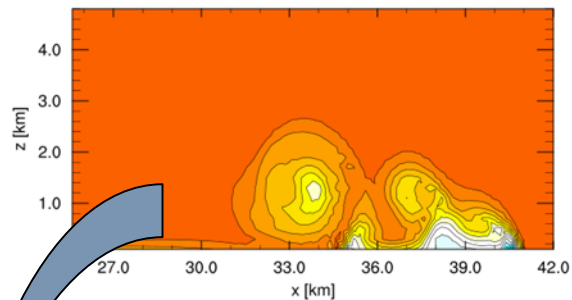


but currently far away from this, only a **toy model for 2D problems exists** with:

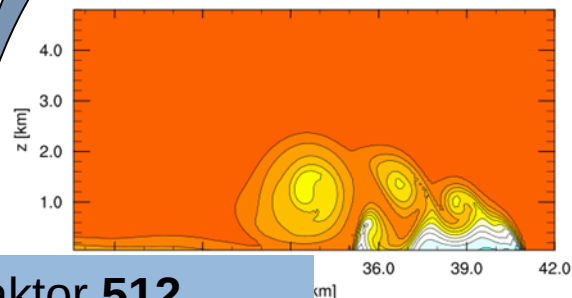
- explicit time integration DG-RK (with Runge-Kutta schemes) or horizontally explicit-vertically implicit (DG-HEVI) (with IMEX-Runge-Kutta)
- ‚local DG‘ (LDG) option for PDEs with higher spatial derivatives
- use of a triangle grid (also on the sphere) is optional

COSMO

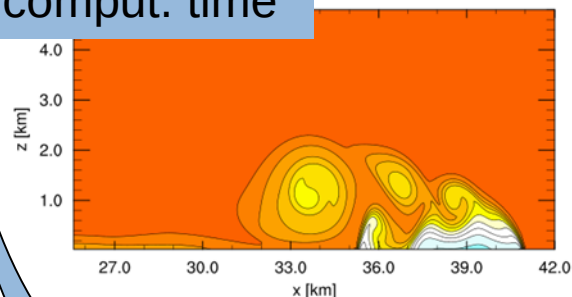
t=00:15:00, dx=dz=200m,



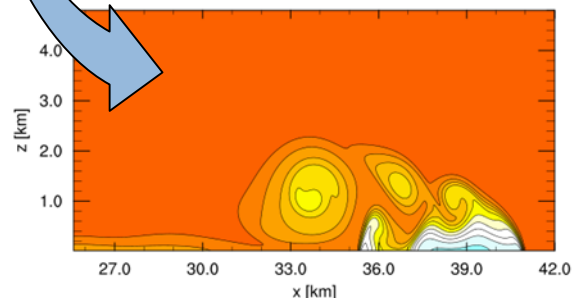
t=00:15:00, dx=dz=100m,



dx=dz=50m,



t=00:15:00, dx=dz=25m,



Faktor 512
in comput. time

Theta (in K)



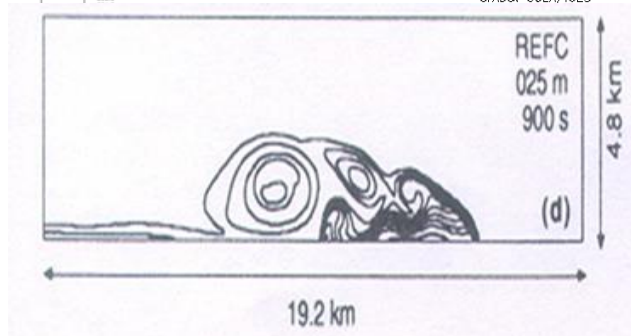
Theta (in K)



Reference solution
from Straka et al. (1993)



Theta (in K)



