

Status Overview on PP KENDA

Km-scale ENsemble-based Data Assimilation

Deutscher Wetterdienst
Wetter und Klima aus einer Hand



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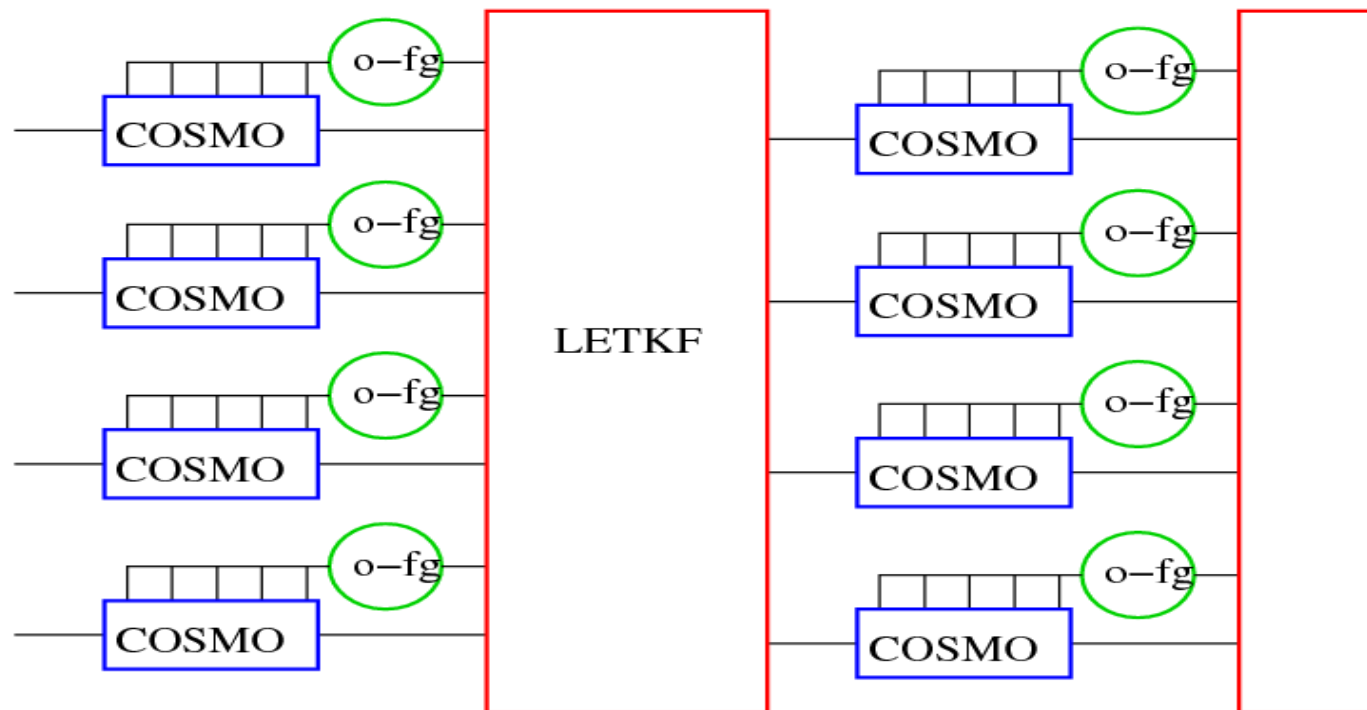
- Task 1: General issues in the convective scale (e.g. non-Gaussianity)
- Task 2: Technical implementation of an ensemble DA framework / LETKF
- Task 3: Tackling major scientific issues, tuning, comparison with nudging
- Task 4: Inclusion of additional observations in LETKF



Task 2: Technical implementation of an ensemble DA framework / LETKF

analysis step (LETKF) outside COSMO code

- ensemble of independent COSMO runs up to next analysis time
- separate analysis step code, LETKF included in 3DVAR code of DWD



collect obs from
 $]t_{ana}-3h, t_{ana}]$

old setup: obs from
 $]t_{ana}-1.5h, t_{ana}+1.5h]$
thinned

Task 2: Technical implementation of an ensemble DA framework / LETKF



- modifications in COSMO fully implemented, but not yet in official code (also in order to have a sub-hourly update frequency)
- first version LETKF (KENDA) implemented in NUMEX, to be tested
→ still use scripts to do a few stand-alone cycles with LETKF
- deterministic analysis implemented → talk of Hendrik
- LBC at analysis time : use (perturbed) analysis ensemble member itself instead of (temporally interpolated) LBC fields from steering ensemble system
- next step: LBC from GME LETKF ensemble
interpolate GME ensemble perturbations and add to deterministic COSMO-DE LBC



Analysis for a deterministic forecast run :
use Kalman Gain \mathbf{K} of analysis mean

