



# Implementation of stochastic physics in *COSMO*: recent tests

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*14°COSMO-General Meeting, Lugano, 10-13 September 2012*





# Outline

- ◆ Buizza stochastic physics and modifications
- ◆ Implementation in COSMO
- ◆ Experiment: 05 June 2011 case
- ◆ Conclusions





# Stochastic Physics

- Model uncertainty could be represented also with a stochastic physics scheme (Buizza et al, 1999; Palmer et al, 2009) implemented in the prognostic model
- This scheme perturbs physics tendencies by adding perturbations, which are proportional in amplitude to the unperturbed tendencies  $X_c$ :

$$X_p = (1 + r\mu)X_c$$

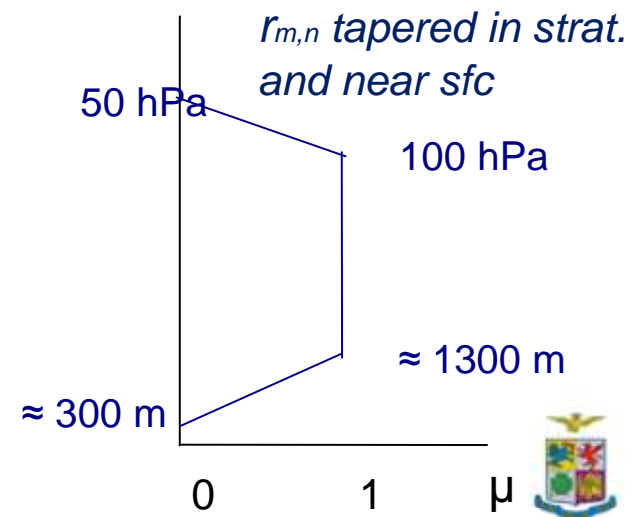
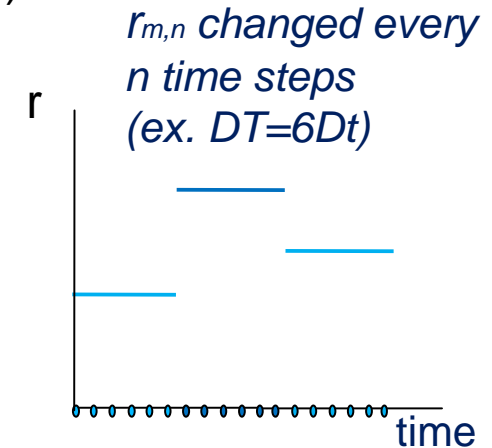
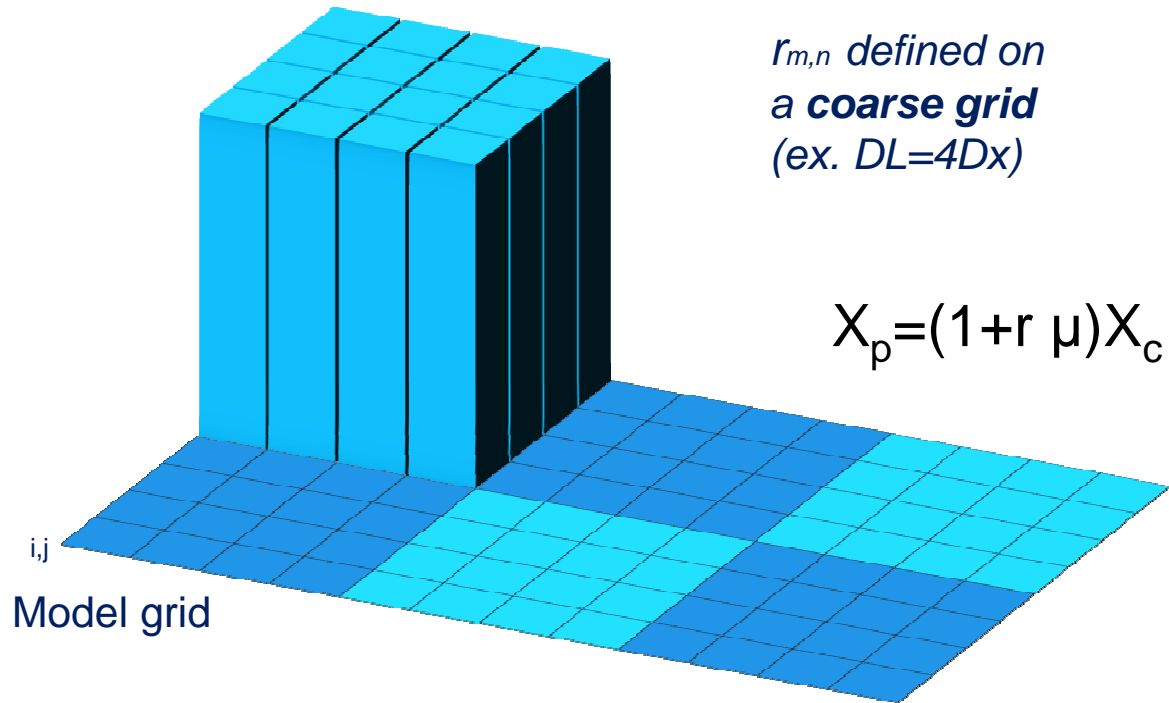
where  $r$  is a random number and  $\mu$  is a tapering factor ( $\mu=1$  in Buizza et al, 1999)





# Stochastic Physics

In Buizza et al. 1999: Spatial correlation is imposed using the same  $r$  in a whole column and drawing  $r$  for a coarse grid with spacing  $DL$  (boxes). Temporal correlation is achieved by drawing  $r$  every  $n$  time steps ( $DT$ )



Univariate distribution with reduced perturbation close to the surface and in the stratosphere (as in Palmer et al, 1999) modified to have a smoother pattern in time and horizontally in space