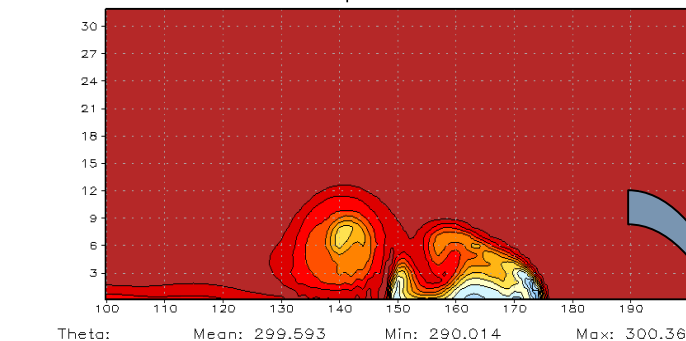
RK2-SLC
dt=0.0826446280992
dx=200.0
dy=200.0

DG explicit

2nd order

dx=dz=200m

Theta p=1 t=900.0

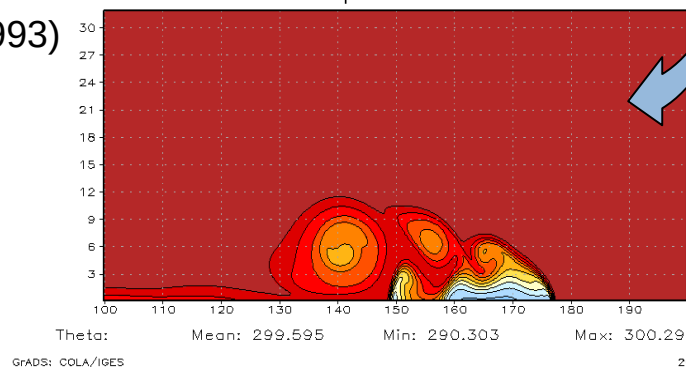


Faktor 4.3
in comput. time

3rd order

dx=dz=200m

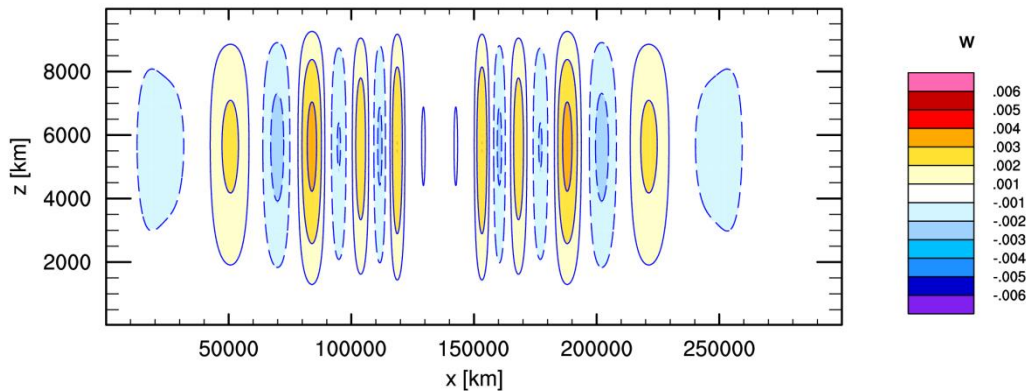
Theta p=2 t=900.0



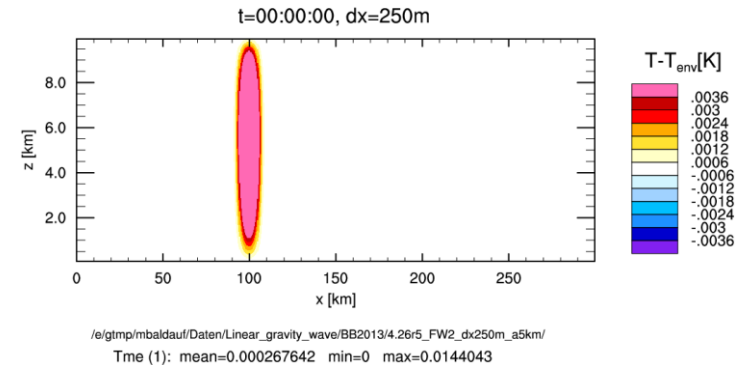
setup similar to *Skamarock, Klemp (1994) MWR*