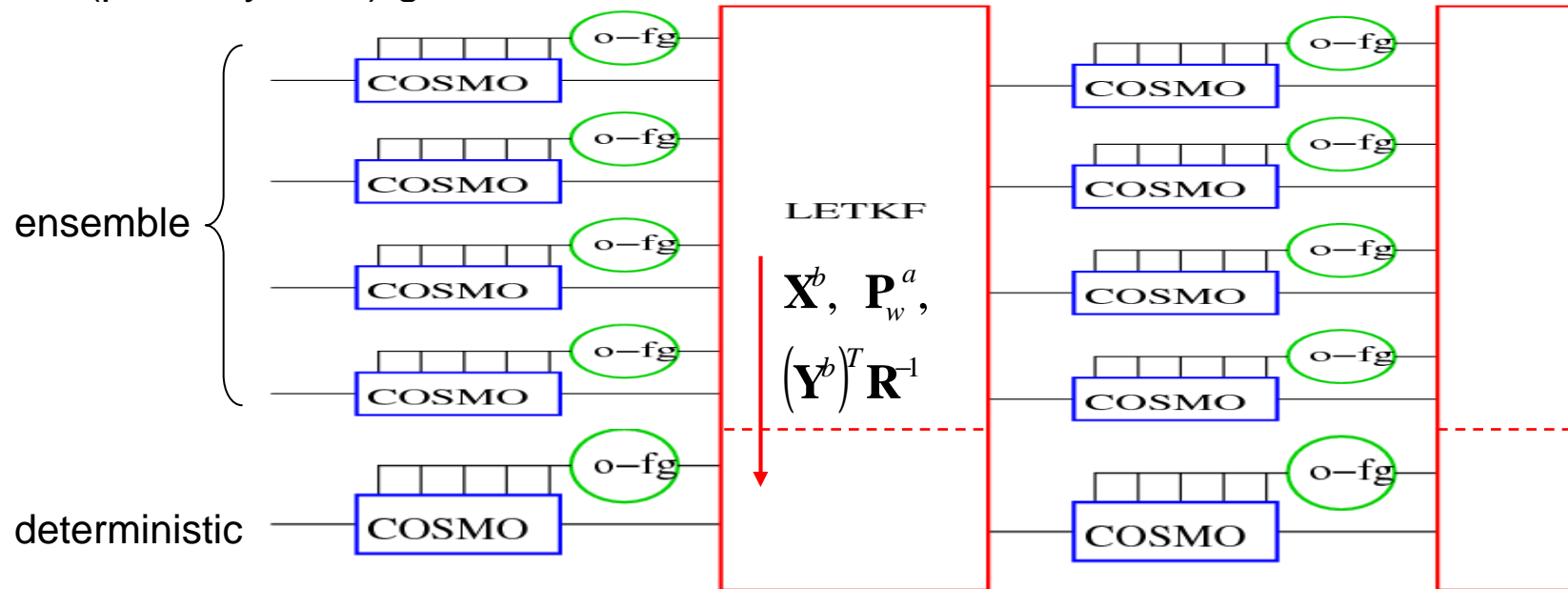


$$\mathbf{x}^A = \mathbf{x}^B + \mathbf{K} [\mathbf{y}^O - H(\mathbf{x}^B)] \quad , \quad \mathbf{K} = L \mathbf{X}^b (\mathbf{P}_w^a)^{-1} (\mathbf{Y}^b)^T \mathbf{R}^{-1}$$

$$\mathbf{P}_w^a = (k-1) \mathbf{I} + (\mathbf{Y}^b)^T \mathbf{R}^{-1} \mathbf{Y}^b$$

deterministic analysis
recently implemented

L : interpolation of analysis increments from grid of LETKF ensemble
to (possibly finer) grid of deterministic run



- deterministic run must use same set of observations as the ensemble system !
- Kalman gain / analysis increments not optimal,
if deterministic background \mathbf{x}^B (strongly) deviates from ensemble mean background



Task 2: Technical implementation: verification for KENDA



'stat'-utility:

compute model (forecast) – obs for complete NEFF (NetCDF feedback files) :

- adapt verification mode of 3DVar/LETKF package
- need to implement COSMO observation operators with QC ('library 2') in 3DVar/LETKF package
(→ hybrid 3DVar–EnKF approaches in principle applicable to COSMO)

- build COSMO modules with clean interfaces & introduce them in LETKF
 - no further progress
 - still plan to first implement upper-air obs (needed as input for VERSUS)



Task 2: Technical implementation: verification for KENDA

NEFFprove (Amalia Iriza, NMA) :

tool to compute and plot verification scores based on NetCDF feedback files (NEFF)

- uses NEFF files from 3DVAR and COSMO environments
- computes statistical scores for different runs ('experiments'),
focus: use exactly the same observation set in each experiment !
 - sort observations in a unique order (lon / level / time / lat / variable / code ..)
 - select obs according to namelist values (area, obs quality, status, ...)
- compute scores (for each experiment, variable, forecast time, vertical level, ... ,
upper-air obs, screen-level obs, RR, cloud, etc.)
 - deterministic continuous: BIAS, RMSE, MSE
 - dichotomous: accuracy, Heidke Skill Score, Hanssen and Kuiper discriminant
 - ensemble scores (to be done): reliability, ROC, Brier Skill Score,
(continuous) Ranked Probability Score
- write scores to ASCII files (1 file for each variable, with all scores for each exp.)
- plot the scores, using GNUplot (work is in progress)

Task 3: scientific issues & refinement of LETKF



- lack of spread: (partly ?) due to model error and limited ensemble size which is not accounted for so far
to account for it: covariance inflation, what is needed ?
 - multiplicative $X_b \rightarrow \rho \cdot X_b$ (tuning, or adaptive ($y - H(x) \sim R + H^T P_b H$))
 - additive ((e.g. statistical 3DVAR-B), stochastic physics (Torrissi))
- localisation (multi-scale data assimilation,
2 successive LETKF steps with different obs / localisation ?)

→ talk by Hendrik Reich

- model bias
- update frequency : up to now only 3-hourly
- noise control : new: use IC as initial LBC
- non-linear aspects, convection initiation (latent heat nudging ?)
- technical aspects: efficiency, system robustness
(Nov. 2012: regular 'pre-operational' LETKF suite)



- **stochastic physics (Palmer et al., 2009)** : by Lucio Torrisi, available Oct. 2011
→ talk by Lucio Torrisi
- **objective estimation and modelling of model (tendency) errors**
→ talk by Michael Tsyruльников

- **investigate LETKF in Observing System Simulation Experiments (OSSE)**
 - apply LETKF to idealized convective weather systems, tune LETKF settings (localization, covariance inflation)
 - quantify non-Gaussianity and spin-up time in the evolving LETKF ensemble during the assimilation of a convective storm
 - develop ways to improve Gaussianity of errors and reduce spin-up time of the LETKF → test “running in place” (Kalnay and Yang, 2010) and “outer loop” (Yang and Kalnay, 2010) algorithms

MeteoSwiss: plan for 2-year project not accepted

- however master thesis (5 months): Manuel Bischof, MeteoSwiss
 - perturbation for ensemble generation:
 - which quantities, describing the storm environment, suitable to perturb ?
 - meaningful perturbation amplitudes ?