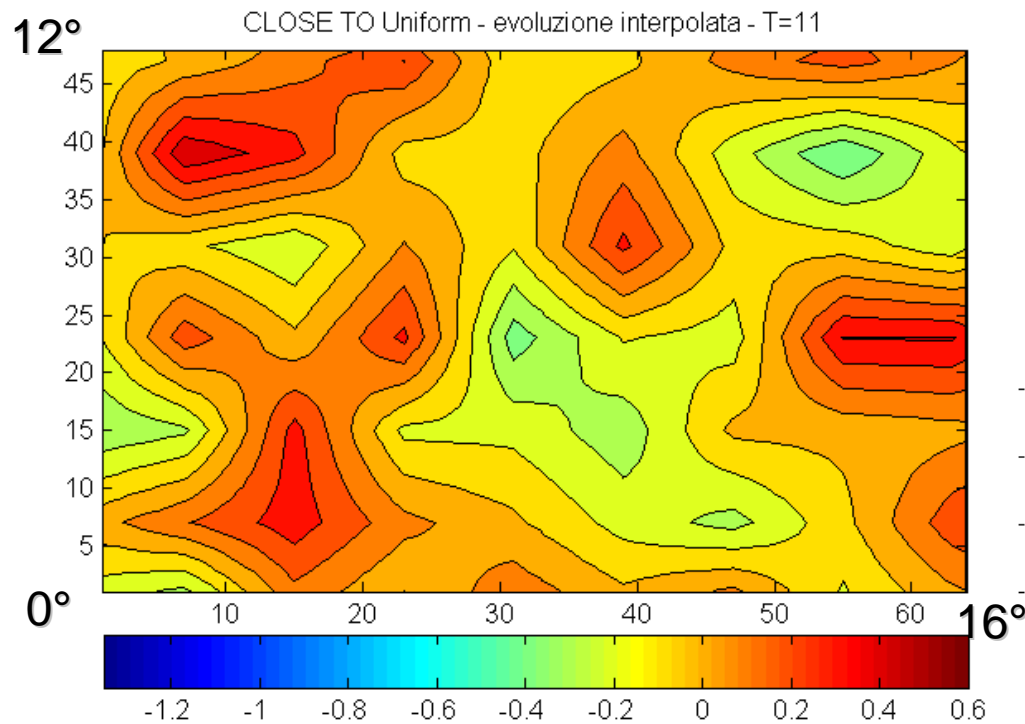


# Stochastic Physics

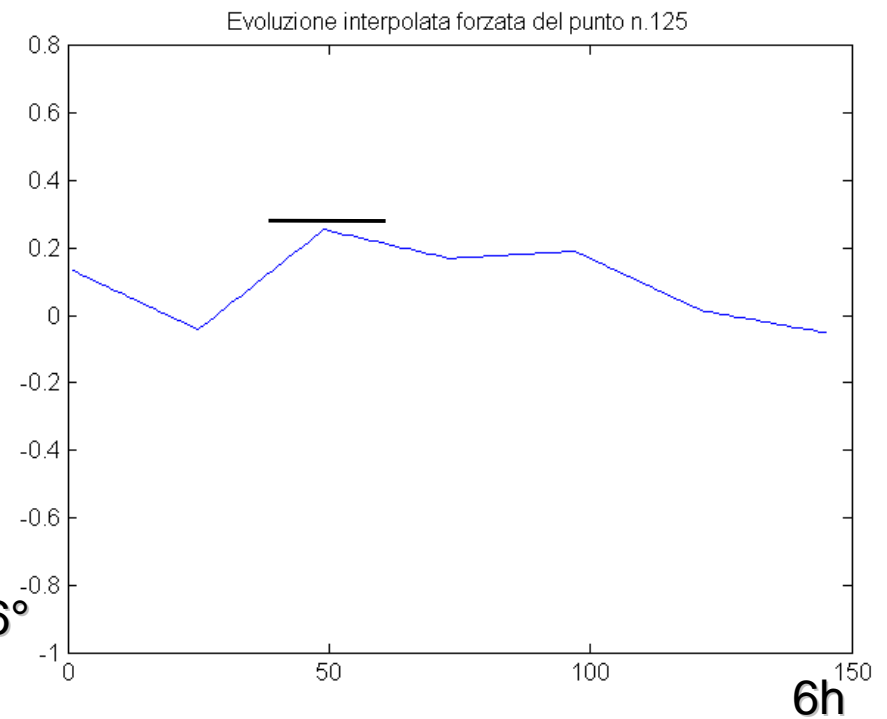
*Toy model and plots by A. Cheloni*

Model grid spacing:  $0.25^\circ$  (28 km)

Time step: 150 s



2.5° coarse grid with bilin. interp.



1h coarse time grid with lin. interp.





# Stochastic Physics

**Modified Version** (in blue, differences from Buizza et al, 1999)

- For **all variables** ( $u, v, T, qv$ ), the random numbers  $r$  are drawn from a uniform distribution in a certain range  $[-0.5, 0.5]$  **or a gaussian distribution with stdv (0.2-0.5) bounded to a certain value (range=  $\pm 2-3$  stdv)**
- A tapering factor  $\mu$  is used to reduce  $r$  close to the surface and in the stratosphere (Palmer et al, 2009)
- The perturbations of  $T$  and  $qv$  are not applied if they lead to particular humidity values (exceeding the saturation value **or negative values**)
- Spatial correlation is imposed using the same  $r$  in a whole column and drawing  $r$  for a coarse grid with spacing  $DL$  (boxes); **then they are *bilinearly interpolated* on the finer grid to have a smooth pattern in space**
- Temporal correlation is achieved by drawing  $r$  every  $n$  time steps ( $Dt$ ); **then they are *linearly interpolated* for the intermediate steps to have a smooth pattern in time**





# Stochastic Physics in COSMO

- Two new modules:
  - `random_numbers.f90` to generate machine-independent pseudo-random numbers (same to ECMWF version)
  - `stoch_physics.f90` to calculate the physics perturbations
- The stochastic physics is called by `organize_eps.f90`, if `lstoch_phys=.true.` in namelist `EPS_INPUT`.
- Other namelist parameters are:
  - `lqv_pertlim`, `lvtaper_rn` (perturbation limit)
  - `lhorint_rn`, `adlat_rn`, `adlon_rn` (horiz. interp,)
  - `ltimeint_rn`, `nfr_rn`, `hfr_rn` (time interp.)
  - `amag_rn` (uniform distribution)
  - `lgauss_rn`, `stdv_rn`, `range_rn` (gaussian distribution)





# Stochastic Physics in COSMO

- Updates:
  - Bug fixed in perturbations distribution
  - Microphysics tendencies perturbed
  - Reproducibility using restart option
  - Optionally namelist specification of seed number
  - SW corner of perturbation (coarse) grid changing with seed number
- Unresolved problems:
  - Moisture turbulent tendencies not perturbed (they should be calculated before the dynamics step)
  - Improve the criterion to evaluate, if a grid point has supersaturated or negative humidity (currently not perturbed) and introduce a decrease of the perturb.
  - Some physical tendencies could not be perturbed



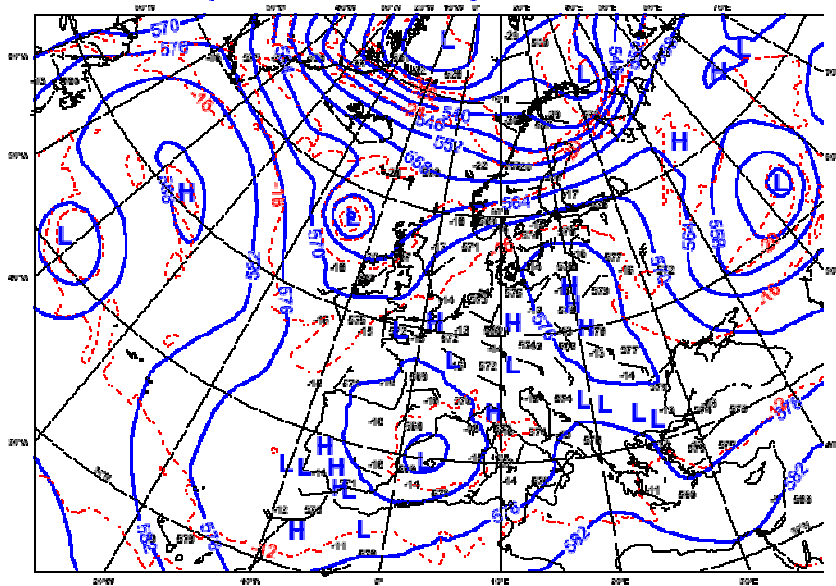




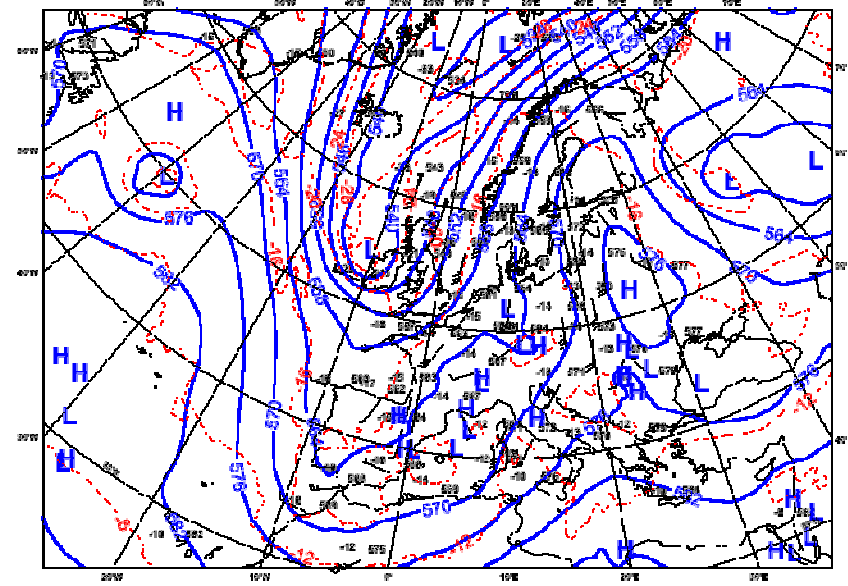
# Experiment: 05 June 2011 case

Situation over Italy: southwesterly flow from North Africa

**ROME Analysis VT:Domenica 5 Giugno 2011 00UTC**  
**Geopotenziale 500 hPa + Temperatura 500 hPa n.a.**



**ROME Analysis VT:Martedì 7 Giugno 2011 00UTC**  
**Geopotenziale 500 hPa + Temperatura 500 hPa n.a.**





# Experiment: 05 June 2011 case

COSMO-ME (7km)