$$\Delta x = 500\text{m}, \Delta z = 250\text{m}$$

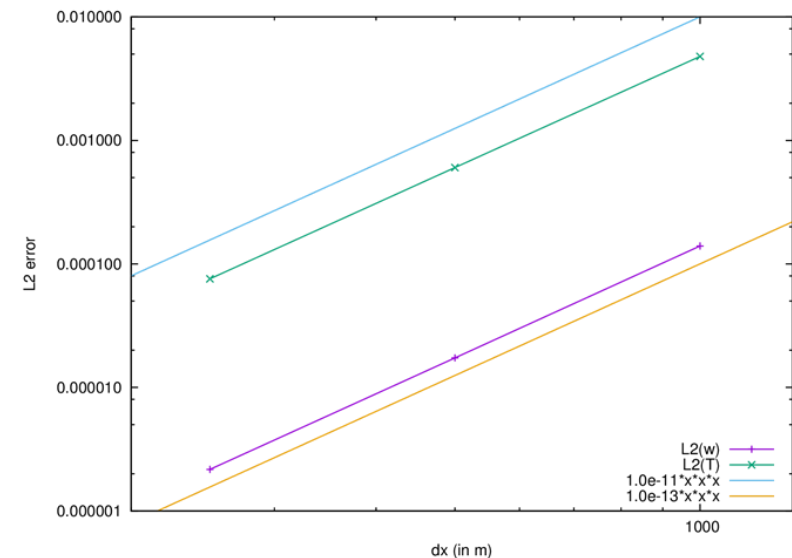
w, t=00:30:00



colors : simulation with $p=2/\text{RK3-SSP}$ (i.e. 3rd order explicit DG)
blue lines: analytic solution for compressible, non-hydrostatic Euler eqns. (*Baldauf, Brdar (2013) QJRMS*)



Exact 3rd order convergence for w and T':



Test case: flow over steep mountains, vertically stretched grid *Schaer et al. (2002) MWR (case 5b: $U_0=10\text{m/s}$, $N=0.01\text{ 1/s}$)*

$$H_{\text{oro}} = 1000\text{m},$$

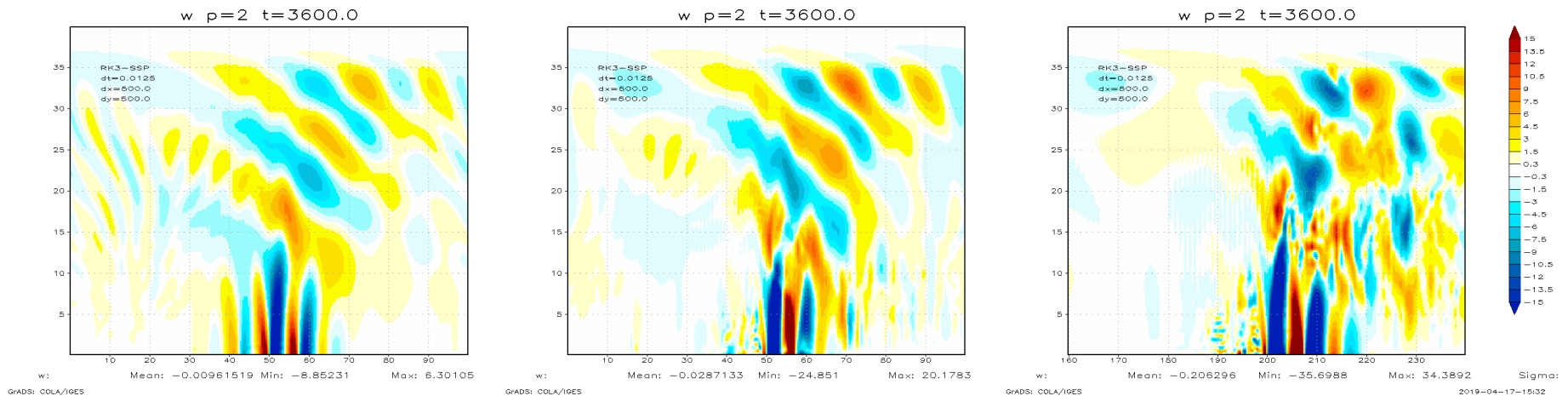
$$\alpha_{\text{max}} = 38^\circ$$

$$H_{\text{oro}} = 2000\text{m},$$

$$\alpha_{\text{max}} = 57^\circ$$

$$H_{\text{oro}} = 3000\text{m},$$

$$\alpha_{\text{max}} = 61^\circ$$



with vertical grid stretching $\sim 1:20$, $\Delta z_{\text{min}} \sim 50\text{m}$