→ talk by Daniel Leuenberger

- **radar : (3-D) radial velocity and (3-D) reflectivity**

Uli Blahak (DWD), Yuefei Zeng, Dorit Epperlein (KIT Karlsruhe)

- Implement full, sophisticated observation operators (by end 2011)
- then develop different approximations and test their validity in order to obtain sufficiently accurate and efficient operators
 - by looking at the simulated obs (by March 2012)
 - doing assimilation experiments : OSSE setup

Particular issues for use in LETKF:

obs error variances and correlations,
superobbing, thinning,
localisation

→ talk by Uli Blahak

- **ground-based GNSS slant path delay** (early 2012, N.N.)
 - implement non-local obs operator in parallel model environment
 - test and possibly compare with using GNSS data in the form of IWV or of tomographic refractivity profiles

Particular issue: localisation for (vertically and horizontally) non-local obs

- **cloud information based on satellite and conventional data**

Annika Schomburg (DWD / Eumetsat)

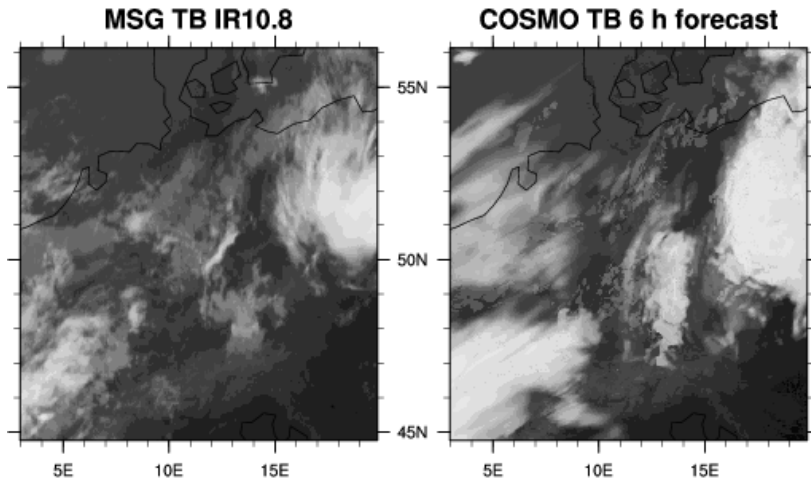
- derive incomplete analysis of cloud top + cloud base, using conventional obs (synop, radiosonde, ceilometer) and NWC-SAF cloud products from SEVIRI
- use obs increments of cloud or **cloud top / base height** or derived humidity
- use SEVIRI **brightness temperature** directly in LETKF in cloudy (+ cloud-free) conditions, in view of improving the horizontal distribution of cloud and the height of its top
- compare approaches

Particular issues: **non-linear observation operators,**
non-Gaussian distribution of observation increments

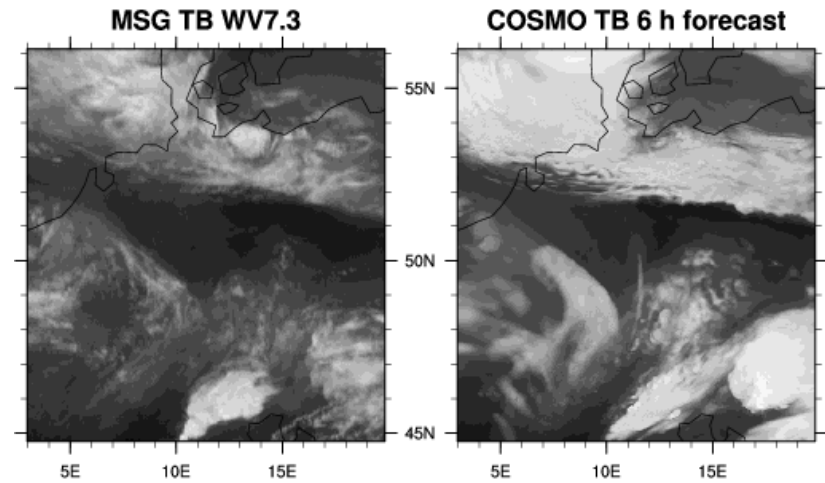
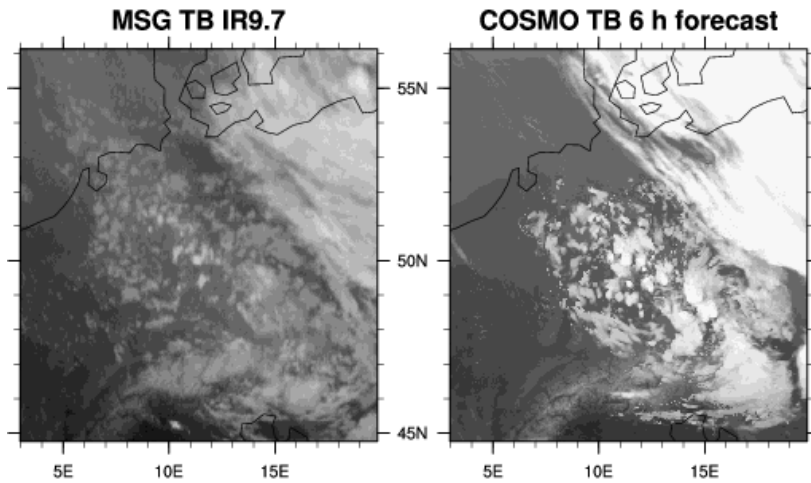
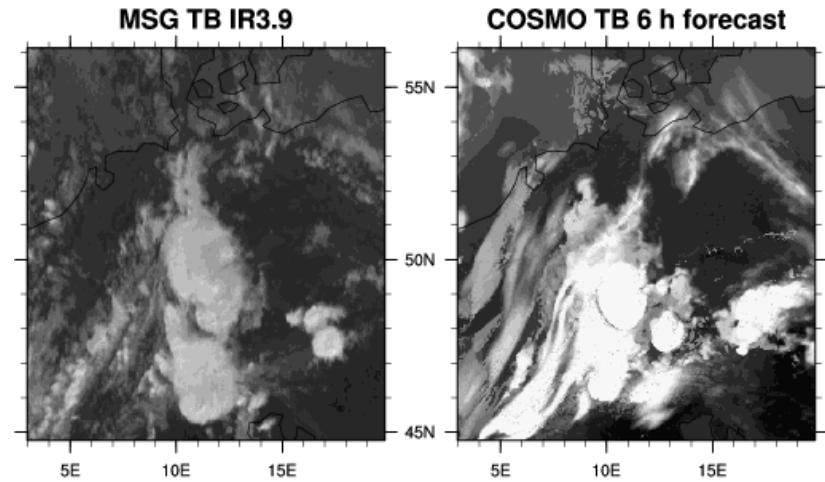
Task 4.4: use of cloud info