$$X_p = (1 + r \mu) X_c$$

**10 members**

Options used:

leps = T

lstoch\_phys = T

lqv\_pertlim = T

lvtaper\_rn = T

lhorint\_rn = T

adlat\_rn = 5°

adlon\_rn = 5°

ltimeint\_rn = T

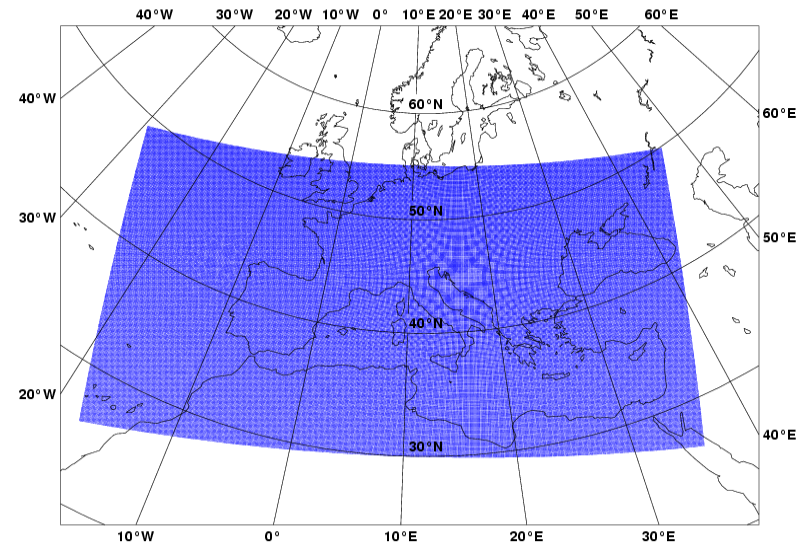
hfr\_rn = 6h

lgauss\_rn = T

stdv\_rn = 0.25,0.5

range\_rn = 0.75,1.

- no qv-T perturbation, if  $qv < 0$  or  $qv > qvs$
- stratosph. / boundary layer tapering of random numbers  $r$  (define  $\mu$ )
- random numbers horizontal interpolation
- same random number for a spatial box  $5^\circ \times 5^\circ$
- random numbers time interpolation
- new random numbers every 6h
- random numbers from gaussian distribution
- standard deviation of random numbers from gauss. distr.
- cutoff value of random numbers from gauss. distr.



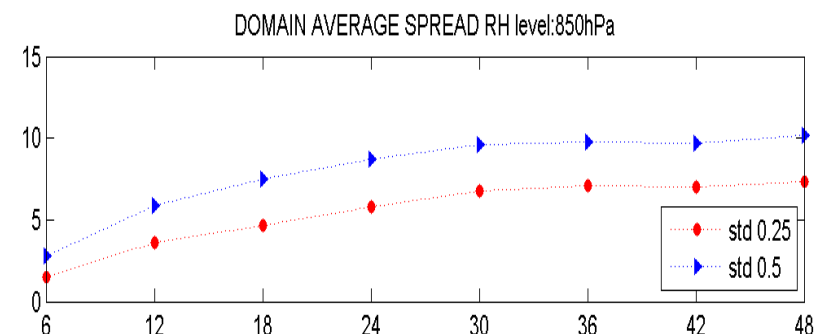
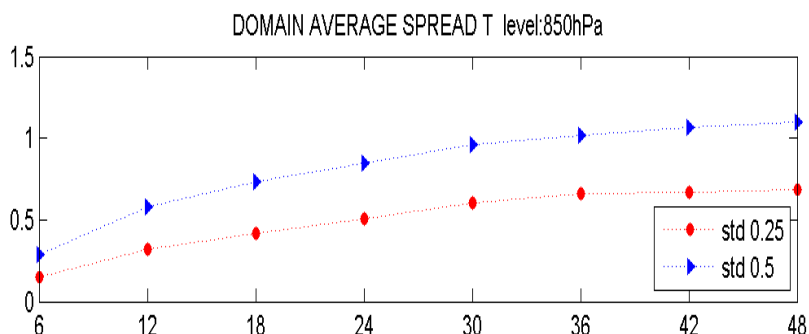
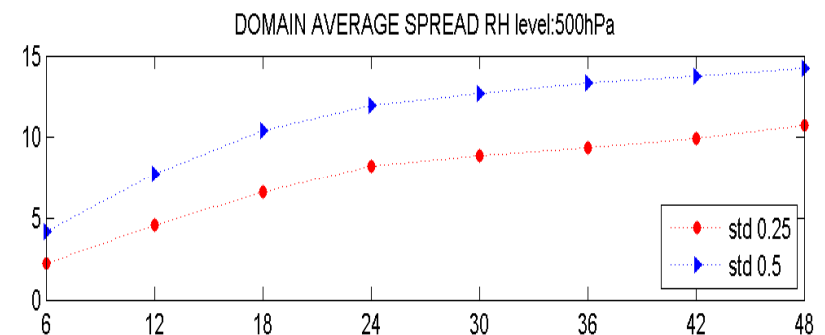
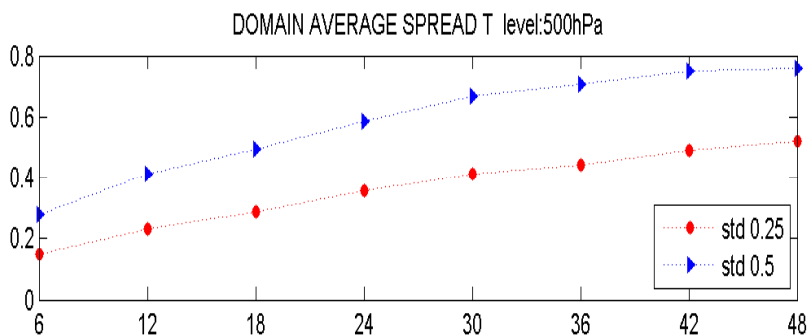
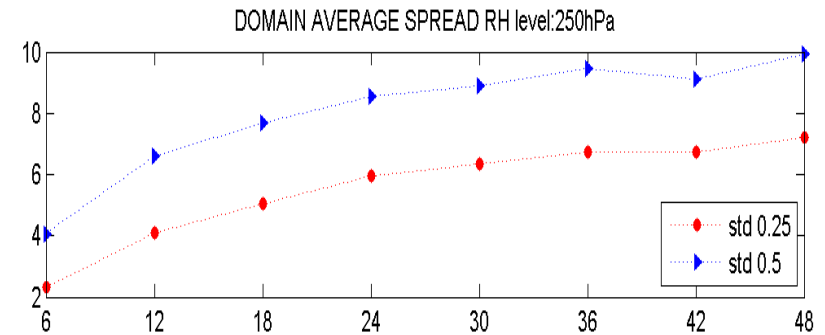
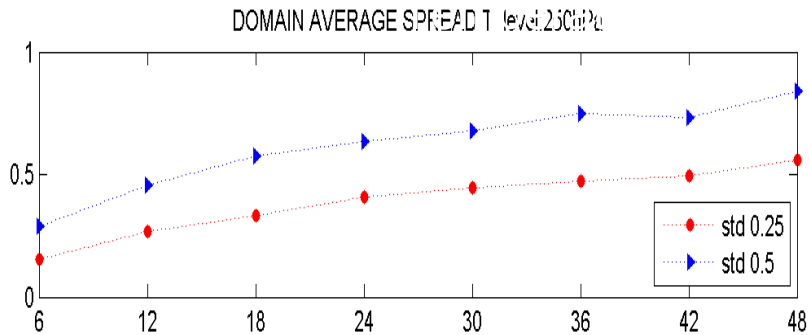