Explicit DG simulation (3rd order) remains stable even for steeper slopes!
(remark: diffusion switched off $\rightarrow$ strong gravity wave breaking occurs)

HEVI-scheme with DG

$$\frac{\partial q^{(s)}}{\partial t} + \underbrace{\nabla \cdot \mathbf{f}_{slow}^{(s)}}_{\text{explicit}} + \underbrace{\nabla \cdot \mathbf{f}_{fast}^{(s)}}_{\text{implicit}} = \underbrace{S_{slow}^{(s)}}_{\text{explicit}} + \underbrace{S_{fast}^{(s)}}_{\text{implicit}}$$

$$\mathbf{f}_{fast}^{(s)} = f_{z,fast}^{(s)} \mathbf{e}_z$$

$$f_{z,fast}^{(s)} = \sum_{s'} H^{ss'} q^{(s')}$$

Treatment of numerical diffusion in the local Lax-Friedrich flux:

$$\left(\mathbf{f}_{slow}^{(s)} + \mathbf{f}_{fast}^{(s)} \right)^{(num)} \cdot \mathbf{n} = \frac{1}{2} \left(\mathbf{f}_{slow,+}^{(s)} + \mathbf{f}_{slow,-}^{(s)} + \underbrace{\mathbf{f}_{fast,+}^{(s)} + \mathbf{f}_{fast,-}^{(s)}}_{\text{implicit}} \right) \cdot \mathbf{n} - \frac{\lambda_{slow}}{2} \left(q_+^{(s)} - q_-^{(s)} \right) - \underbrace{\frac{\lambda_{fast}}{2} \left(q_+^{(s)} - q_-^{(s)} \right)}_{\text{implicit}}$$

Blaise et al. (2016) IJNMF

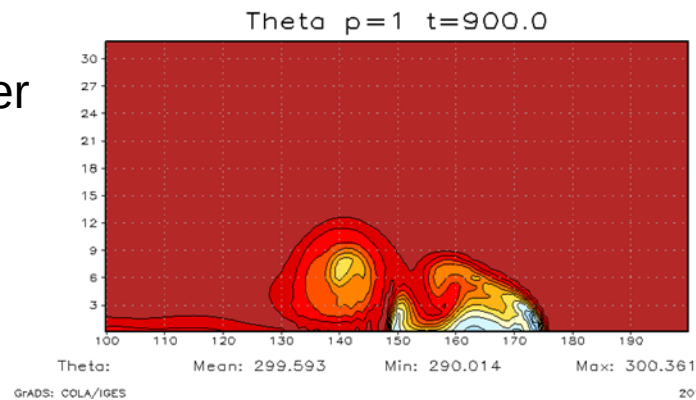
Test case: falling cold bubble (Straka et al. (1993))

Comparison explicit vs. HEVI scheme

2nd order

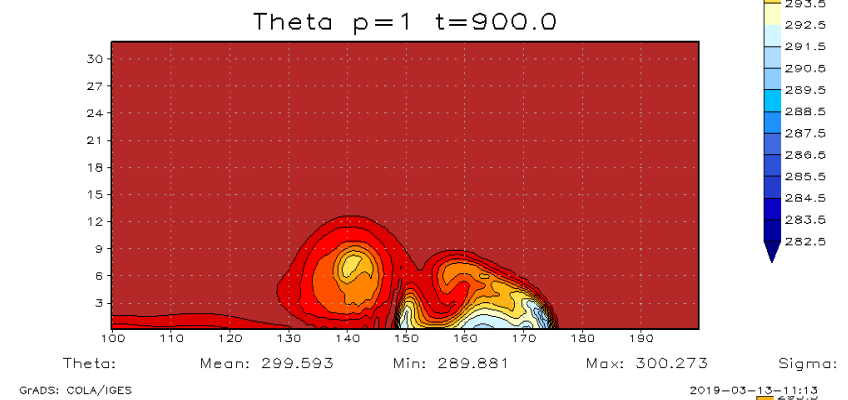
RK2-SLC
dt=0.0826446280992
dx=200.0
dy=200.0