- make observations available
 - aggregation / interpolation to COSMO-DE / COSMO-EU grids
 - regular archiving:
 - NWC-SAF cloud products available since May 2011:
(CT: cloud type, CTT: cloud top temperature, CTH: cloud top height)
 - BT available since June 2011: all 12 SEVIRI channels
- assumption in LETKF: no bias
 - look at systematic differences in BT / cloud products

Task 4.4: use of cloud info



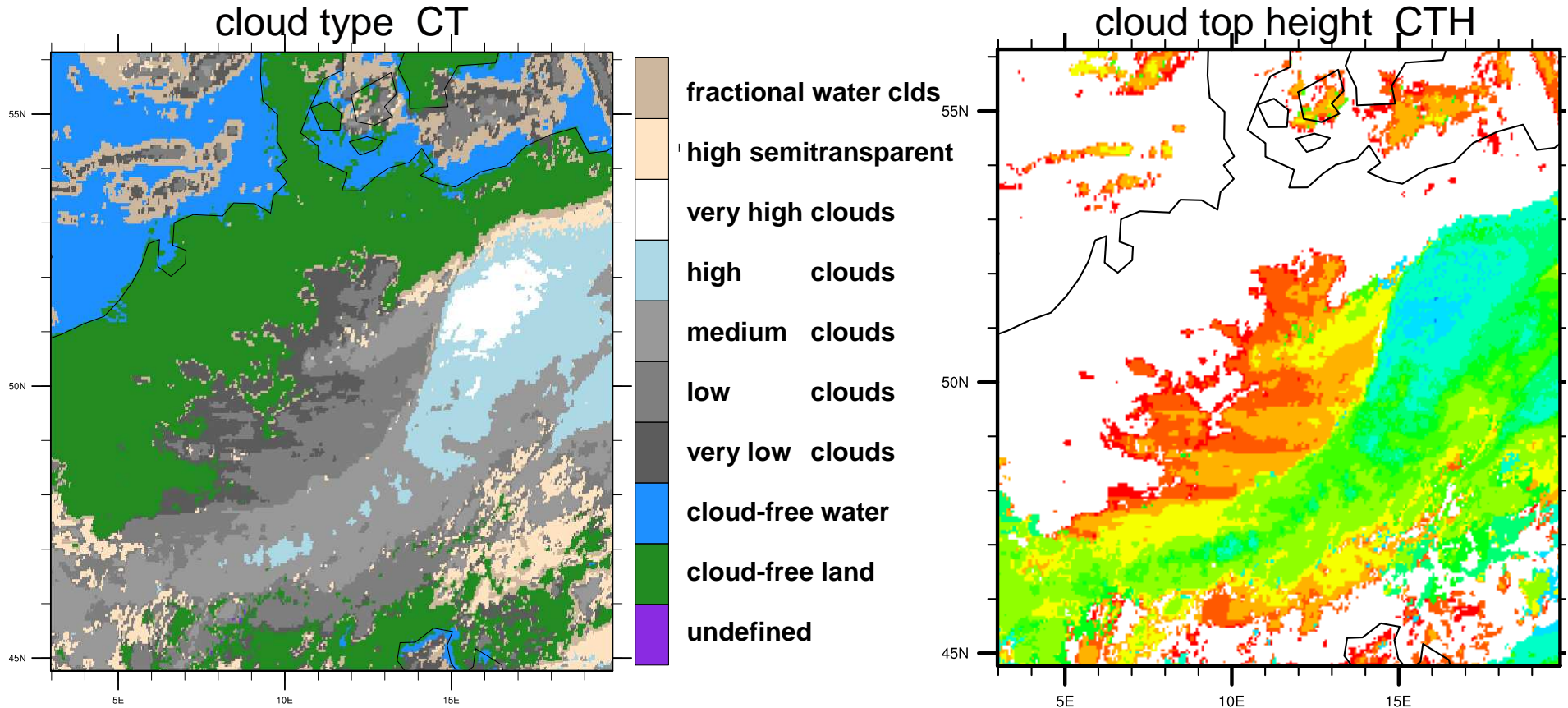
BT



high clouds: cold bias of model-BT, (partly ?) due to known bias in RTTOV-9