# Experiment: 05 June 2011 case

## Domain Averaged Spread for 10 members

### Temperature

### Relative Humidity



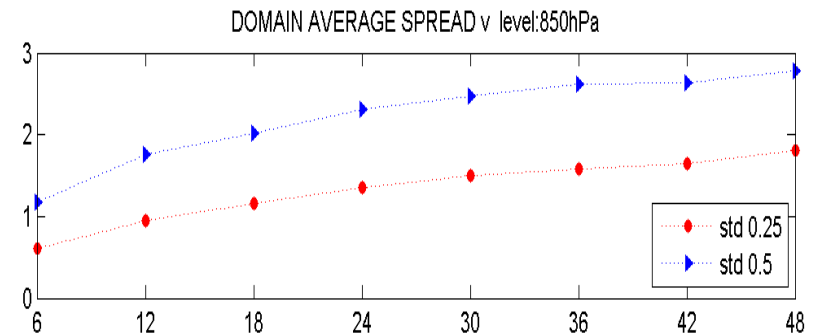
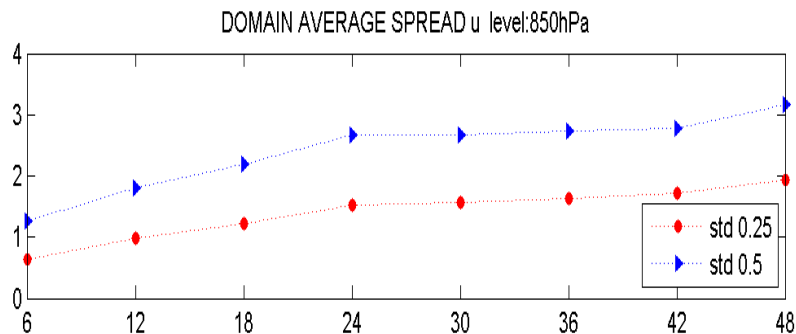
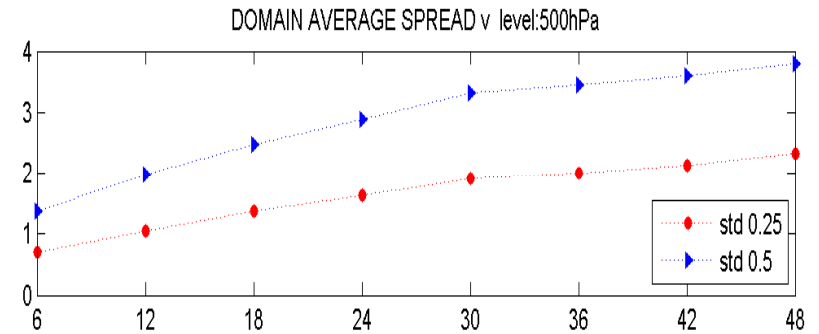
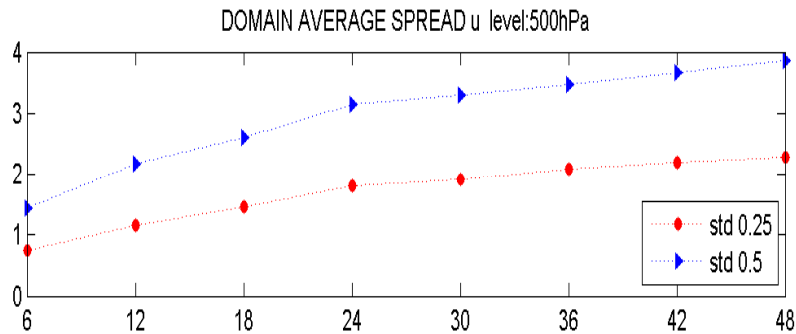
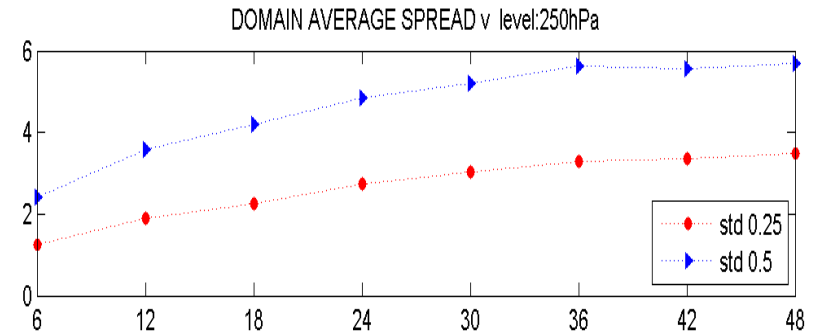
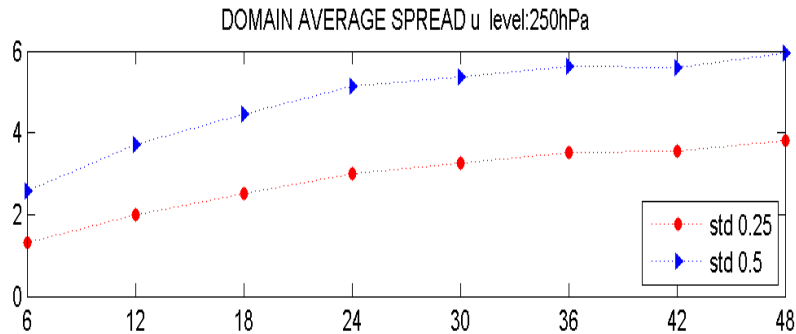