DG explicit

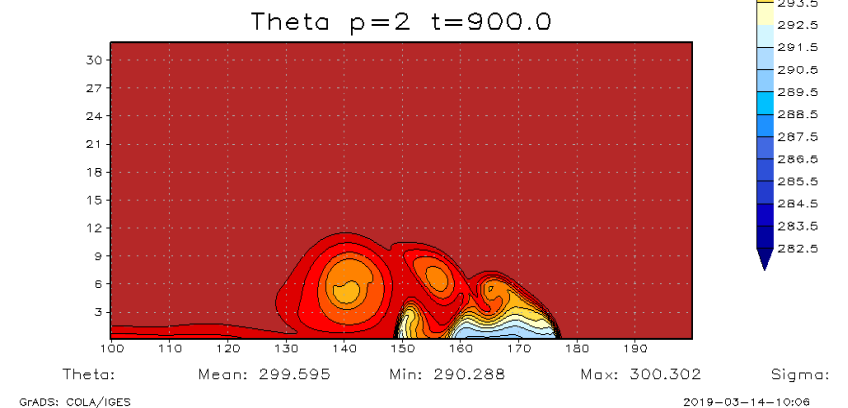
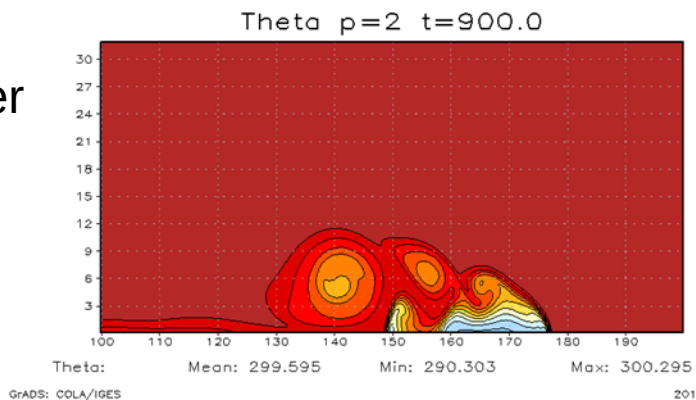


SSP3-3-3-2
dt=0.16393442623
dx=200.0
dy=200.0

DG HEVI



3rd order



How to bring DG on the sphere ...

Idea to avoid pole problem and to keep high order discretization:
use **local (rotated) coordinates** for every (triangle) grid cell,
i.e. rotate every grid cell towards $\lambda \approx 0$, $\varphi \approx 0$.

→ geometry is treated exactly

→ transform fluxes between neighbouring cells

shallow water equations

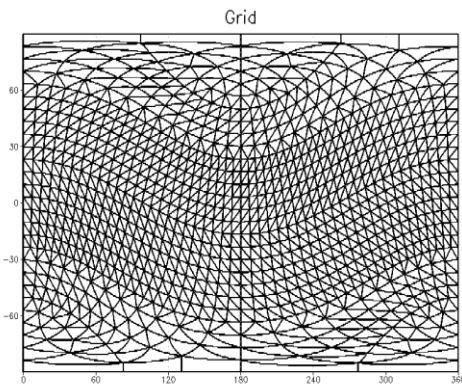
covariant formulation (here: without bathymetry)

$$\begin{aligned}\frac{\partial \sqrt{G}H}{\partial t} + \frac{\partial}{\partial x^i} \sqrt{G}m^i &= 0 \\ \frac{\partial \sqrt{G}m^i}{\partial t} + \frac{\partial}{\partial x^j} \sqrt{G}T^{ij} &= \sqrt{G}(F_{Cor}^i - \Gamma_{jk}^i T^{jk}) \\ T^{ij} &= \frac{m^i m^j}{H} + \frac{1}{2} g^{ij} g_{grav} H^2\end{aligned}$$