Task 4.4: use of cloud info

NWC-SAF SEVIRI cloud products: example

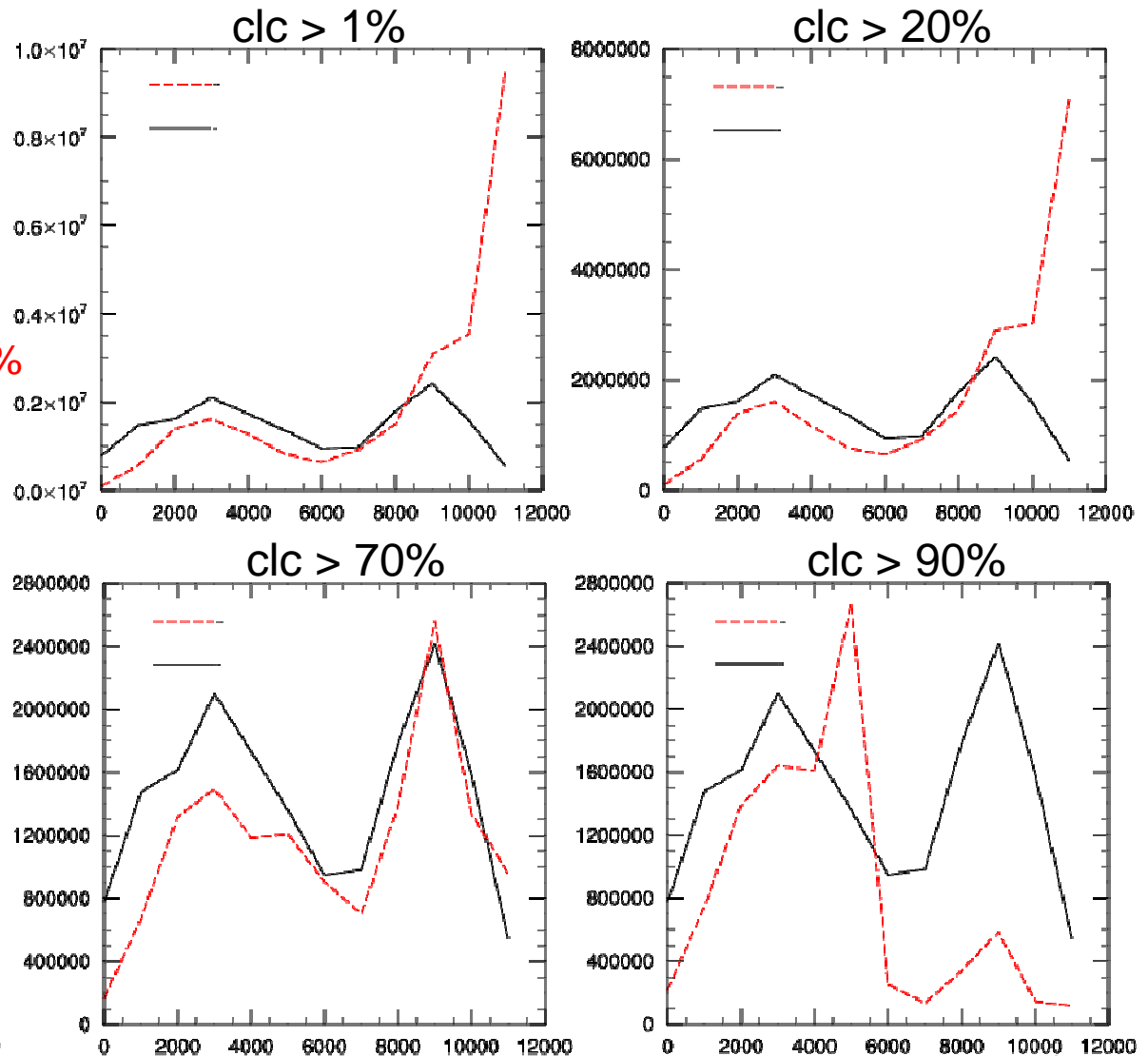


COSMO: cloud water $q_c > 0$, or cloud ice $q_i > 5 \cdot 10^{-5} \text{ kg/kg}$ \rightarrow $clc = 100 \%$
 subgrid-scale clouds \rightarrow $clc = f(\text{RH}; \text{shallow convection}; q_i, q_{i,\text{sgs}}) < 100 \%$

Task 4.4: use of cloud info

cloud top height CTH
distribution
May – July 2011

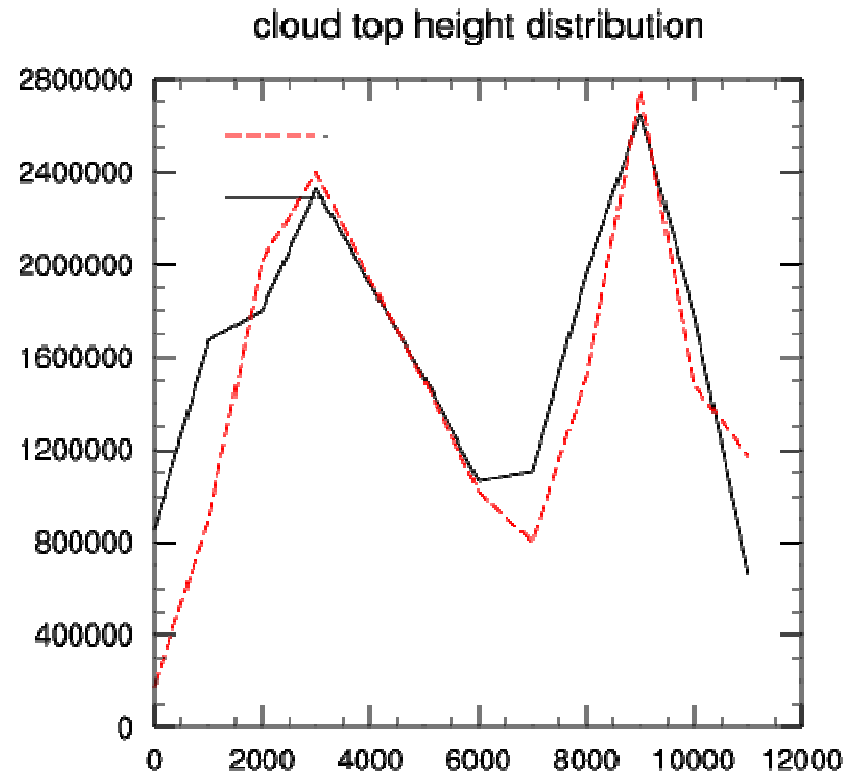
— NWC-SAF CTH 'obs'
 - - - COSMO CTH for $clc > x \%$