# Experiment: 05 June 2011 case

## Domain Averaged Spread for 10 members

### Zonal wind

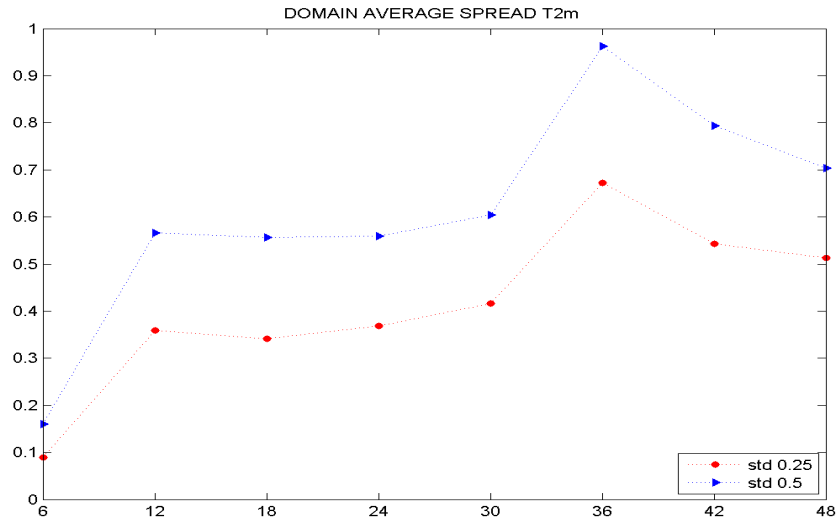
### Meridional Wind



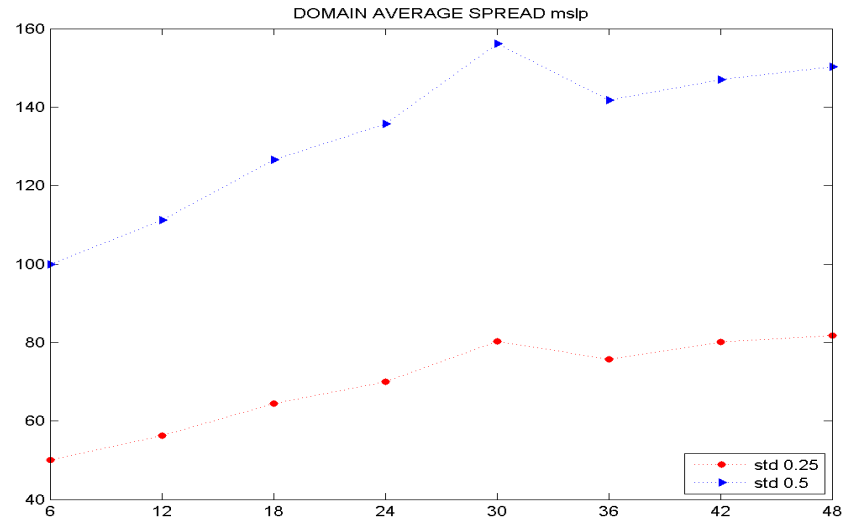


# Experiment: 05 June 2011 case

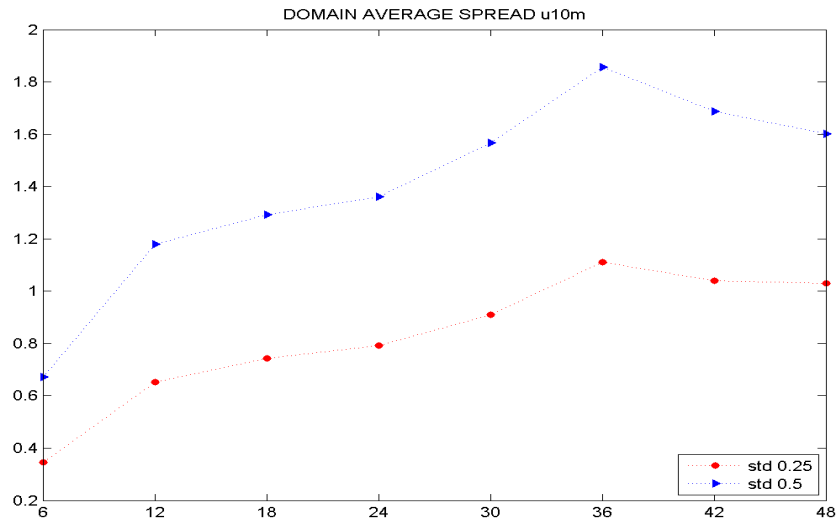
## Temperature



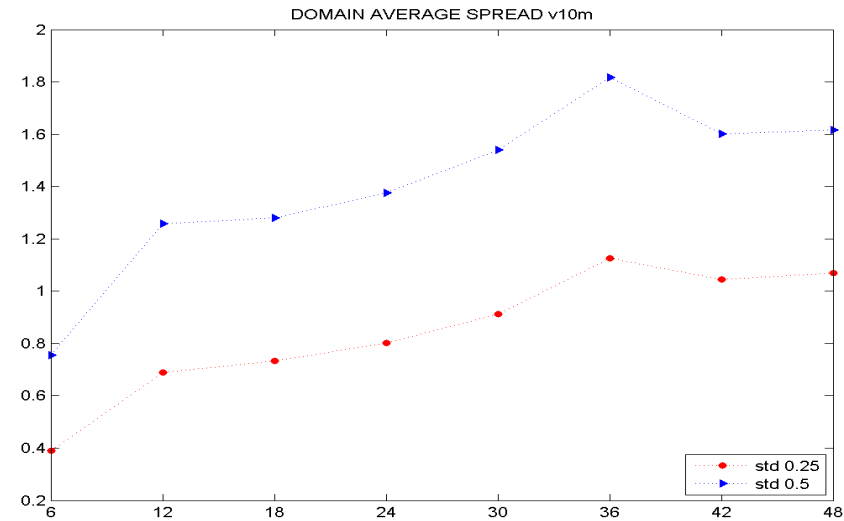
## MSL Pressure



## 10m Zonal Wind



## 10m Meridional Wind





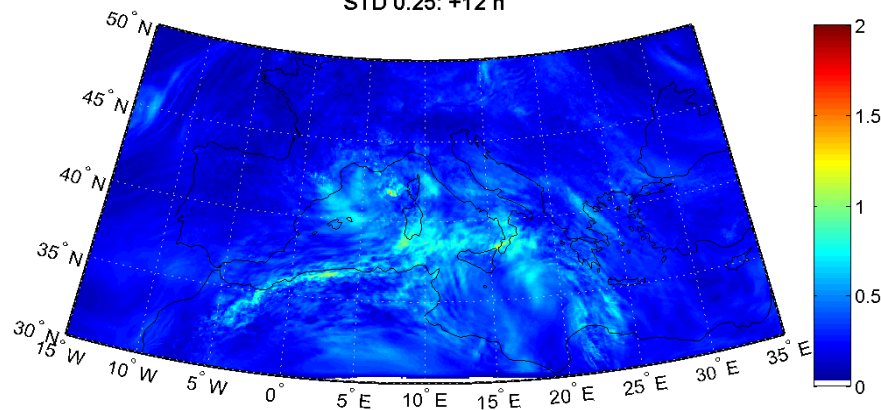
# Experiment: 05 June 2011 case

500 hPa Temperature Spread for 10 members

stdv=0.25 range=0.75

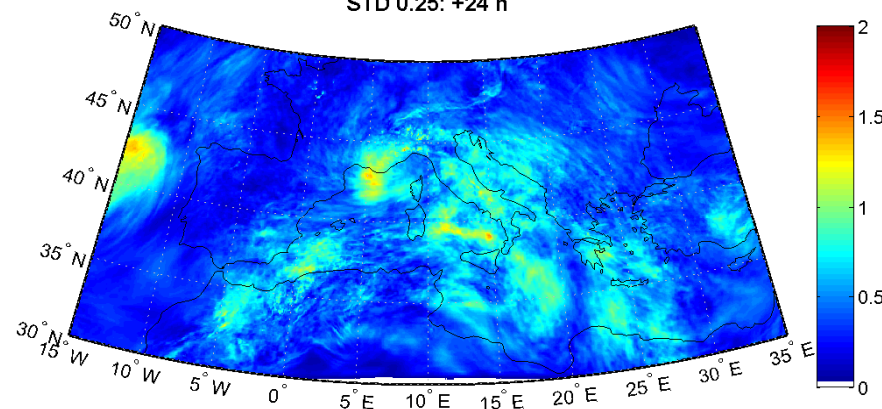
T+12h

STD 0.25: +12 h



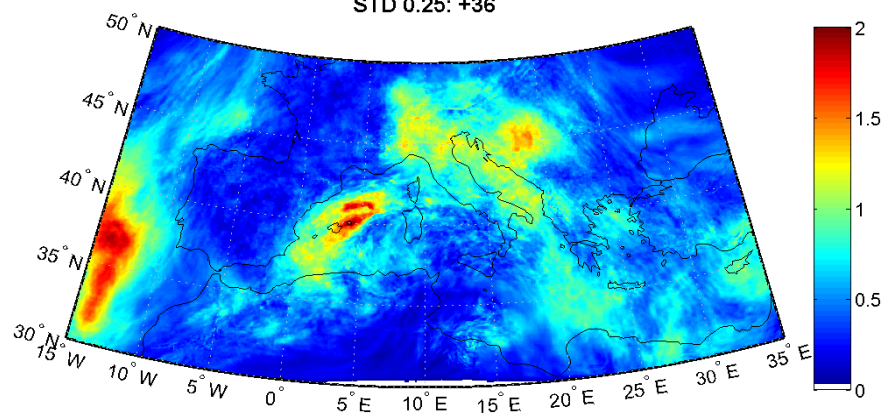
T+24h

STD 0.25: +24 h



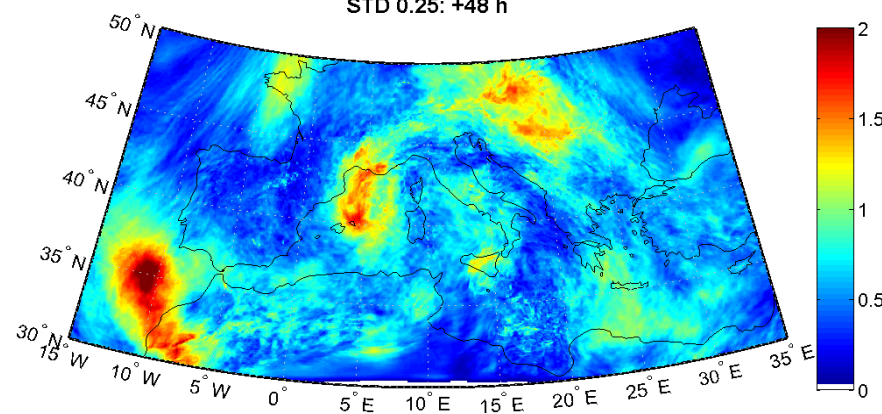
T+36h

STD 0.25: +36



T+48h

STD 0.25: +48 h





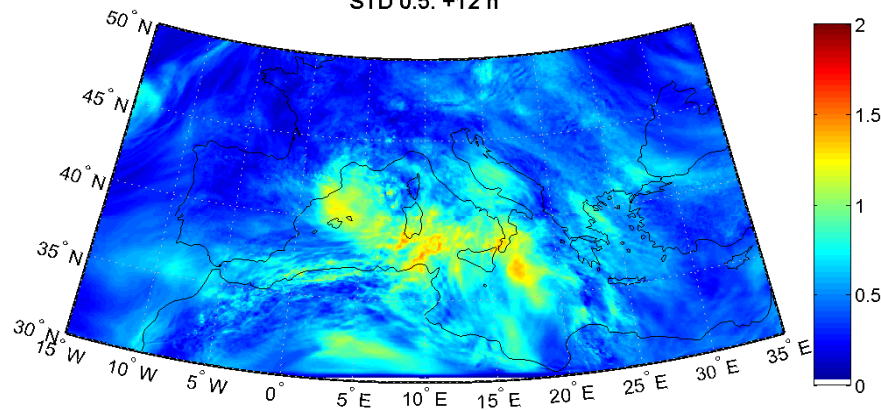
# Experiment: 05 June 2011 case

500 hPa Temperature Spread for 10 members

stdv=0.5 range=1.