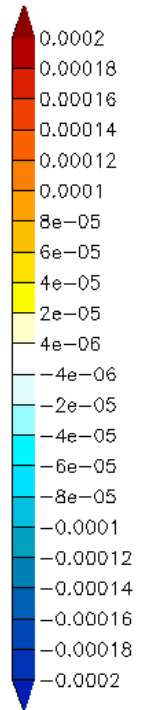
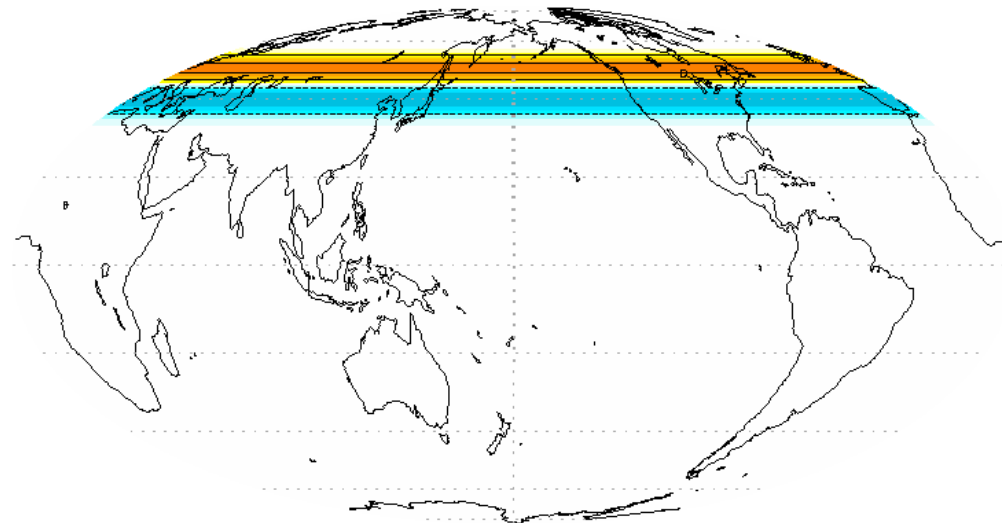
Barotropic instability test *Galewsky et al. (2004)*

4th order DG scheme
without additional diffusion
 $dx \sim 67$ km, $dt = 15$ sec.

simple triangle grid
on the sphere
 $dx \sim 500$ km:



rel. Vortic., ord=4 t=0d00h00m0.0s



relVort:

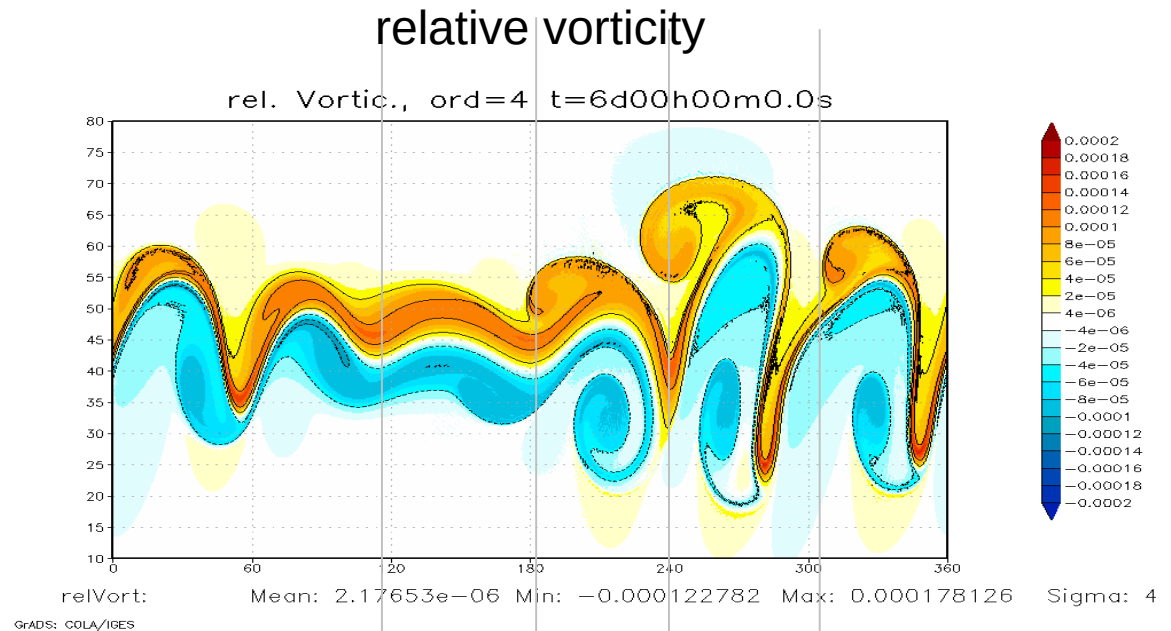
Mean: 7.53016e-07 Min: -9.8335e-05 Max: 0.000112421 Sigma: 2