Task 4.4: use of cloud info

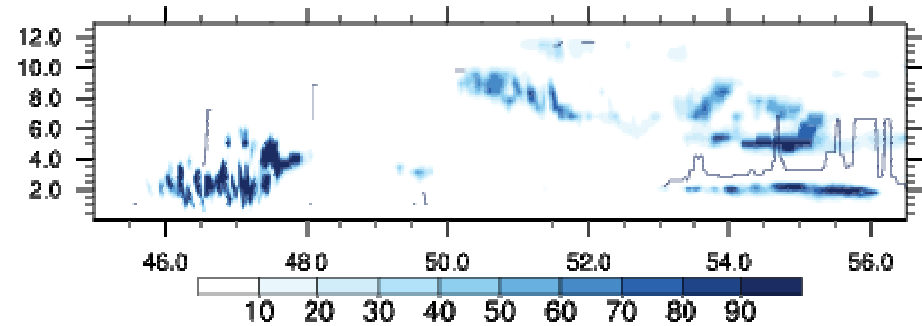
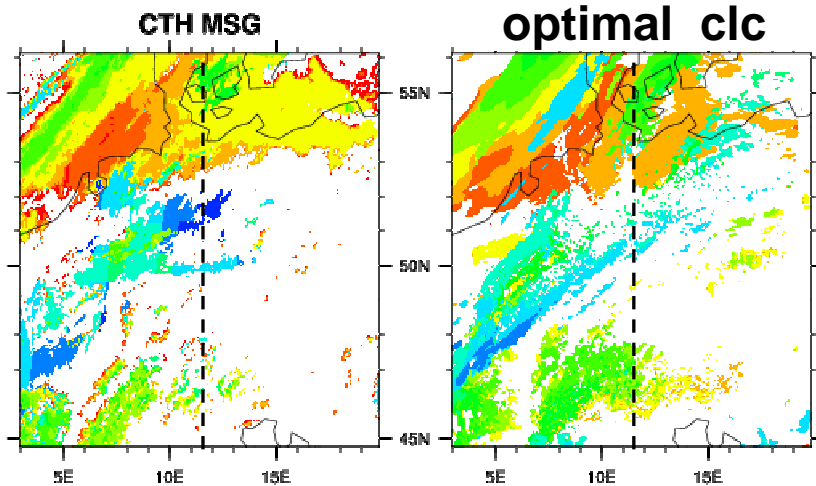
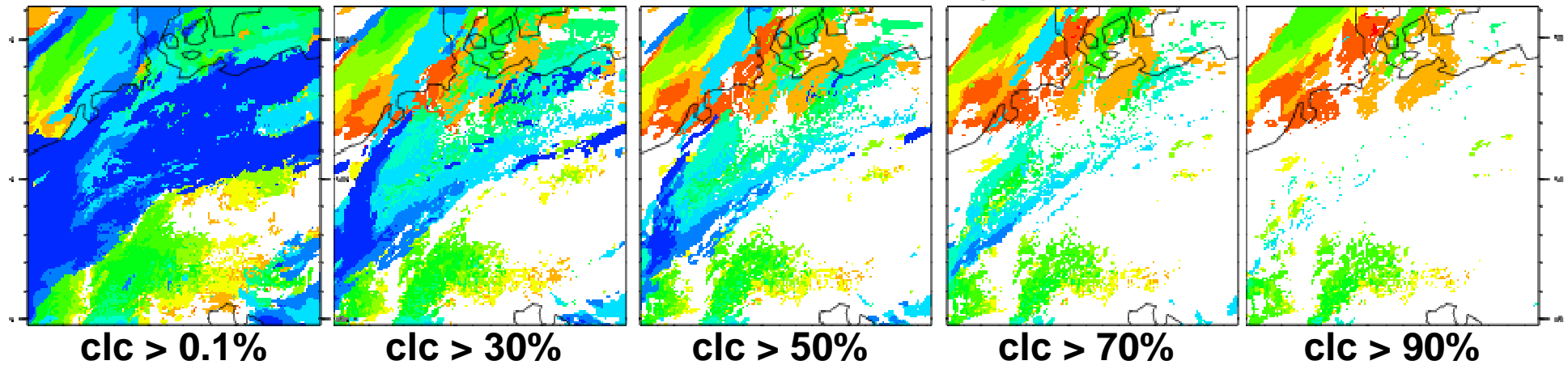
cloud top height CTH distribution May – July 2011

- NWC-SAF CTH 'obs'
- COSMO CTH with optimal threshold: for $cl_c > 70\%$ for levels above 5000 m ,
for $cl_c > 40\%$ for levels below 5000 m

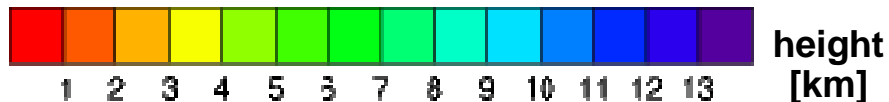


Task 4.4: use of cloud info

CTH COSMO 6-h forecast, for 18 May 2011, 18 UTC



cloud cover [%] , shading: model
isoline: CTH MSG



Technical implementation of verification for KENDA

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Proposition of WG3b (Jean-Marie Bettems) for other WGs :

WG1 : Assimilation of land surface data

To overcome some climatological assumptions, evaluate usage of LandSAF products

MSG/SEVIRI based products

- **Wild Fires** (Fire Radiative Power, Fire Risk Map, Fire Detection and Monitoring)
- **Vegetation Parameters** (Fraction of Vegetation Cover, Leaf Area Index, Fraction of Absorbed Photosynthetic Active Radiation)
- **Snow Cover** (daily, 15 mins)
- **Other** : Bi-Directional Reflectance Factor, Land Surface Emissivity
- **Surface Albedo**
- **Land Surface Temperature** (15 mins)
- **Down-welling Surface Fluxes** (short-wave, long-wave)
- **Evapotranspiration** (30 mins, daily) → additional control variable in var SMA ?