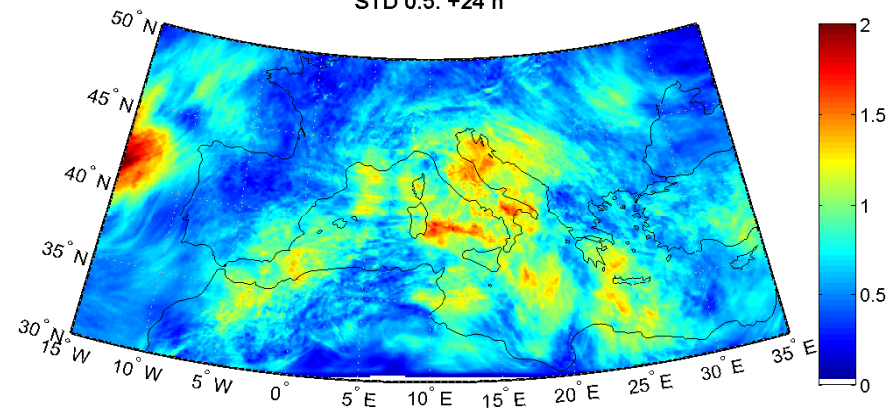
T+12h

STD 0.5: +12 h



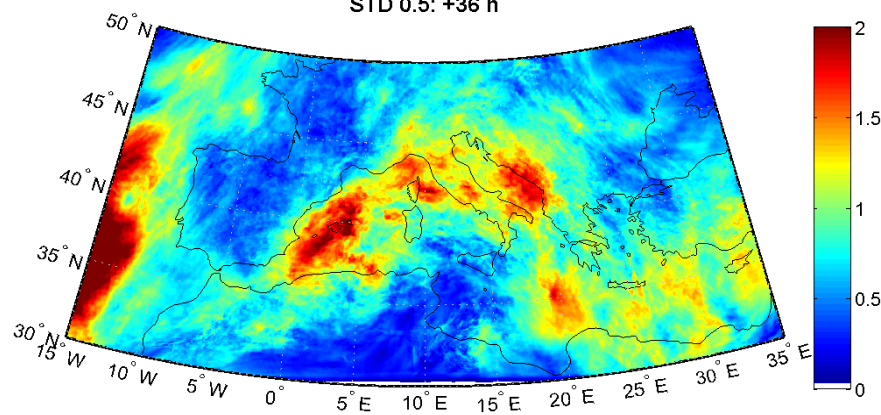
T+24h

STD 0.5: +24 h



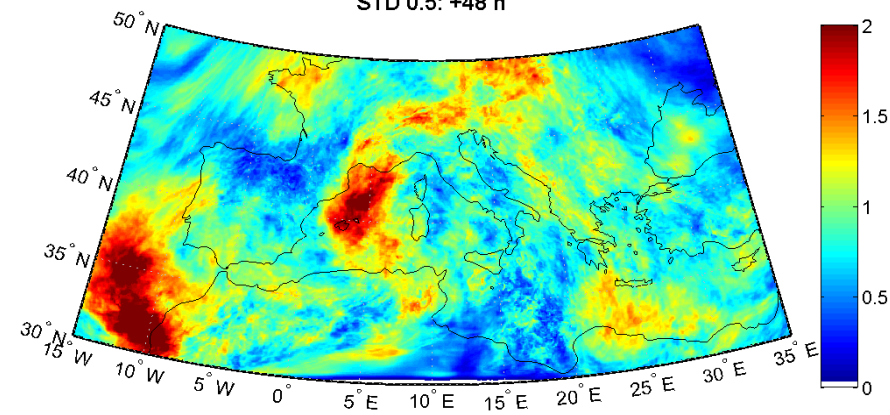
T+36h

STD 0.5: +36 h



T+48h

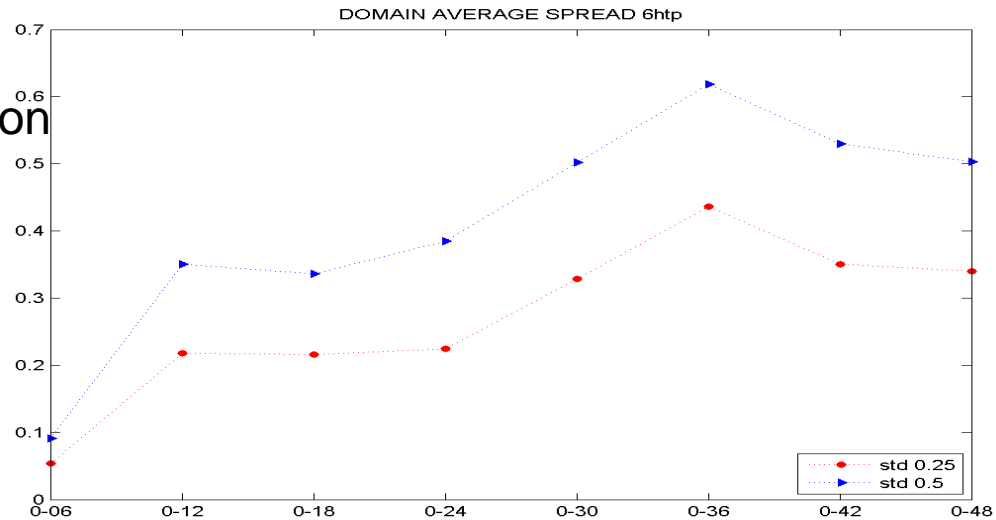
STD 0.5: +48 h



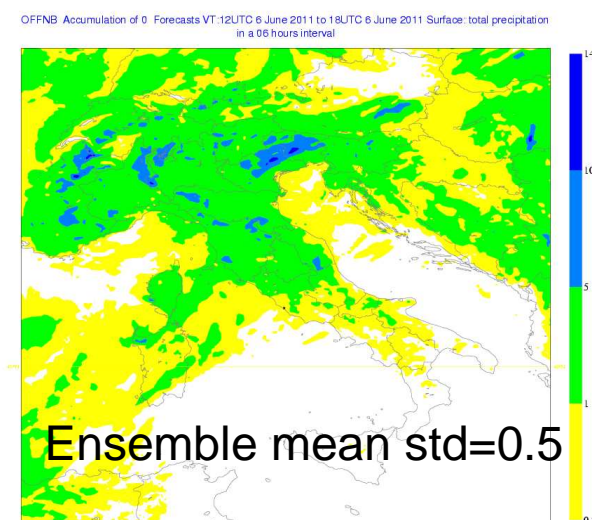
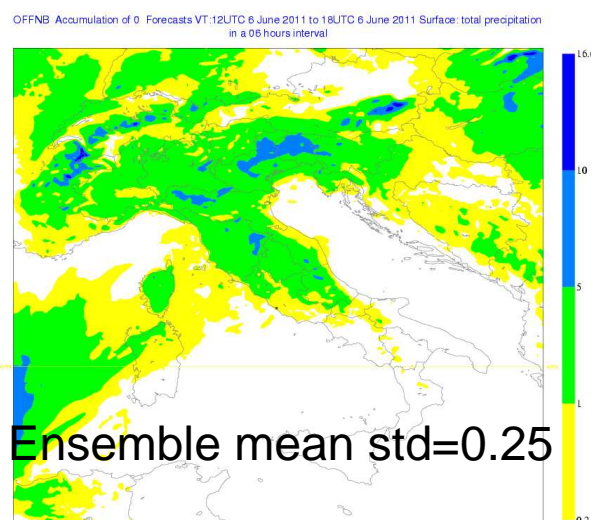