GrADS: COLA/IGES

2019-09-03-17:03

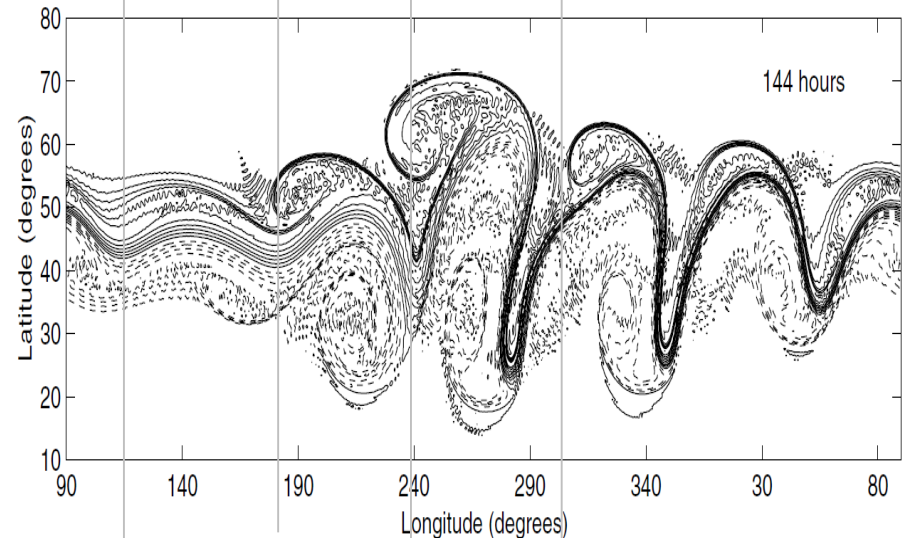
Barotropic instability test *Galewsky et al. (2004)*

4th order DG scheme
without additional diffusion
dx~67 km, dt=15 sec.



FMS-SWM (Geophys. Fl. Dyn. Lab.)
without additional diffusion
dx~60 km (T341), dt=30 sec.

Fig. 4 from *Galewsky et al. (2004)*

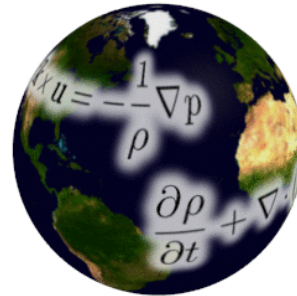


Summary

- **2D toy model** for
 - explicit DG-RK (on arbitrary unstructured grids with triangle or quadrilateral grid cells) and
 - **HEVI DG-IMEX-RK**works for several idealized tests (also Euler equations with terrain-following coordinates), correct convergence behaviour, ...
- **DG on the sphere** by use of local (rotated gnomonial) coordinates

Outlook

- **further design decisions:** nodal vs. modal, local DG vs. interior penalty vs. ..., ..
- coupling of **tracer advection** (mass-consistency)?
- improve **efficiency** in the HEVI direct solver
- further **milestones** (for the next years!)
 - development of a 3D prototype DG-HEVI solver
 - choose optimal convergence order p and grid spacing
estimated: $p_{\text{horiz}} \sim 3 \dots 6$, $p_{\text{vert}} \sim 3 \dots 4$ ($p_{\text{time}} \sim 3 \dots 4$)



PDEs

THE WORKSHOP ON PARTIAL
DIFFERENTIAL EQUATIONS ON THE SPHERE

Announcement:

The next

„Partial differential equations on the sphere“ – workshop

will take place at
DWD, Offenbach, Germany
5-9 October 2020