MetOp/AVHRR based products

- **Land Surface Temperature**
- **Down-welling Surface Fluxes** (long-wave)



Technical implementation of verification for KENDA

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- Post-processing routine to prepare statistics for ensemble runs
- Verification based on netcdf feedback files
- Uses files from the 3DVAR and COSMO environments
- Computes statistical scores for different runs, based on exactly the same observations for each of the used experiments



- namelist reading: specifications on models + observation types used, selection criteria for the obs (lat/lon, quality, state, ...) , etc.
- select obs for each 3-hrly time box :
 - sort observations to get a unique order which allows the identification of the observations in all experiments (reduces computing time for the next step):
 - determine maximum number of observations
 - sort observations according to three criteria:
 - longitude, vertical level
 - time, latitude, observed variable
 - code type, sub-centre, status
 - select the observations (dep. on namelist):
 - eliminate the observations that are not used in all the experiments, or do not correspond to at least one of the criteria specified in the namelist
 - select according to verification times



- update contingency tables:
 - compute intermediate scores for upper air temperature, relative humidity, surface pressure, wind components, vertical velocity;
 - update contingency tables for dichotomous scores (for precipitation amount and total cloud amount).
- compute scores (for each experiment, each variable, forecast time, vertical level, ...)
 - deterministic continuous: BIAS, RMSE, MSE
 - dichotomous: accuracy, Heidke Skill Score, Hanssen and Kuiper discriminant
 - ensemble score (to be done): reliability, ROC, Brier Skill Score, (continuous) Ranked Probability Score
- write scores to ASCII files (1 file for each variable, containing all computed scores for each experiment).
- plot the scores, using GNUplot (work is in progress)