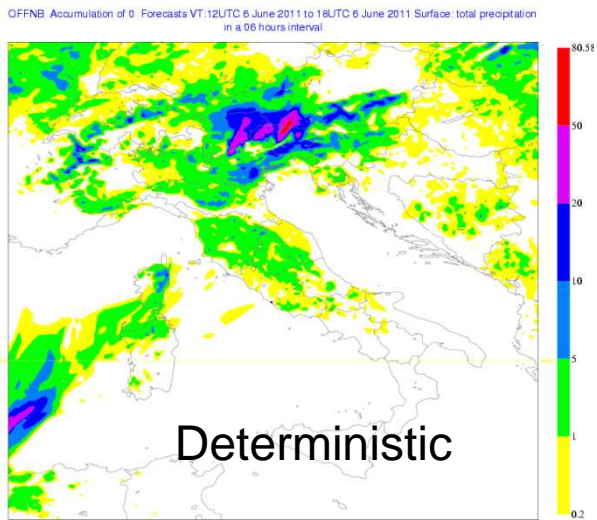


# Experiment: 05 June 2011 case

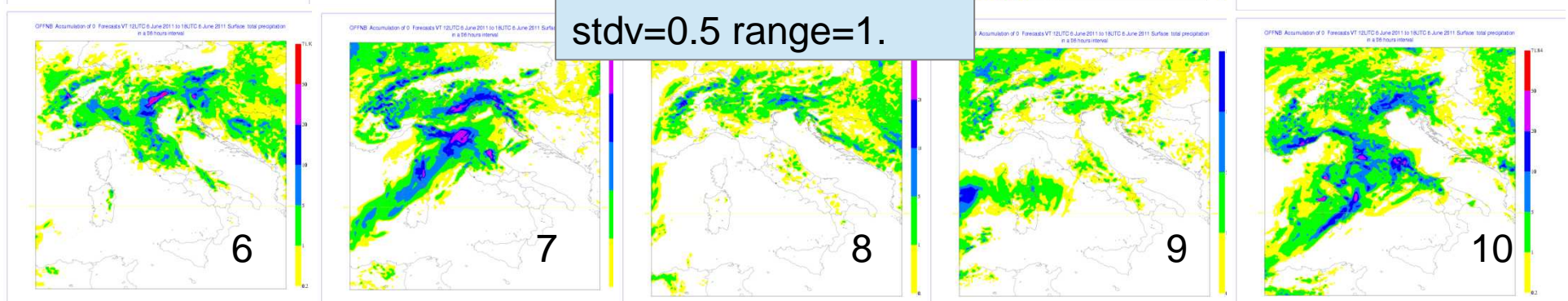
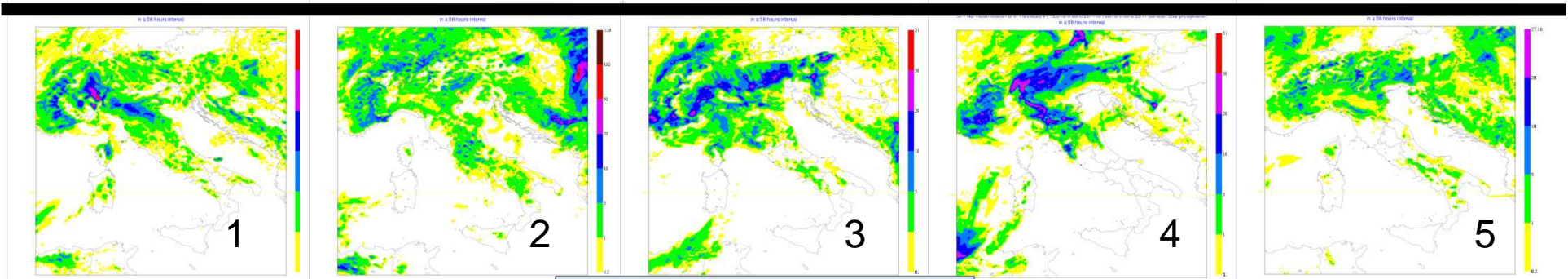
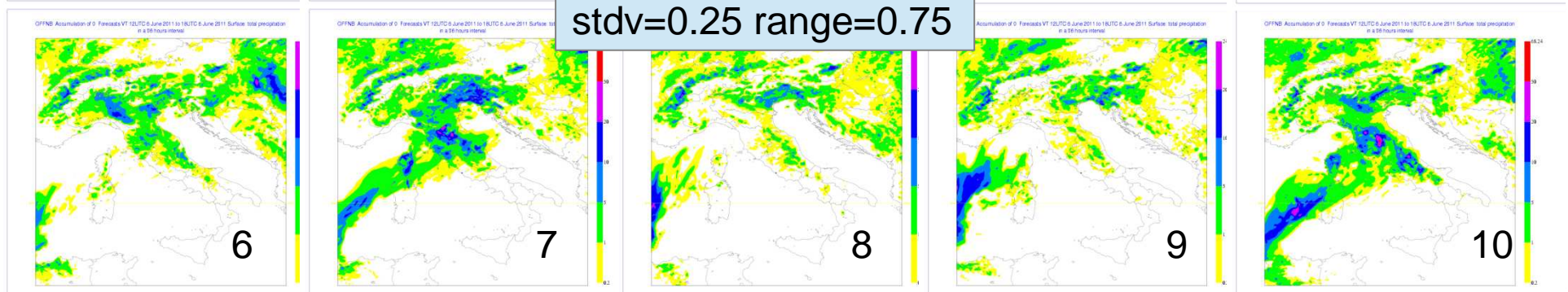
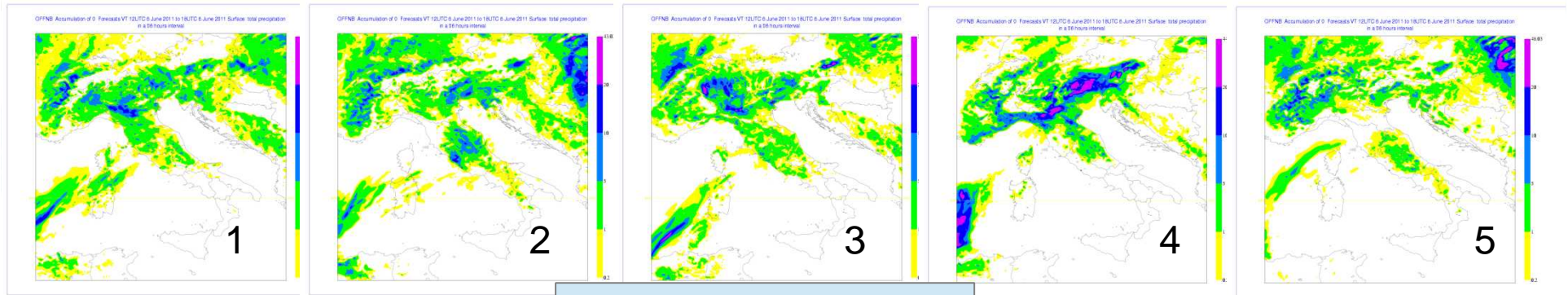
6h Accumulated Precipitation  
Domain Averaged Spread  
for 10 members