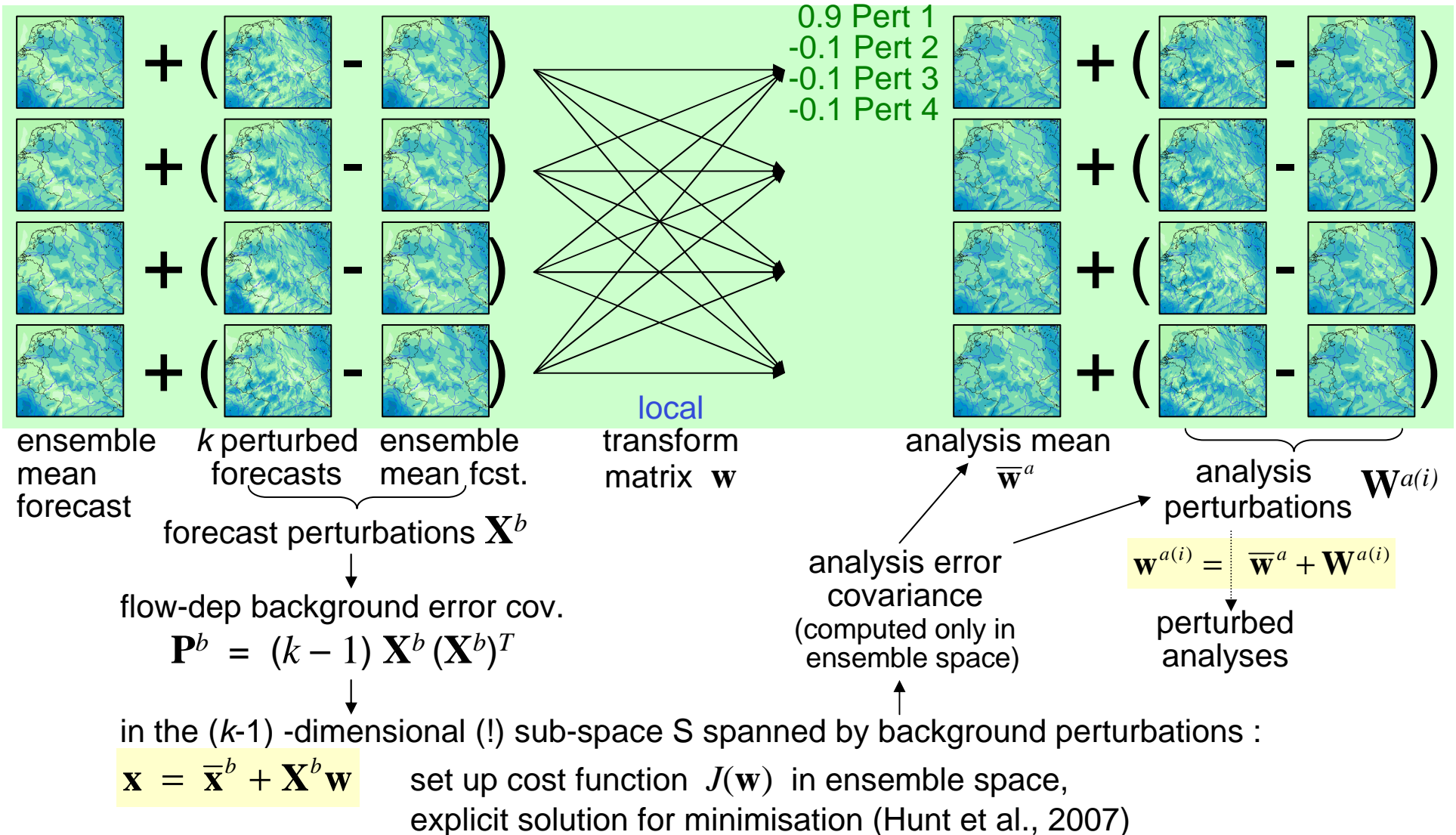


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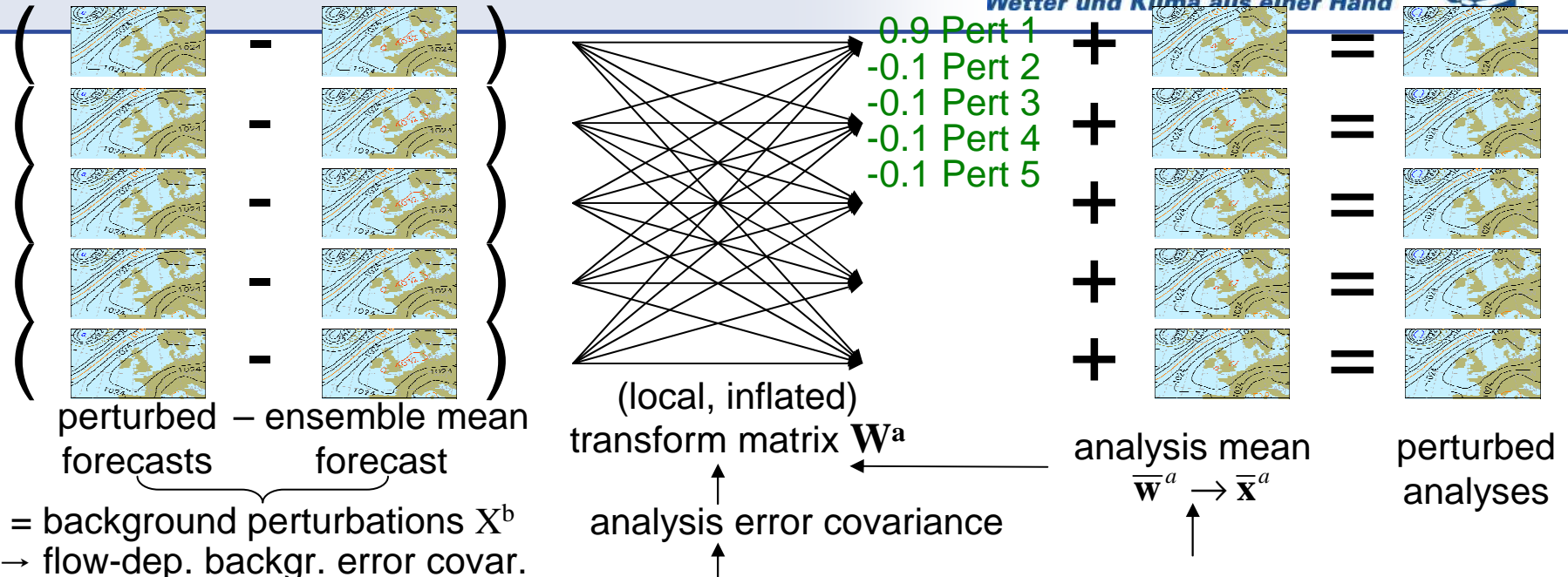


Local Ensemble Transform Kalman Filter LETKF (Hunt et al., 2007)



Local Ensemble Transform Kalman Filter (LETKF)

(chart based on a slide of Neil Bowler, UK MetOffice)



in the (5-dimensional) sub-space \mathbf{S} spanned by background perturbations :
(linearise obs operator around the mean of ensemble of simulations of an obs) $\bar{\mathbf{y}}^b \equiv \mathbf{y}^{b(i)} \equiv H(\mathbf{x}^{b(i)})$
explicit solution for minimisation of cost function (Hunt et al., 2007)

→ analysis (mean) and analysis error covariance : columns of $\mathbf{Y}^b : \mathbf{y}^{b(i)} - \bar{\mathbf{y}}^b$

$$\bar{\mathbf{w}}^a = \mathbf{P}_w^a (\mathbf{Y}^b)^T \mathbf{R}^{-1} [\mathbf{y}^o - \bar{\mathbf{y}}^b], \quad \mathbf{P}_w^a = \left[(k-1) \mathbf{I} + (\mathbf{Y}^b)^T \mathbf{R}^{-1} \mathbf{Y}^b \right]^{-1}$$

and in model space : $\bar{\mathbf{x}}^a = \bar{\mathbf{x}}^b + \mathbf{X}^b \bar{\mathbf{w}}^a, \quad \mathbf{P}^a = \mathbf{X}^b \mathbf{P}_w^a (\mathbf{X}^b)^T$

→ analysis ensemble member : $\mathbf{x}^{a(i)} = \bar{\mathbf{x}}^a + \mathbf{X}^{a(i)} = \bar{\mathbf{x}}^b + \mathbf{X}^b (\bar{\mathbf{w}}^a + \mathbf{W}^{a(i)}) \quad \mathbf{W}^a = \left[(k-1) \mathbf{P}_w^a \right]^{1/2}$



LETKF (COSMO) : method to deal with non-Gaussianity



- „No cost smoother“ (Kalnay and Yang, QJRS, submitted)
 - Apply weights from time $t+1$ at time t
- „Running in place“
 - Use observations more than once by iterating several times over same assimilation window using the no cost smoother until „convergence“
- „Outer loop“ in LETKF (Yang and Kalnay, MWR, submitted)
 - Bring analysis mean closer to observations by iterative integration of ensemble mean from time $n-1$ to n (re-centering)
 - Advance to next assimilation time by integration of whole ensemble to time $n+1$
- Have proven to improve LETKF in presence of nonlinearity / non-Gaussianity in Lorenz model