6h Accumulated Precipitation (T+36 - T+42h)









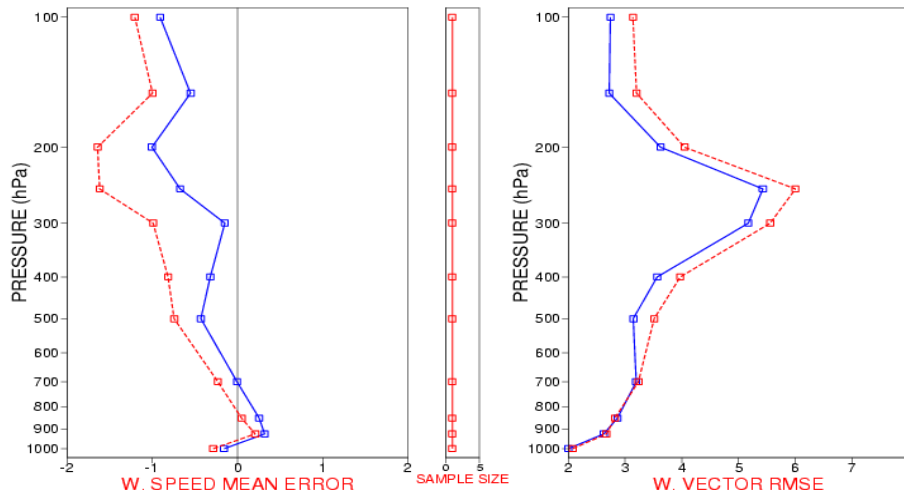
# Experiment: 05 June 2011 case

## Ensemble Mean Forecast against IFS Analysis

WIND (m/s) 00 UTC FC + 24 h

Verification 05/06/11

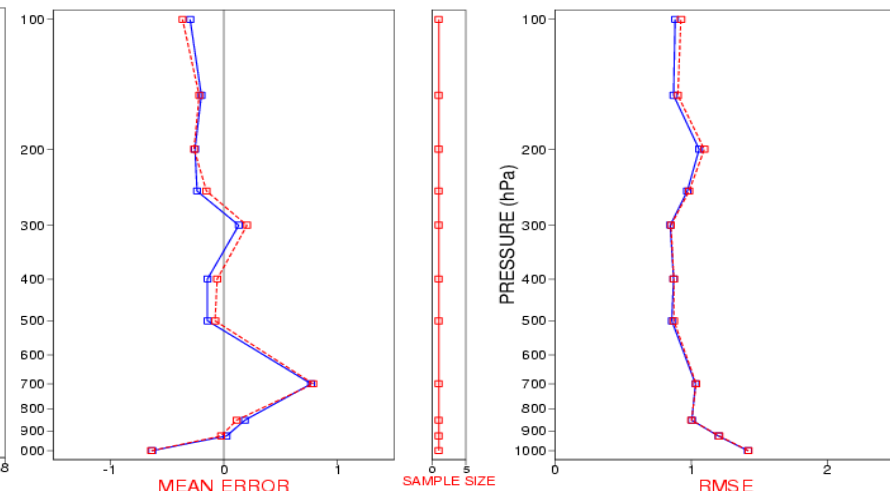
COSMO-ME\_0.25: Blue COSMO-ME\_0.5: Red



TEMPERATURE (°C) 00 UTC FC + 24 h

Verification 05/06/11

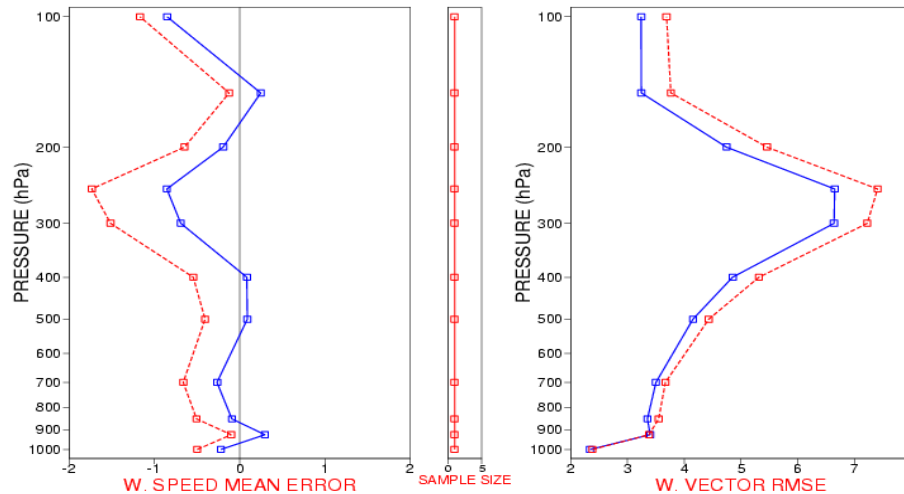
COSMO-ME\_0.25: Blue COSMO-ME\_0.5: Red



WIND (m/s) 00 UTC FC + 48 h

Verification 05/06/11

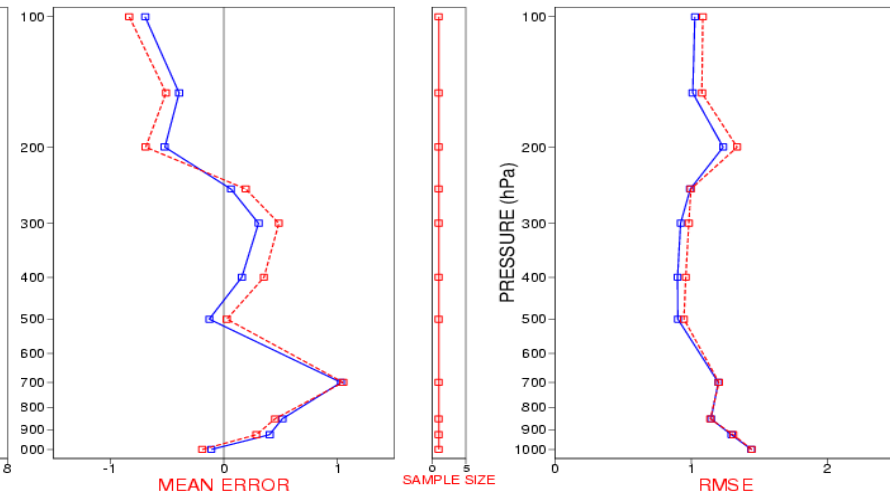
COSMO-ME\_0.25: Blue COSMO-ME\_0.5: Red



TEMPERATURE (°C) 00 UTC FC + 48 h

Verification 05/06/11

COSMO-ME\_0.25: Blue COSMO-ME\_0.5: Red





# Conclusions

- Some improvements were done on the modified Buizza stochastic physics implemented in COSMO model
- Two experiments using a gaussian distribution of random numbers with 0.25 and 0.5 standard deviation using 10 integrations of COSMO-ME for 05 July 2011 were performed
- The COSMO-ME ensemble spread increases as a function of forecast time
- From the comparison of the ensemble mean forecasts with IFS analysis and observations the run with  $\text{std}=0.25$  seems to verify better
- More experiments are needed to evaluate the “best tuning” of stochastic physics in COSMO model





Thanks for the attention!  
Any questions?

